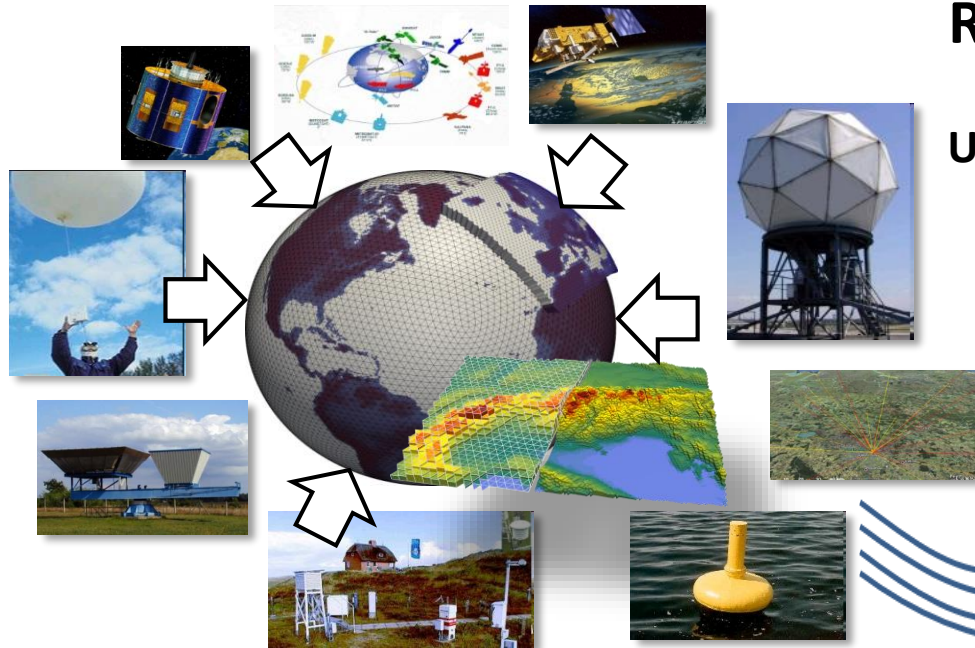


# Ensemble Data Assimilation and Particle Filters for NWP

With the help of many people, in particular:

Anne Walter,  
Andreas Rhodin  
Harald Anlauf,  
Christina Köpken,  
Robin Faulwetter,  
Olaf Stiller,  
Alexander Cress,  
Martin Lange,  
Stefanie Hollborn,  
E. Bauernschubert,  
Christoph Schraff,  
Hendrik Reich,  
Klaus Stephan  
Ulrich Blahak



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University of Reading

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1. Why and Where **Distributions, Risk** and **Uncertainty?**
2. Discussion of **Ensemble (+Particle) Methods**
3. Framework Global+LAM+LES Model: **ICON** and **ICON-EPS** and the **LEKTF+EnVAR/KENDA** System
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# Why Distributions , Risk, Uncertainty?

Extreme Weather Events  
triggered by climate change  
threaten security and  
economy.

Lightning, Cyclones, Heavy  
precipitation events impact  
the life of individuals and the  
society.

**Protect Lives!**

**Protect Property!**

**Protect Society and Business**

**We need Uncertainty  
Estimates for Risk Prediction!**

**National Task: Warn  
and Protect**



Phenomena	Level
Small (5-10)	Moderate
Medium (10-25mm/h)	Strong
Storm Force Gusts, Heavy Rainfall	Strong
Storm Force Gusts, Heavy Rainfall, Hail	Strong

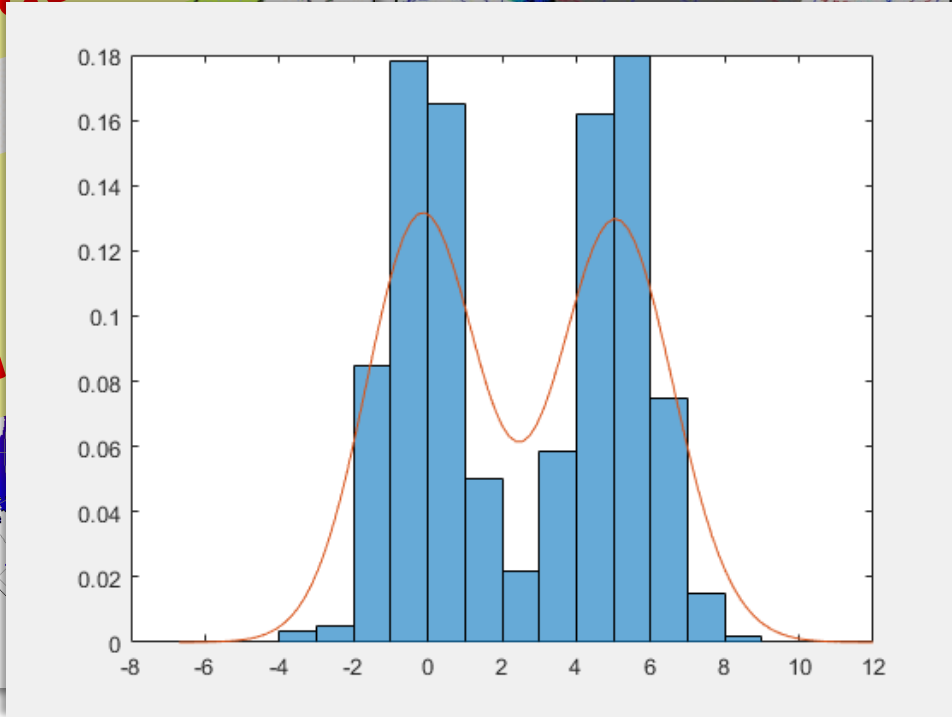
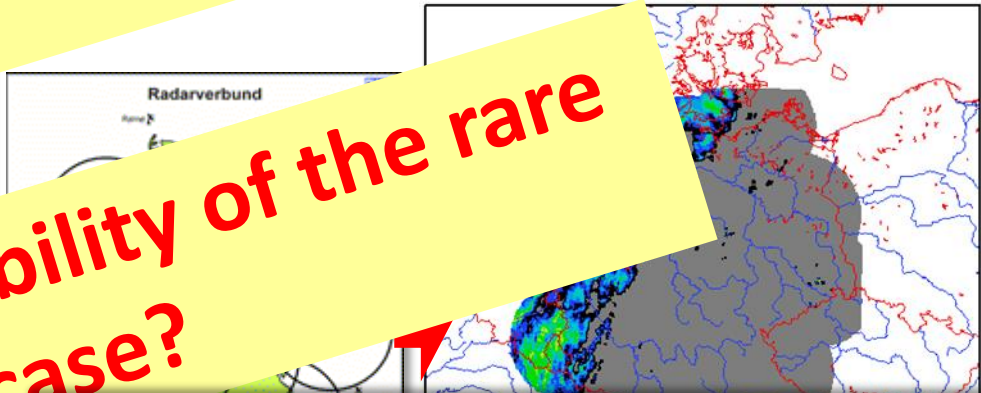
**NOW, 5min, 30min,  
... 1h, 2h, 6h, 24h, 72h, ...**

# Why Distributions, Risk, Uncertainty?

More than what would usually happen!

What is the probability of the rare event, the worst case?

We need Uncertainty Estimates for Special



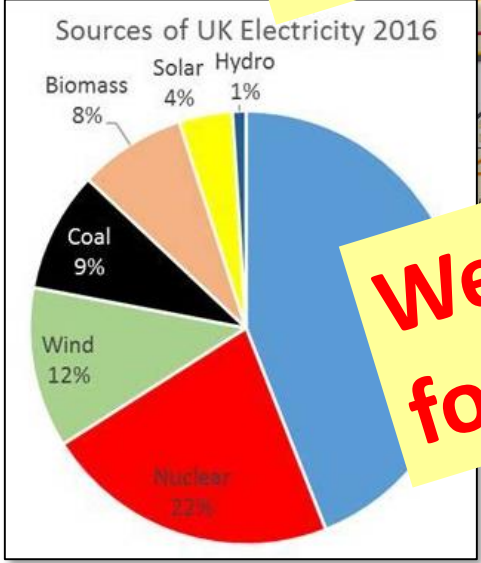
# Why Distributions , Risk, Uncertainty?

**6h Market: Millions**

**Day-Ahead Market: Billions!**

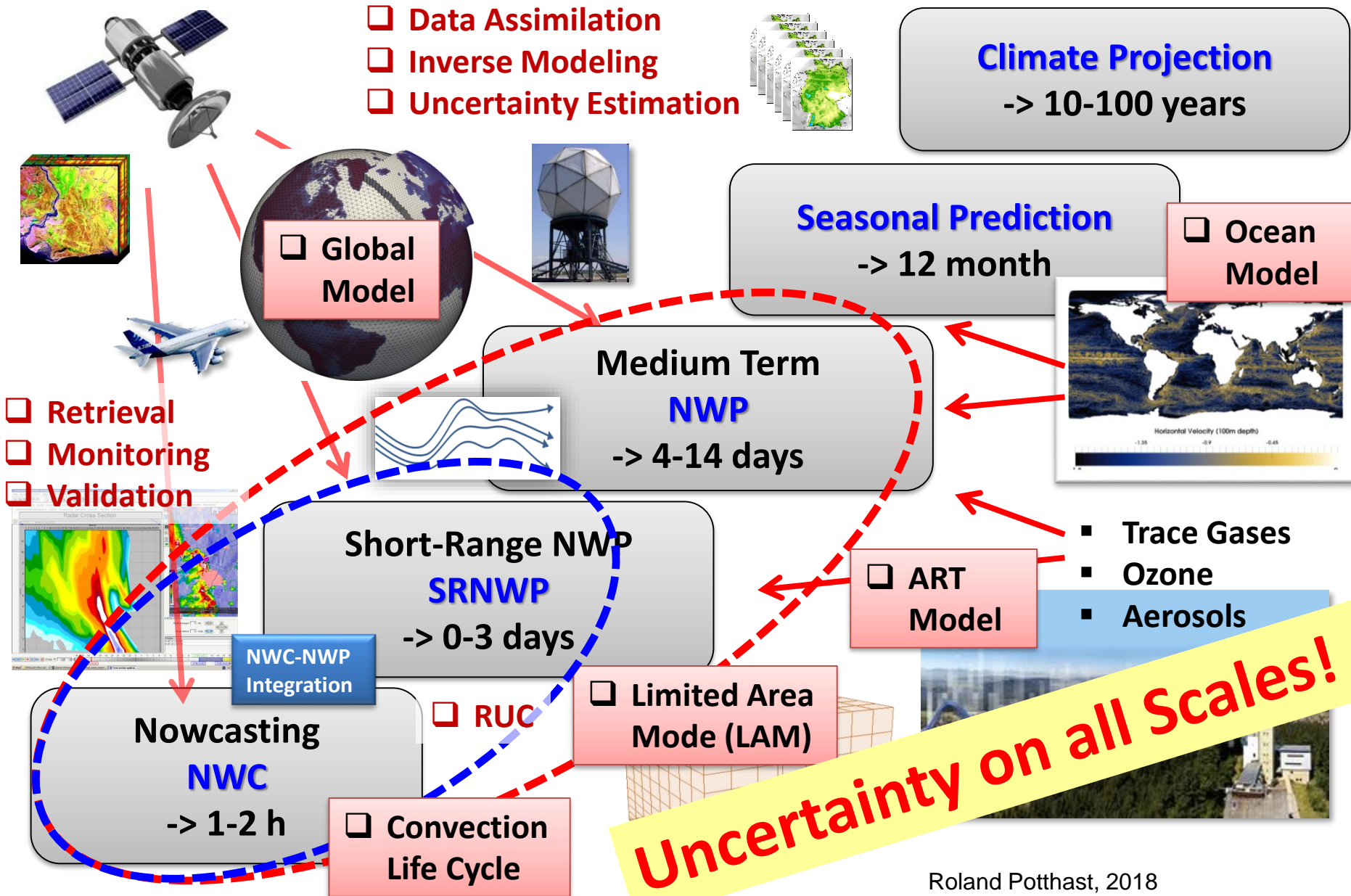
**We need Uncertainty Estimates for Energy Network Security!**

**Renawable Energy Forecasting**



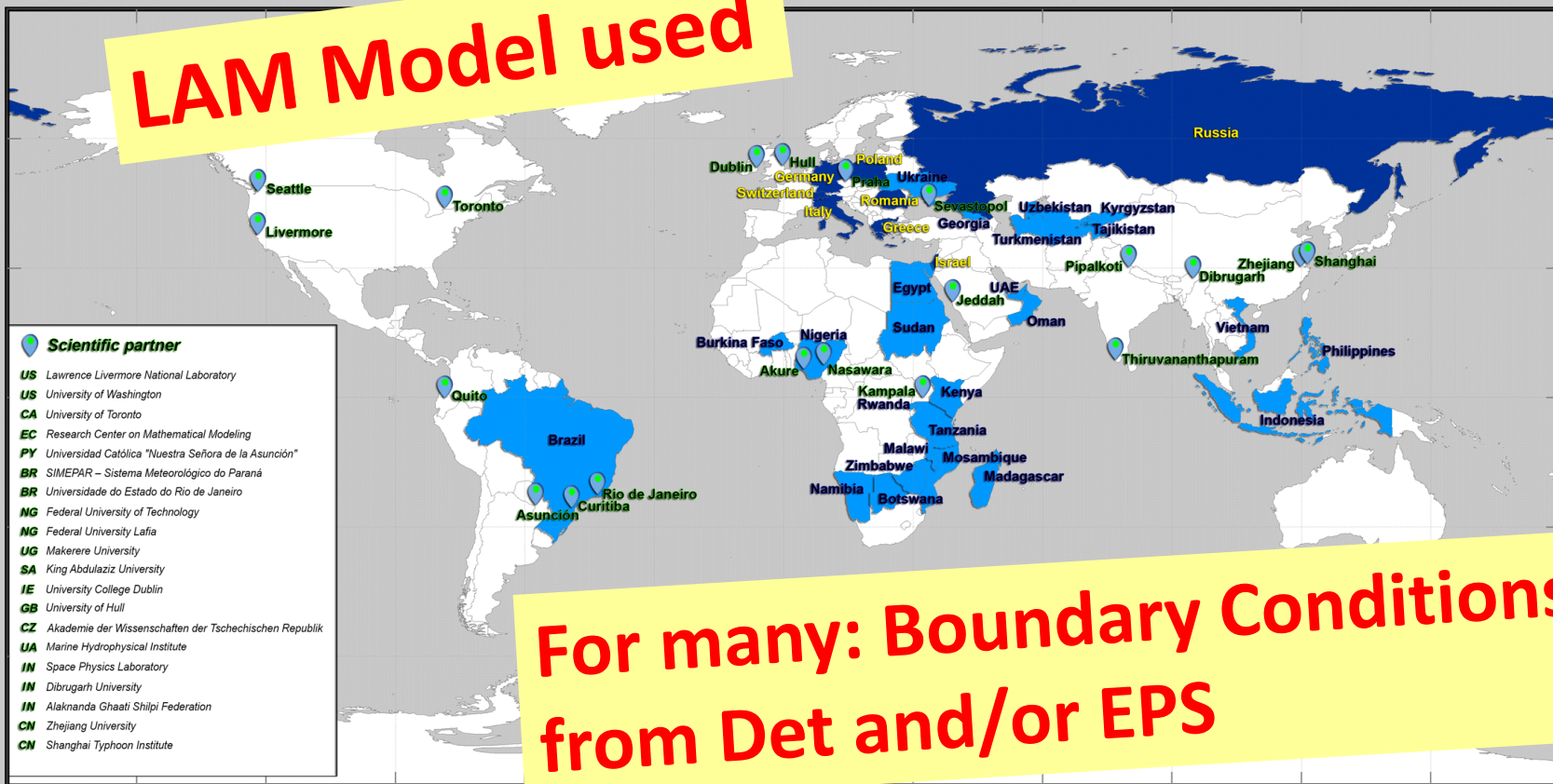
Renawable Energy Forecasting

# Framework Numerical Weather Prediction



40 Countries

LAM Model used



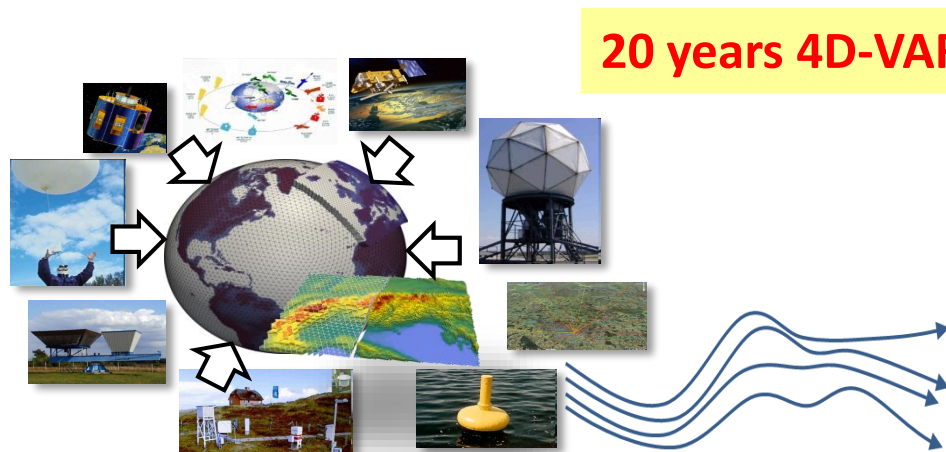
For many: Boundary Conditions from Det and/or EPS

Cosmo User Seminar + Training Course + Symposium

1. Why and Where **Distributions, Risk and Uncertainty?**
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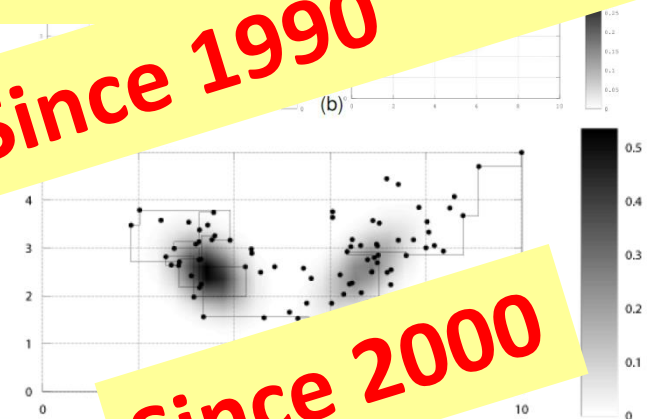


- Why **variational** Data Assimilation (3D/4D-VAR)?



20 years 4D-VAR at ECMWF

Since 1990



- Why **Ensemble** Data Assimilation (EDA)?

Since 2000

- Why **Hybrid** Methods? (3D/4D-EnVAR)

Since 2010

- Why Particle **Filters**? (PF,GPF,ETPF,LAPF,LMCPF)

Since 2020

# Variational Analysis (3D/4D-VAR)

## Recall where we came from ...

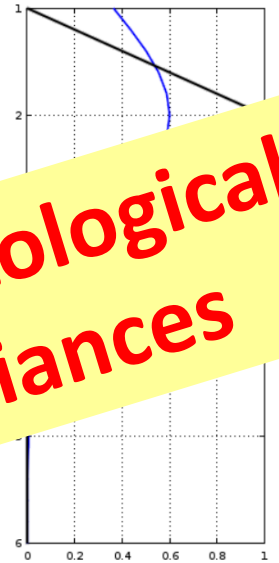
The minimization of

$$J(x) = \|x - x^b\|_{B^{-1}}^2 + \|y - Hx\|_{R^{-1}}^2$$

Obs Operator



Climatological covariances



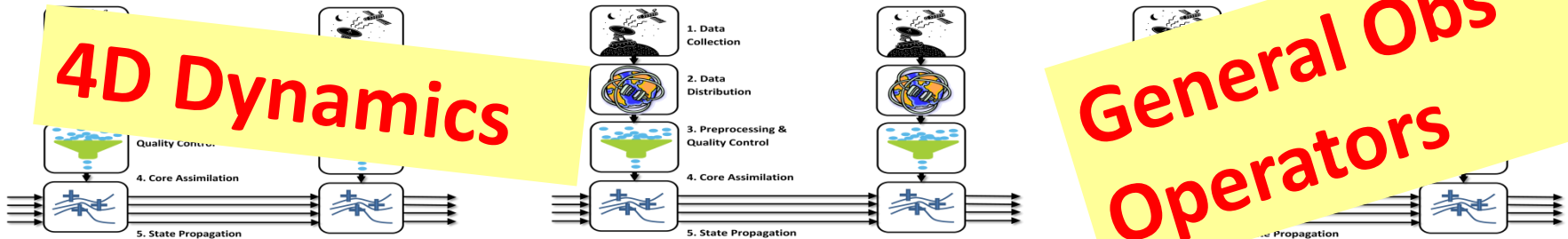
leads to the analysis

Minimization

$$x^a = x^b + B H^T (R + H B H^T)^{-1} (y - H x^b)$$

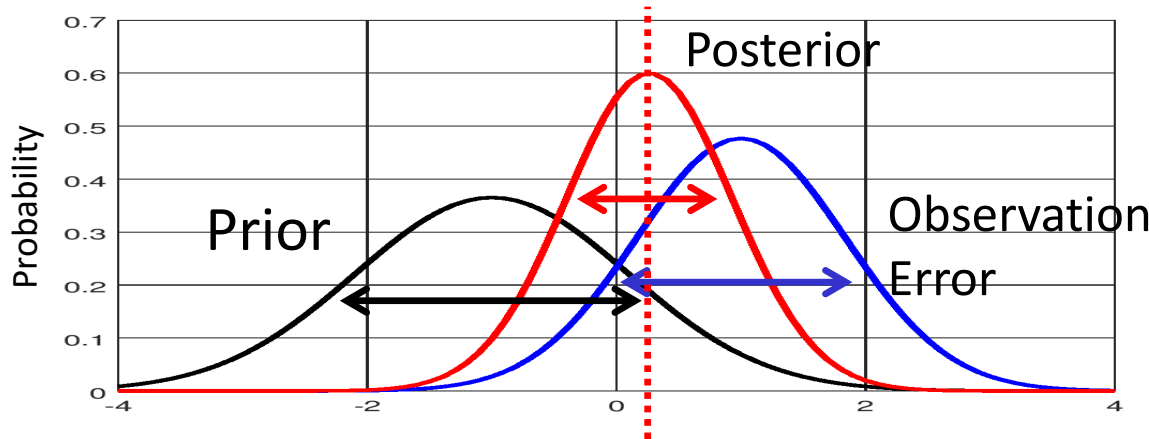
This is the mother of all data assimilation formulas

4D Dynamics



General Obs Operators

# Stochastic View $\Leftrightarrow$ Minimization



**Var == Mean/  
ML estimator!**

Gaussian Prior

$$p(x) = e^{-\frac{1}{2}(x-x^b)^T B^{-1}(x-x^b)}$$

Gaussian Data Error

$$p(y|x) = e^{-\frac{1}{2}(y-Hx)^T R^{-1}(y-Hx)}$$

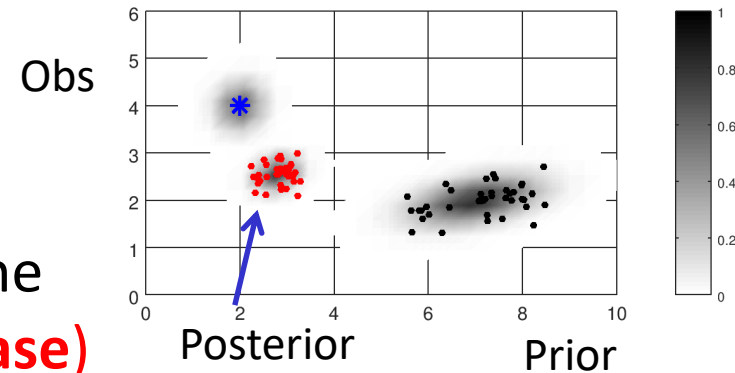
Gaussian Posterior

**Maximum Likelyhood Estimator =  
Minimization of Functional = 3/4DVar**

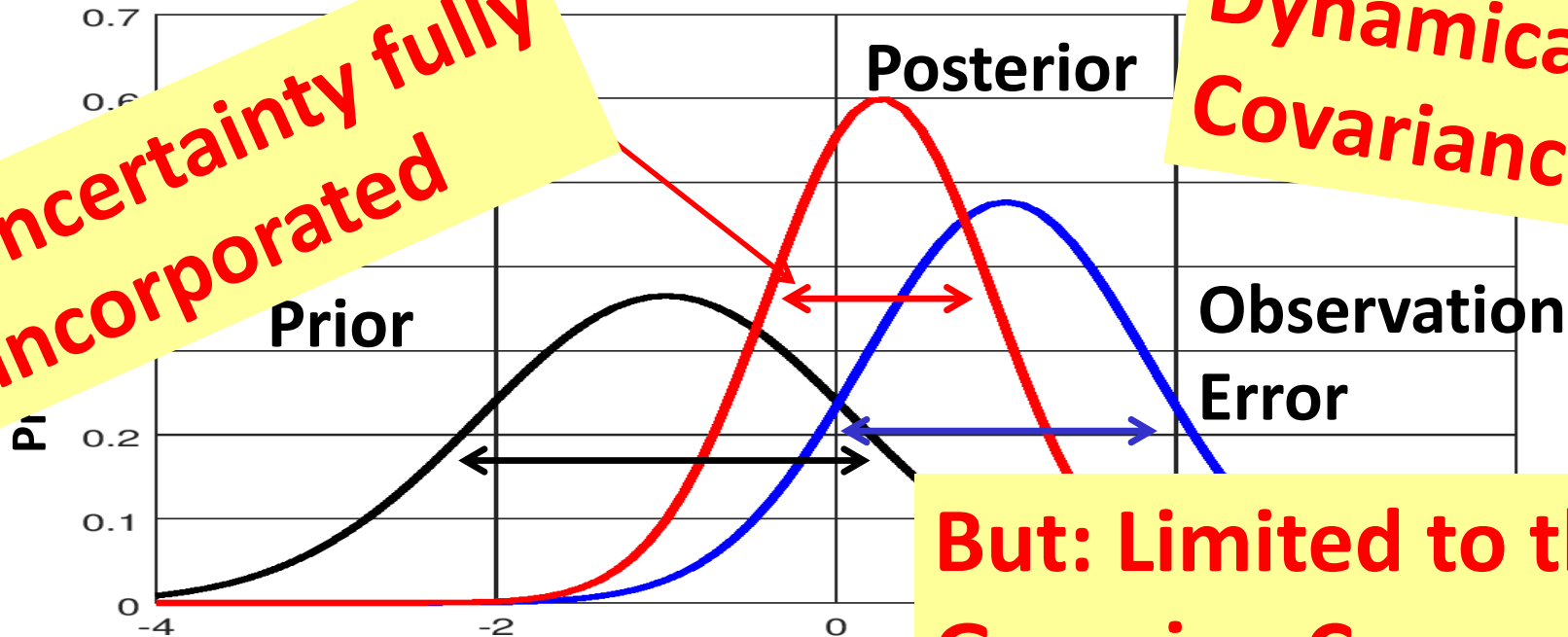
$$p(x|y) = c e^{-\frac{1}{2} \left\{ (x-x^b)^T B^{-1}(x-x^b) + (y-Hx)^T R^{-1}(y-Hx) \right\}}$$

## Basic Idea of the Kalman Filter

- **Sequential** Assimilation of Data
- Do not only adapt the mean, but also the **Covariance B (Uncertainty, Gaussian Case)**



**Uncertainty fully incorporated**

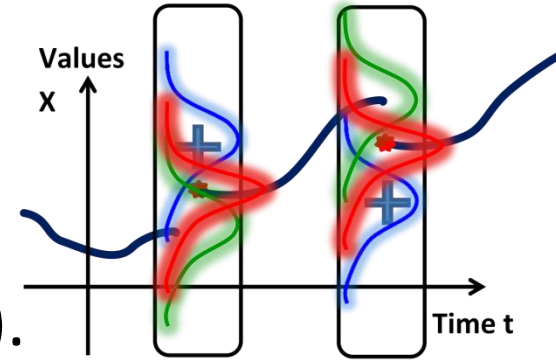


**Dynamical Covariances**

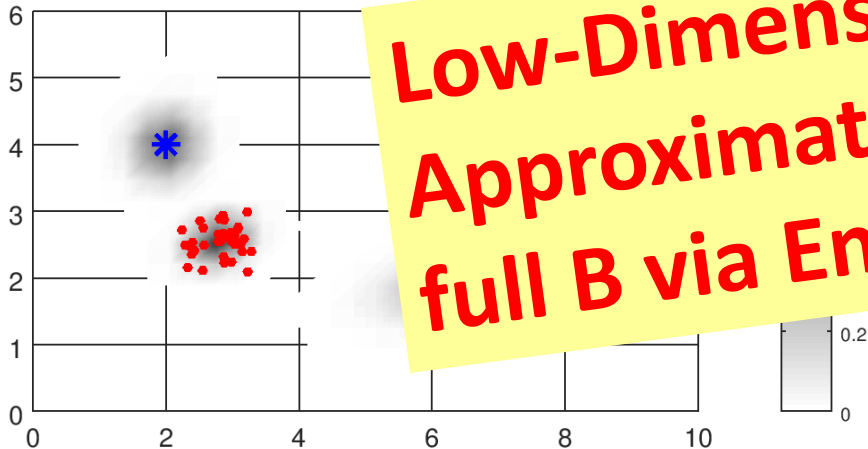
**But: Limited to the Gaussian Case**

# EDA: Ensemble Kalman Filter (EnKF)

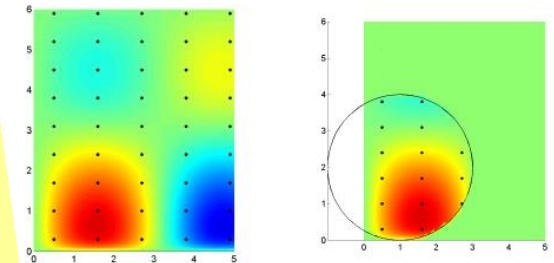
- Kalman Filter needs B update => **expensive!**
- **Estimate** B based on an ensemble of forecasted states (**stochastic estimator**).



B will be **flow-dependent** and variable, depending on the **model dynamics** and on the **observations**



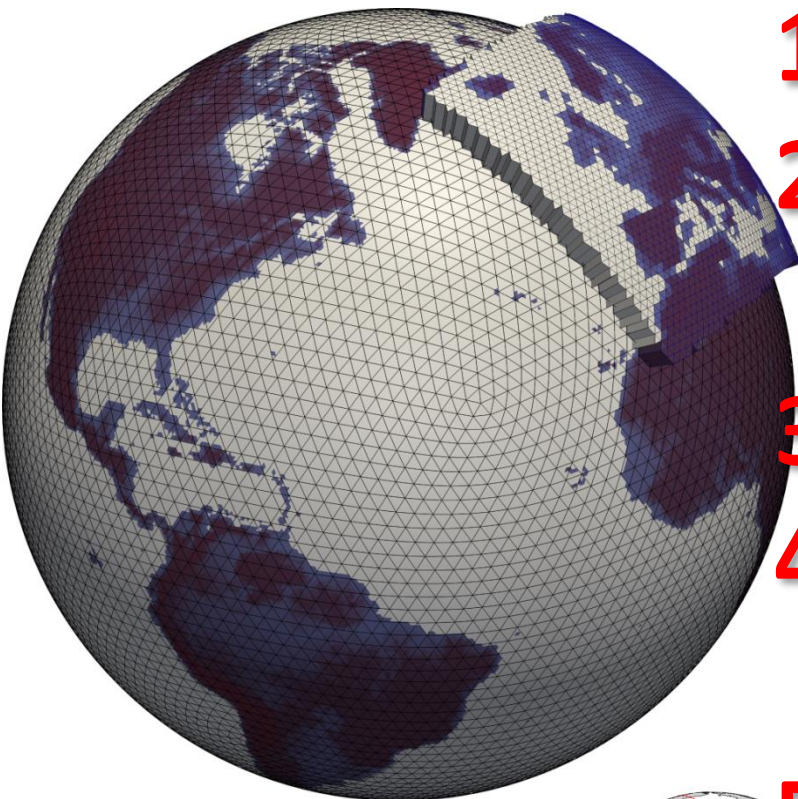
**Low-Dimensional  
Approximation to  
full B via Ensemble!**



**Needs Localization**

**Localized: LETKF**

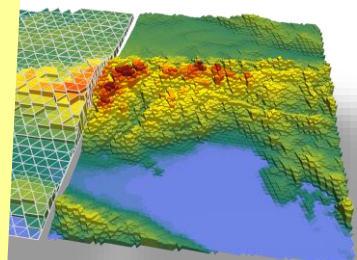
1. Why and Where **Distributions, Risk and Uncertainty?**
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- 1) **ICON Model 13km**
- 2) **Nest over Europe (6.5km; 2-way)**
- 3) **ICON-LAM D2**
- 4) **Nest over Germany (1km; 2-way) D1**
- 5) **NWC Ensemble**



- 3h cycles global+Europe: 180h fc
- 1h ana cycle LAM, 3h fc cycle: 27h fc
- RUC cycle 1h: 6-12h fc



$$\frac{\partial v_n}{\partial t}$$
$$\frac{\partial w}{\partial t}$$
$$\frac{\partial \rho}{\partial t}$$
$$\frac{\partial \rho v_n}{\partial t} + \nabla \cdot (\bar{v} \rho \theta_v) = 0$$

**Conventional and Remote Sensing**

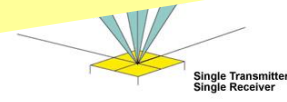
**Exploring Pilot Stations for Boundary Layer Remote Sensing Obs 2019-2025**

Radiosonde

Meteor. Observato  
Lindenberg

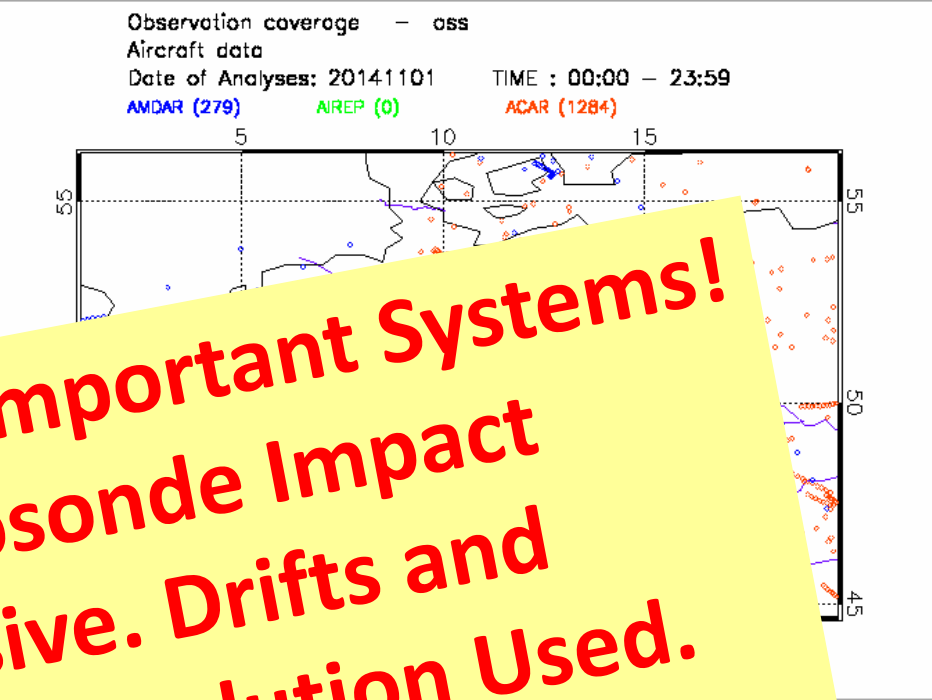
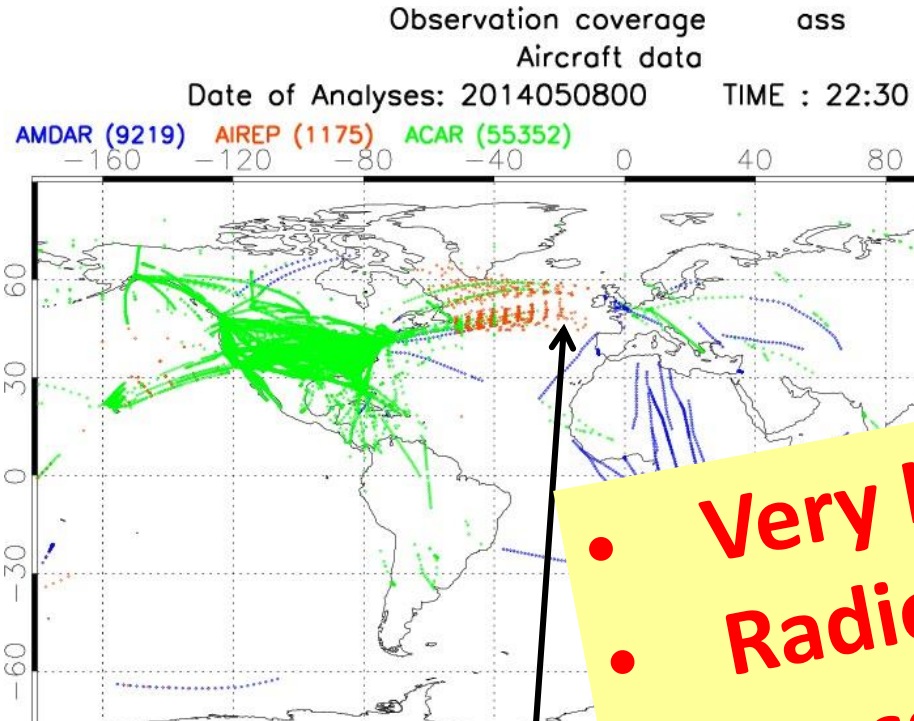
1. Doppler-LIDAR (Wind)
2. DIAL (Humidity)
3. Raman LIDAR (Temp+Hum)
4. MWR (Temp+Hum)
5. GPS STD (Hum)
6. Cloud Radar

Wind-Profiler NCAR/EOL





# Conventional Synop + Airplanes

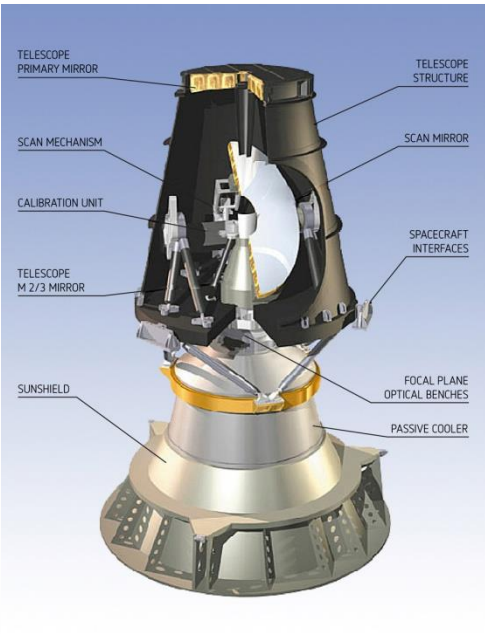


- Very Important Systems!
- Radiosonde Impact massive. Drifts and High-Resolution Used.
- Mode-S Operational over Germany (KENDA)

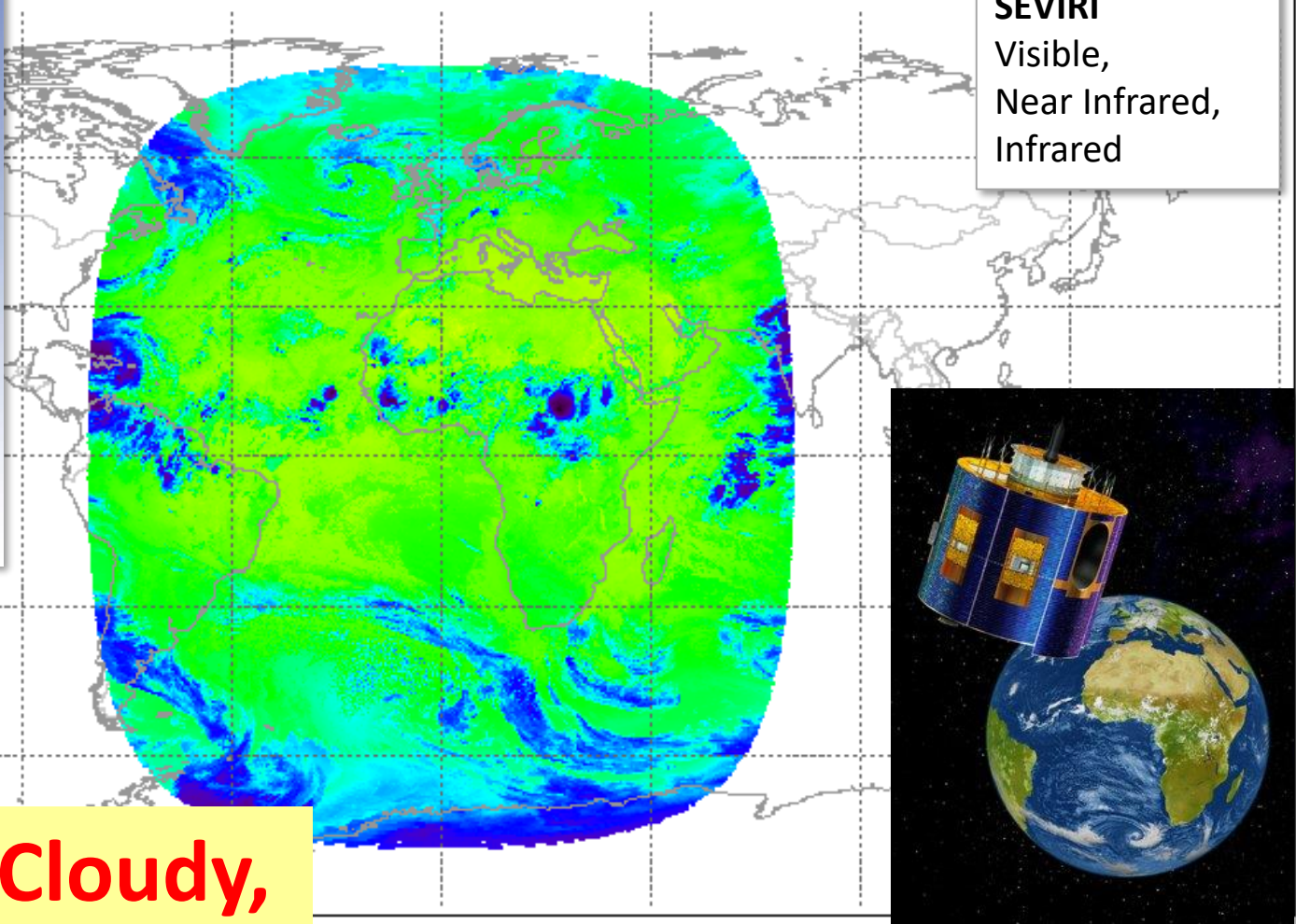


# Observations: Geostationary Satellites

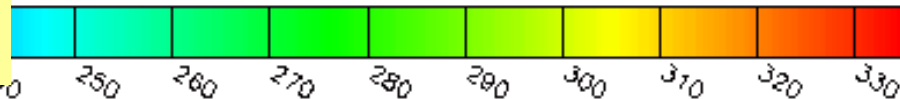
Bild: Robin Faulwetter



**SEVIRI**  
Visible,  
Near Infrared,  
Infrared



**Clear and Cloudy,  
IR, VIS (MFASIS)**



# Observations: Polar Orbiting Satellites

Deutscher Wetterdienst  
Wetter und Klima aus einer Hand



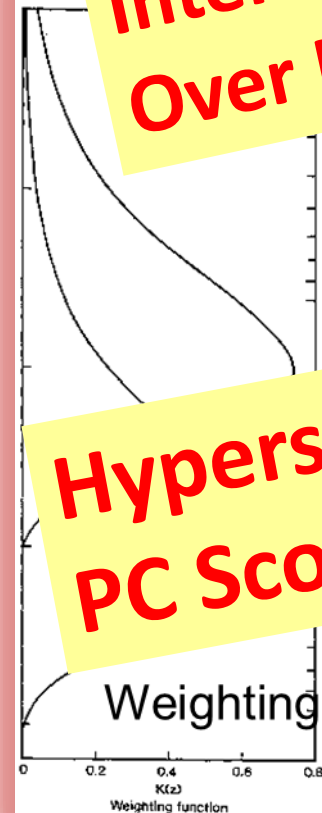
**MW and IR Sounder  
and Imager  
Research Work:  
Interchannel Corr +  
Over Land + All-Sky**

**Aeolus Wind Lidar: Cal-Val +  
Assimilation tests**

**Hyperspectral:  
PC Scores Assimilation**

**EPS-SG coming ...**

- **Currently (Sept 2018):  
3 permanent positions  
advertised**



# Ensemble Datenassimilation EnVar

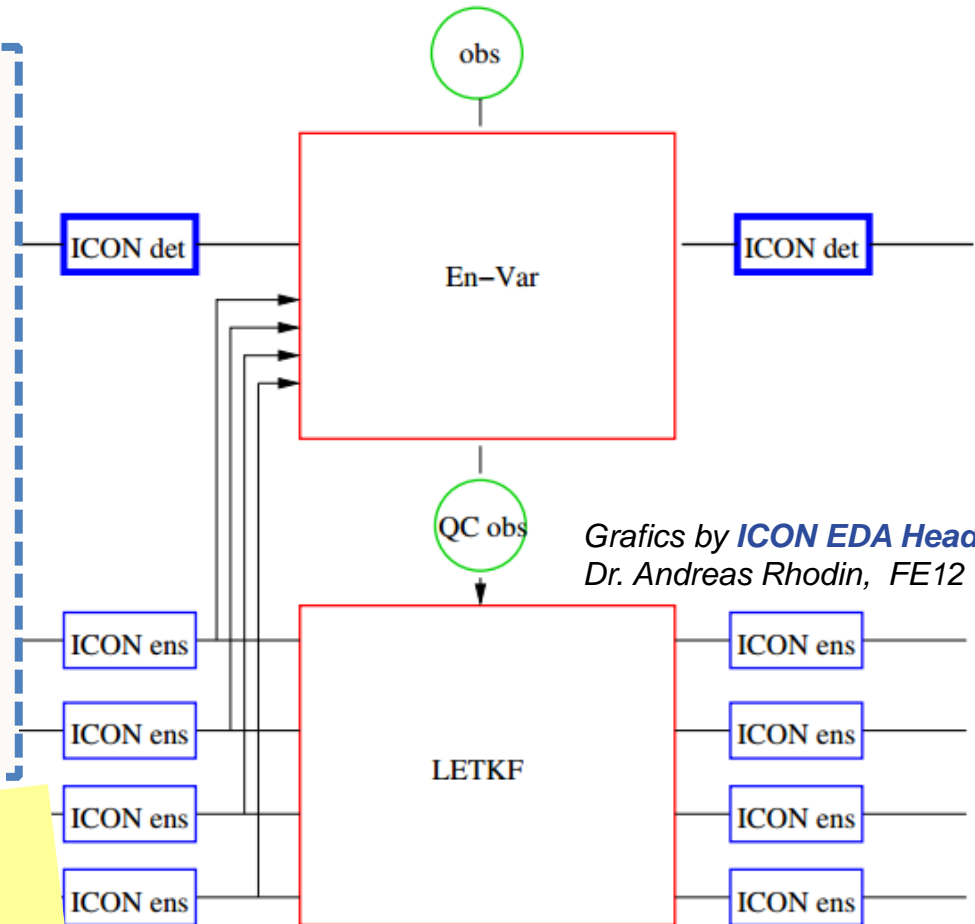
Operational since January 2016

Deutscher Wetterdienst  
Wetter und Klima aus einer Hand



We are running **ICON EDA** in our Routine since Jan 2016

- 40 Members each with 40km global resolution and 20km NEST over Europe
- 1 deterministic 13km/6.5km
- **EPS forecasts** 40 Members 7 Days + 1 Deterministic
- Output for convective-scale EDA/EPS
- **Hybrid System**



Grafics by **ICON EDA Head**  
Dr. Andreas Rhodin, FE12

**Hybrid Ensemble – Variational, 3h cycle**



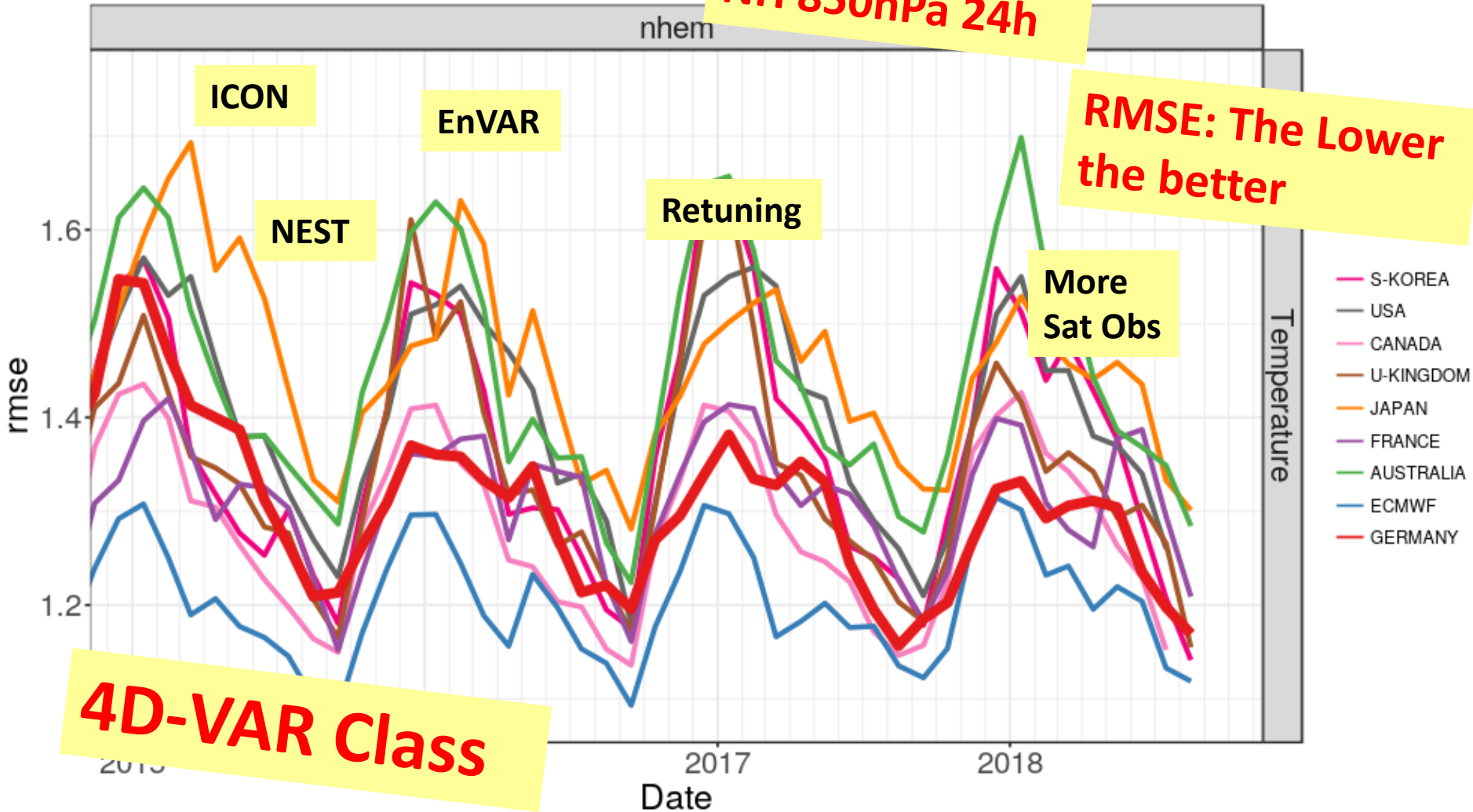
# Hybrid Methods: EnVAR Scores



WMO verification against observations  
lead-time: 24h  
valid-time: 12UTC  
level: 850hPa

Temperature  
forecast quality  
NH 850hPa 24h

RMSE: The Lower  
the better



# Hybrid Methods: EnVAR Scores



Temperature  
forecast quality,  
NH 500hPa 72h

WMO verification against observations  
lead-time: 72h  
valid-time: 12UTC  
level: 500hPa

RMSE: The Lower  
the better

More  
Sat Obs

nhem

ICON

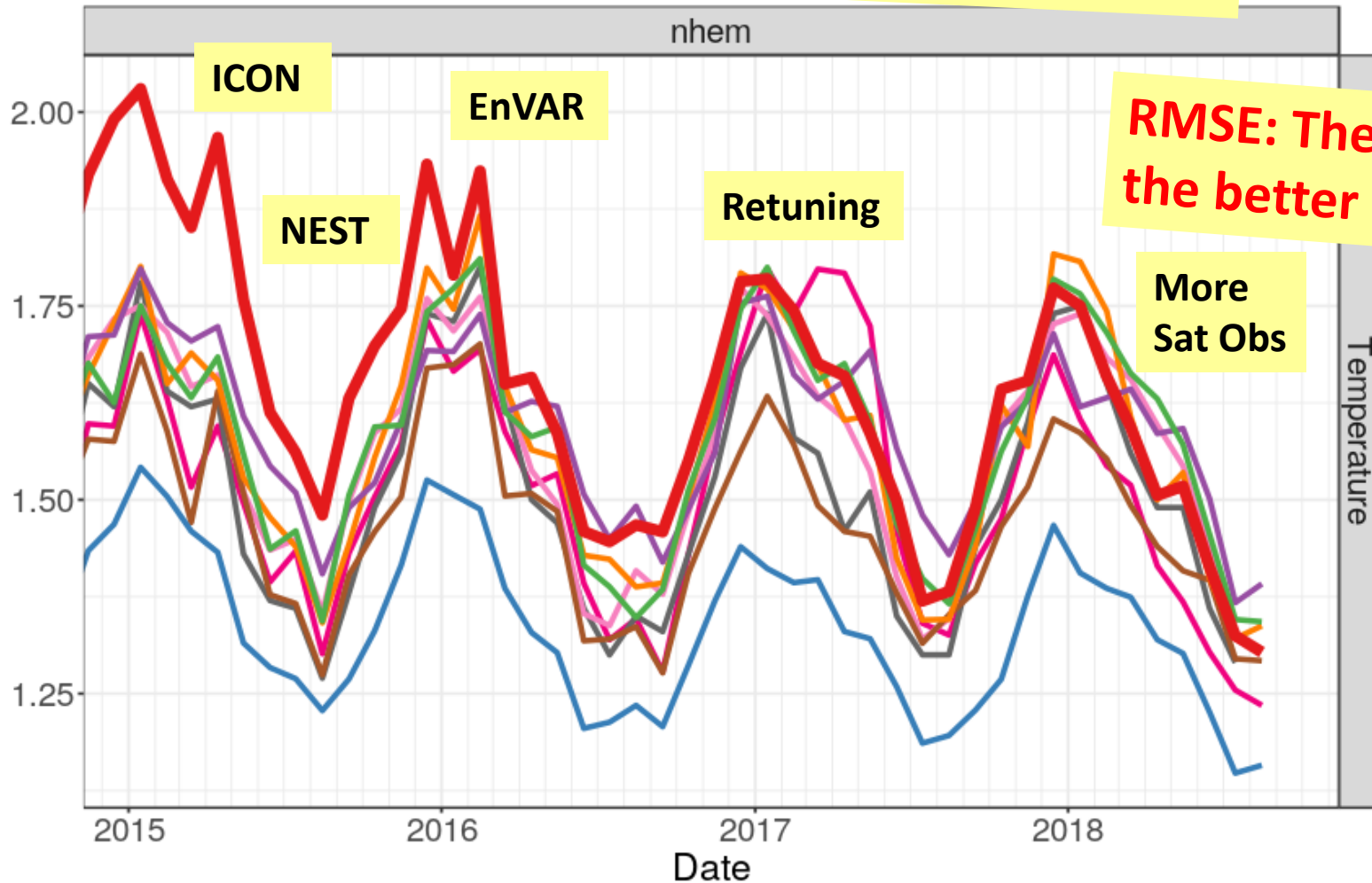
EnVAR

NEST

Retuning

Temperature

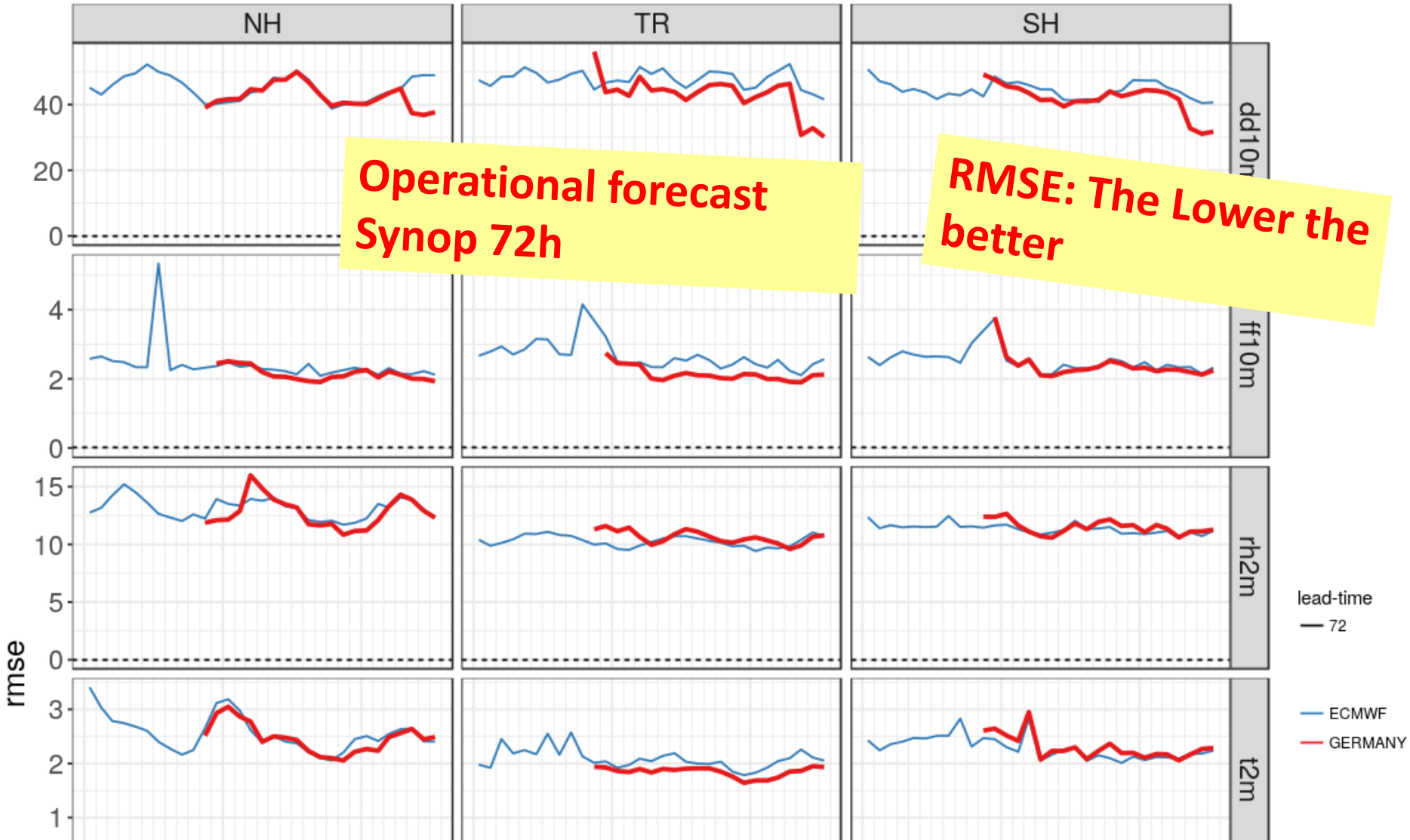
- S-KOREA
- USA
- CANADA
- U-KINGDOM
- JAPAN
- FRANCE
- AUSTRALIA
- ECMWF
- GERMANY



# Hybrid Methods: EnVAR Scores

WMO verification against SYNOP  
lead-time: 72h  
valid-time: 12UTC

2017+2018

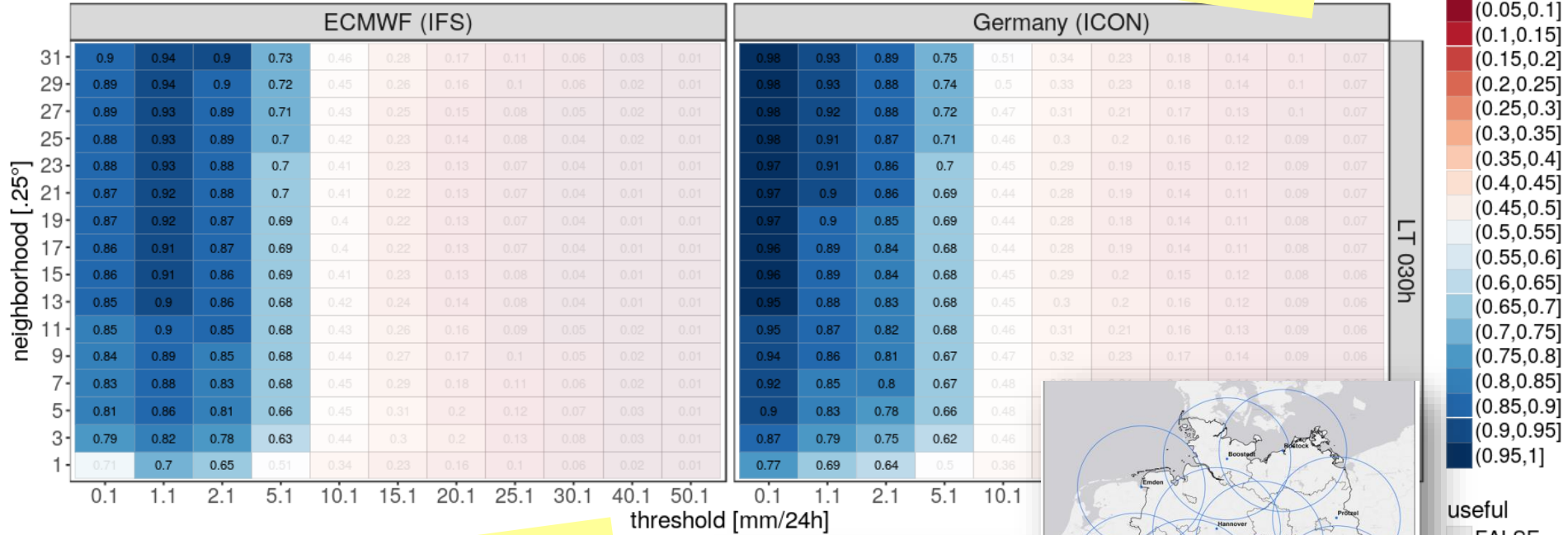


# Hybrid Methods: EnVAR Scores

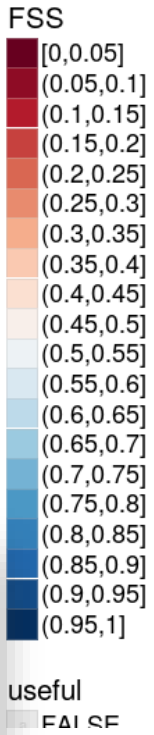
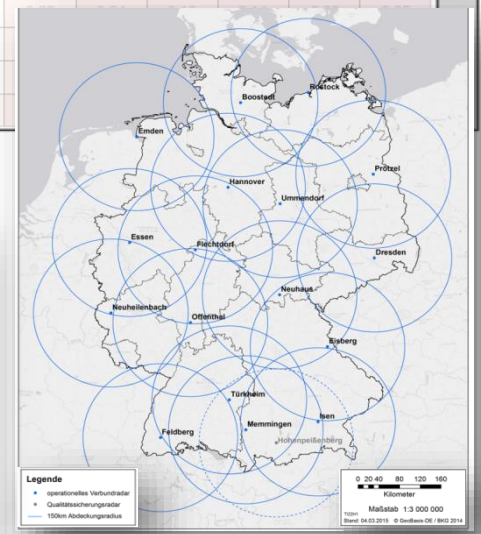
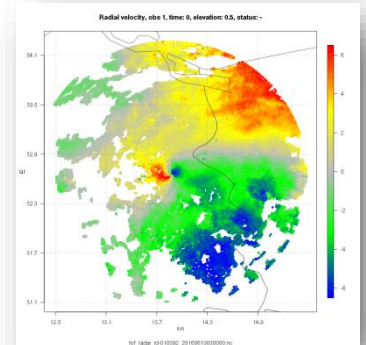
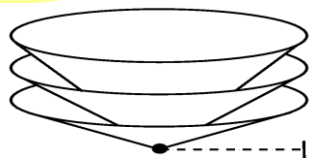


**FSS: the higher the better**

OBS: german radar (rw) 06UTC-06UTC  
 PRED: rr\_24h (0.25°)  
 INI: 00UTC  
 CASES: ECMWF (IFS) 90, Germany (ICON) 90



**Operational forecast  
Precipitation vs RADAR**





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PRIOR

DATA

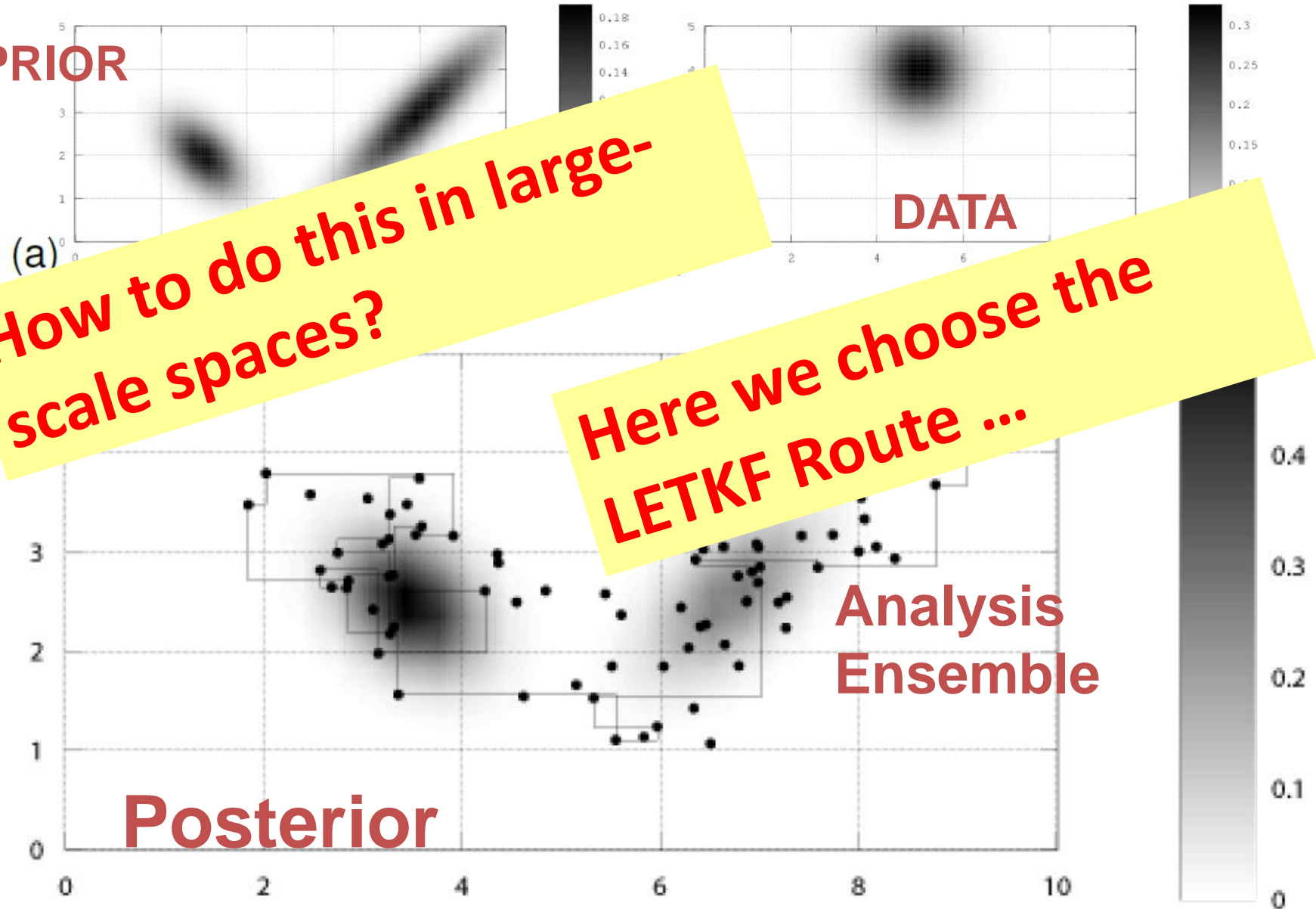
How to do this in large-scale spaces?

Here we choose the LETKF Route ...

Analysis Ensemble

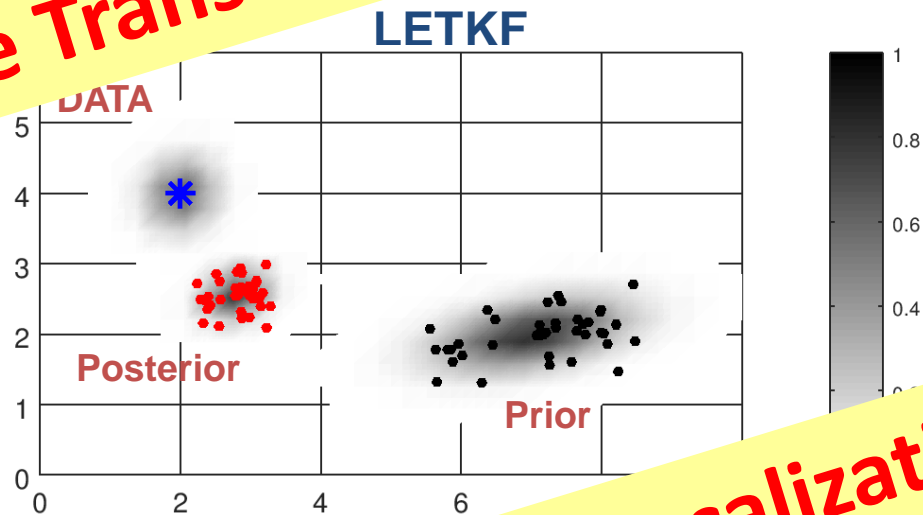
Posterior

(c)



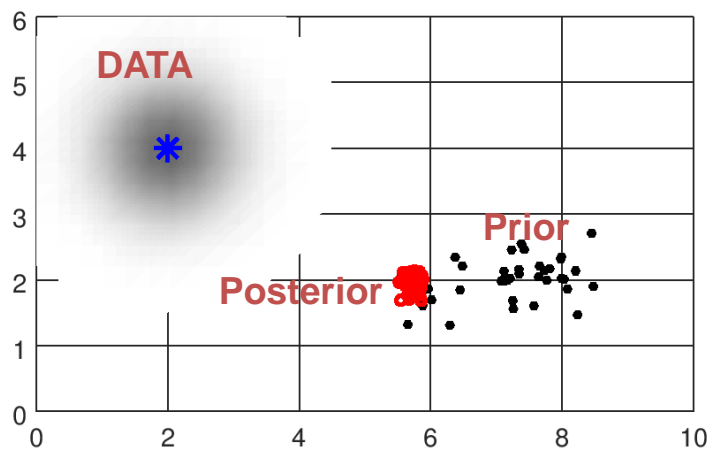
# Bayesian Filtering via PF

1) Ensemble Transform

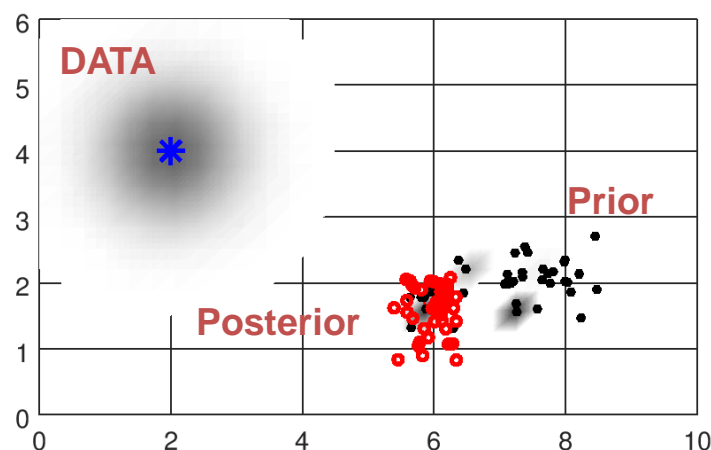


2) Localization

Classical PF



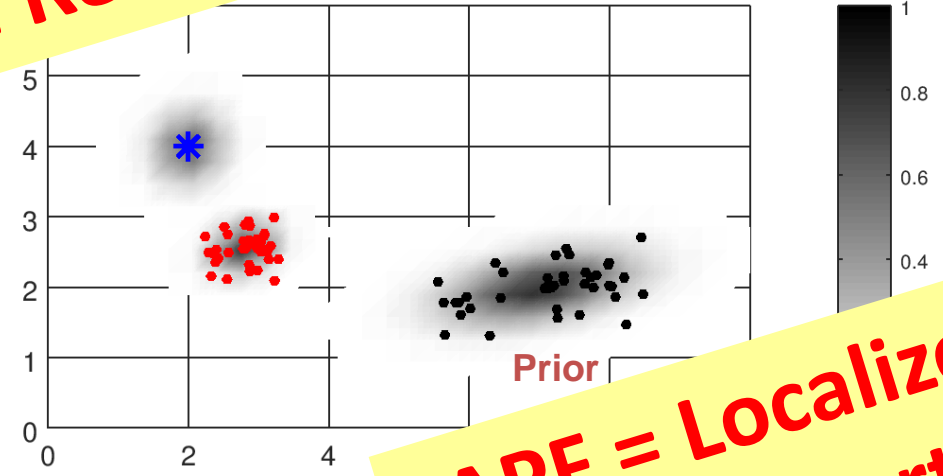
LAPF



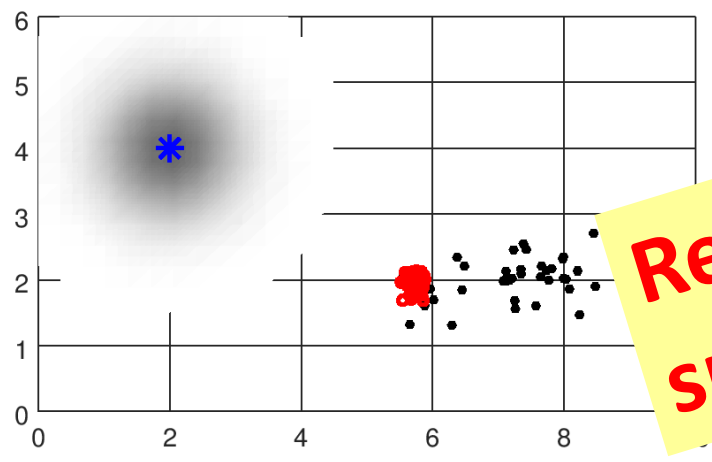
# Bayesian Filtering via PF

## 3) Adaptive Resampling

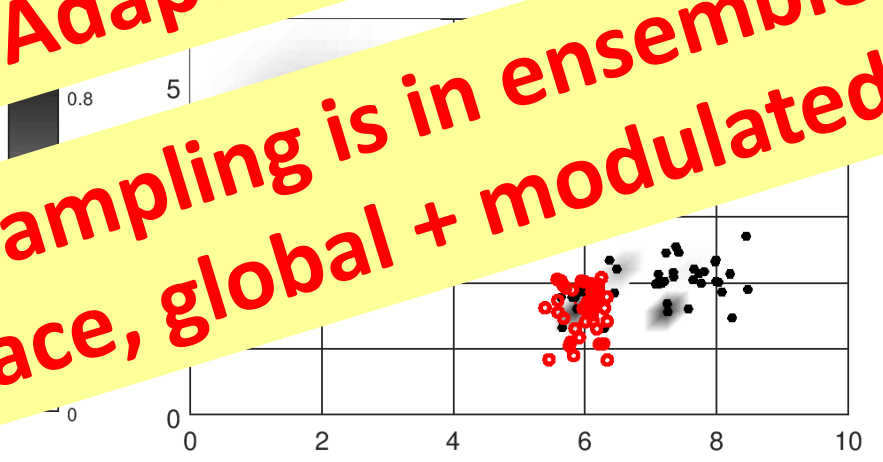
LETKF



Classical PF



LAPF



LAPF = Localized  
Adaptive Particle Filter

Resampling is in ensemble  
space, global + modulated

LAPF = Transform, Localization, Adaptivity with global modulated Resampling

- **Bayes formula** to calculate new analysis distribution

$$p_k^{(a)}(x) := p(x|y_k) = c p(y_k|x) p_k^{(b)}(x), \quad x \in \mathbb{R}^n$$

$c$  is a normalization factor:  $\int_X p_k^{(a)}(x) dx = 1$

## Classical PF Approach

- To carry out the analysis step at time  $t_k$   
**aposteriori weights**  $p_k^{(a)}$  are calculated

$$p_{k,l}^{(a)} = c e^{-\frac{1}{2}(y-Hx^{(l)})^T R^{-1}(y-Hx^{(l)})}$$

$c$  is chosen such that  $\sum_{l=1}^L p_{k,l}^{(a)} = L$

- **Accumulated weights**  $w_{ac}$  are defined:

$$w_{ac_0} = 0$$
$$w_{ac_i} = w_{ac_{i-1}} + p_i^a, \quad i = 1, \dots, L$$

where  $L$  denotes the ensemble size

- Drawing  $r_j \sim U([0,1])$ ,  $j = 1, \dots, L$ , set  $R_j = j - 1 + r_j$  and define **transform matrix**  $W$  for the particles by:

$$W_{i,j} = \begin{cases} 1 & \text{if } R_j \in (w_{ac_{i-1}}, w_{ac_i}], \\ 0 & \text{otherwise,} \end{cases}$$

$i, j = 1, \dots, L$  with  $W \in \mathbb{R}^{L \times L}$ ,  $(s, t]$  denotes the interval of values  $s < \eta \leq t$ .

**Resampling**

### Adaptivity based on o-b statistics

- Based on the adaptive multiplicative **inflation factor**  $\rho$  determined by the LETKF

$$\rho = \frac{\text{E} \left[ \mathbf{d}_{o-b}^T \mathbf{d}_{o-b} \right] - \text{Tr}(\mathbf{R})}{\text{Tr}(\mathbf{H}\mathbf{P}^b\mathbf{H}^T)}$$

- **Weighting factor**  $\alpha$  has been chosen, due to the small ensemble size ( $L = 40$ )

$$\rho_k = \alpha \tilde{\rho}_k + (1 - \alpha) \rho_{k-1}$$

- **Perturbation factor  $\sigma$**  is used to add spread to the system

$$\sigma = \left\{ \begin{array}{ll} c_0, & \rho < \rho^{(0)} \\ c_0 + (c_1 - c_0) * \frac{\rho - \rho^{(0)}}{\rho^{(1)} - \rho^{(0)}}, & \rho^{(0)} \leq \rho \leq \rho^{(1)} \\ c_1, & \rho > \rho^{(1)} \end{array} \right\}$$

where  $c_0 = 0.02$ ,  $c_1 = 0.2$ ,

$\rho^{(0)} = 1.0$  and  $\rho^{(1)} = 1.4$ , with

$\sigma = c_1$  if  $\rho \geq \rho^{(1)}$  and

$\sigma = c_0$  if  $\rho \leq \rho^{(0)}$

**Enforce the  
desired spread!**

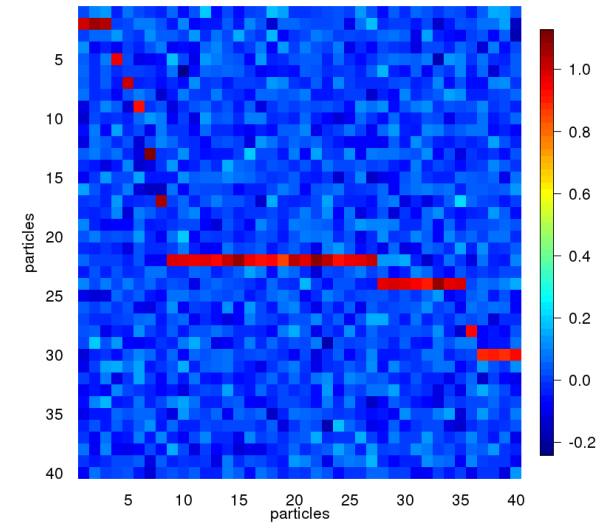
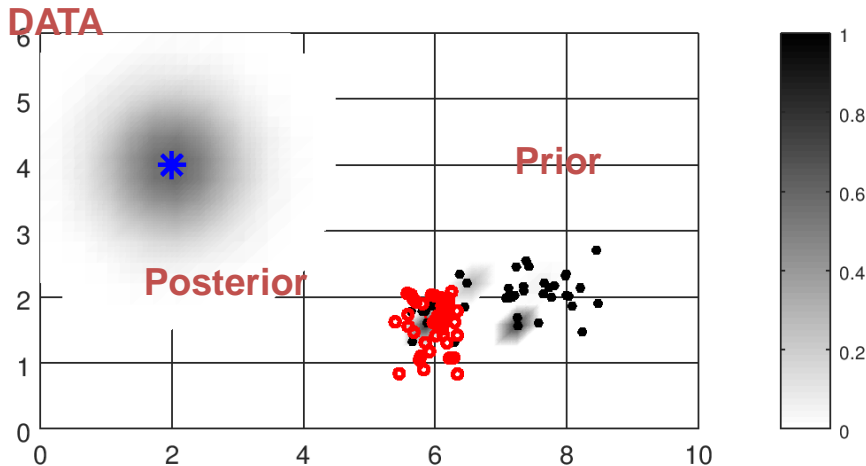


# Fourth Step: Gaussian Resampling

- Weights  $W$  are modified by applying the **perturbation factor**  $\sigma$

$$W = W + R_{nd} * \sigma$$

with  $R_{nd}$  normally distributed random numbers



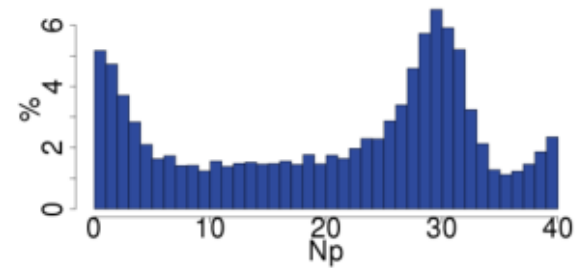
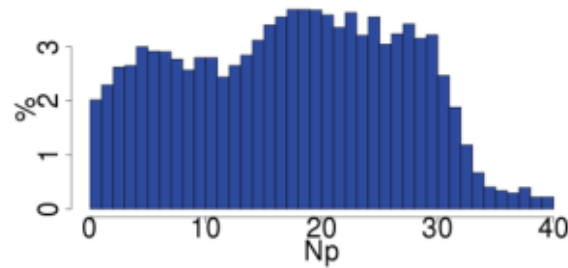
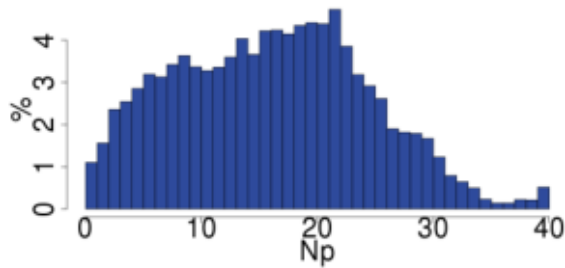
An example for a W-Matrix after applying  $\sigma$  determined with  $\sigma = 0.1$  for 60%

**Enforce the desired spread!**

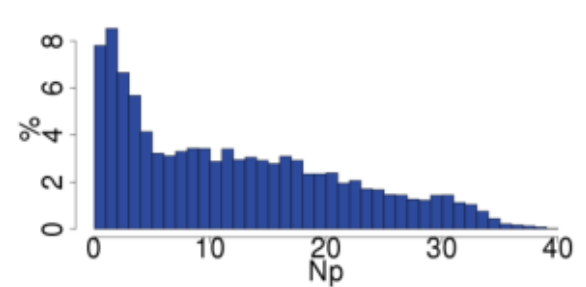
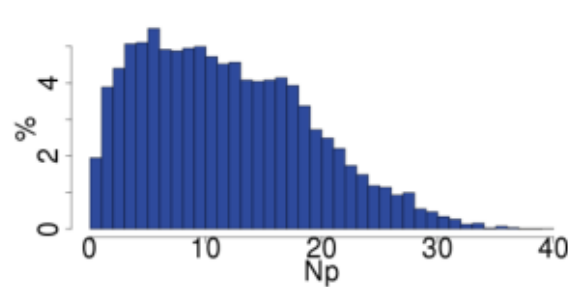
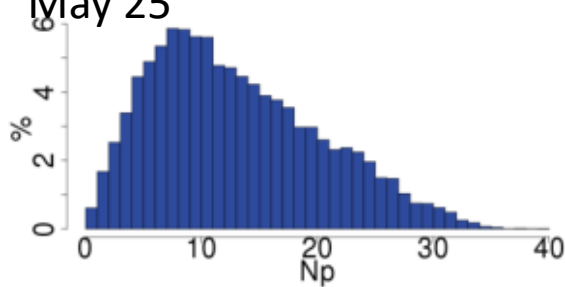
are chosen

# Effective Ensemble Size Distributions

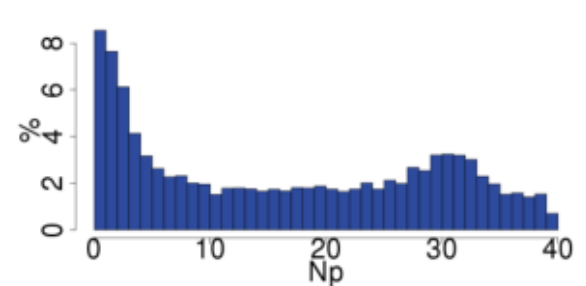
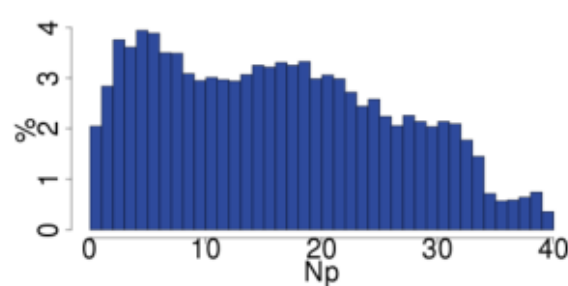
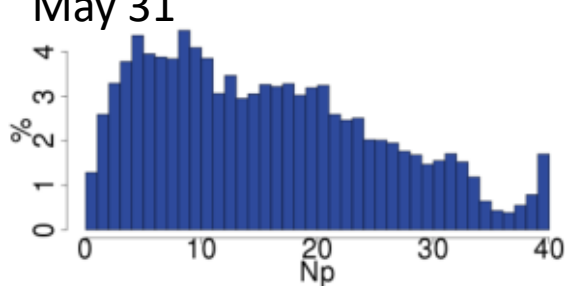
May 20



May 25



May 31



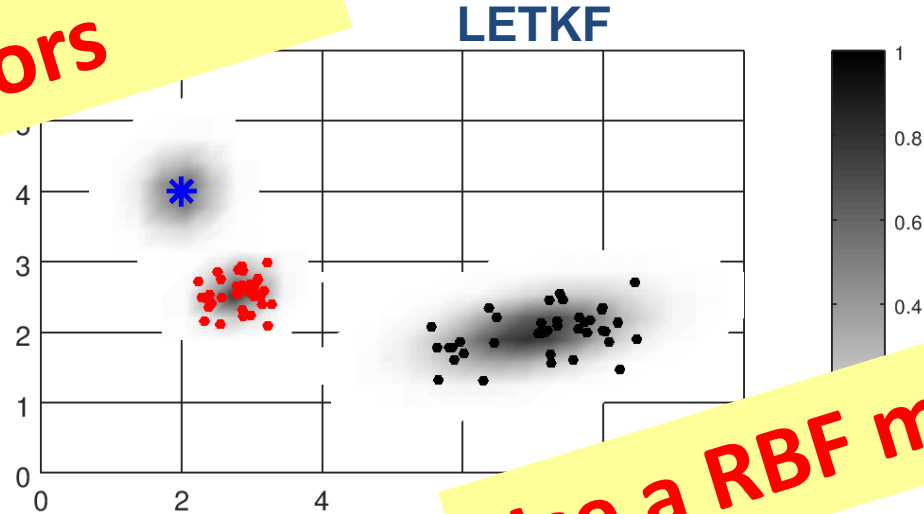
100 hPa

500 hPa

1000 hPa

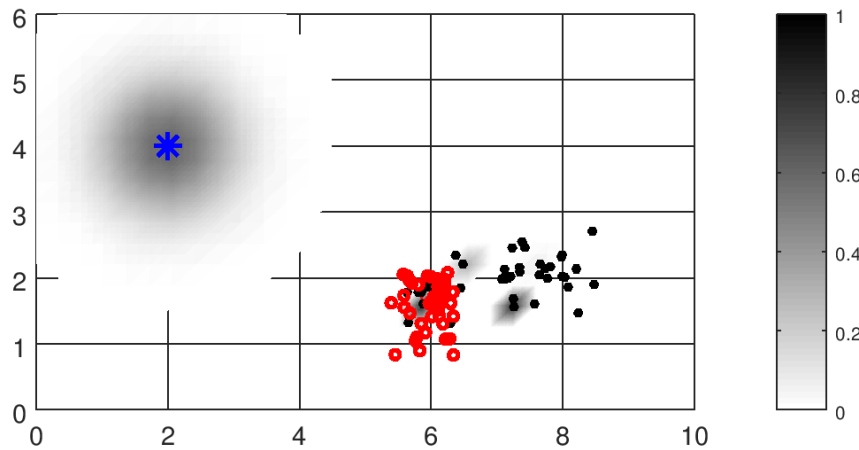
# LAPF versus LMCPF

**LAPF has problems  
with model errors**

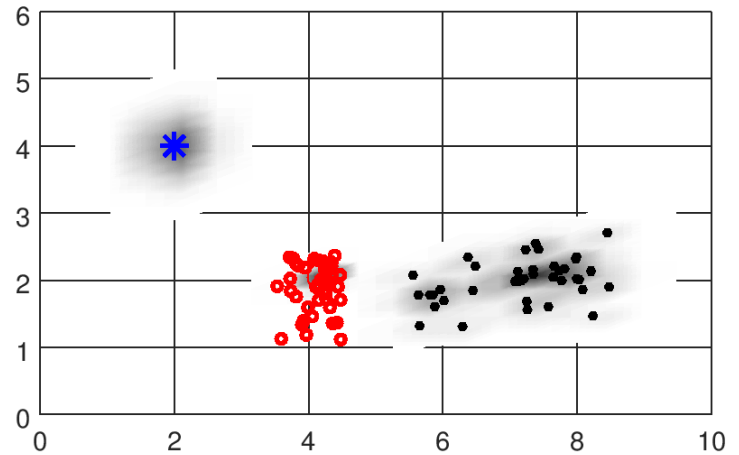


**Use a RBF mixture**

### LAPF



### LMCPF



**LMCPF = Transform, Localization, RBF mixture, Adaptivity**

## ■ Kalman Filter

$$x^{(a)} = x^{(b)} + BH^T(R + HBH^T)^{-1}(y - Hx^{(b)})$$

$$K = BH^T(R + HBH^T)^{-1} \quad \tilde{B} = (I - KH)B$$

## ■ Ensemble B Estimator

$$\bar{x} := \frac{1}{L} \sum_{\ell=1}^L x^{(\ell)}$$

$$B = \frac{1}{L-1} X X^T$$

$$X = (x^{(1)} - \bar{x}, \dots, x^{(L)} - \bar{x}) \in \mathbb{R}^{n \times L}.$$

# LMCPF Basics: B Posterior

$$Y := HX$$

$$\begin{aligned}\tilde{B} &= (I - KH)B \\ &= (I - BH^T(R + HBH^T)^{-1}H)B \\ &= \left(I - \gamma X X^T H^T (R + \gamma HX X^T H^T)^{-1}H\right) \gamma X X^T \\ &= X \left(I - \gamma Y^T (R + \gamma Y Y^T)^{-1}Y\right) \gamma X^T \\ &= X \left(I - \gamma (I + \gamma Y^T R^{-1}Y)^{-1}Y^T R^{-1}Y\right) \gamma X^T \\ &= X \left((I + \gamma Y^T R^{-1}Y)^{-1} (I + \gamma Y^T R^{-1}Y - \gamma Y^T R^{-1}Y)\right) \gamma X^T \\ &= X (I + \gamma Y^T R^{-1}Y)^{-1} \gamma X^T \\ &= X \left(\frac{1}{\gamma} I + Y^T R^{-1}Y\right)^{-1} X^T\end{aligned}$$

**RBF Basis Function  
in Ensemble Space**

$$Y^T (R + \gamma Y Y^T)^{-1} = (I + \gamma Y^T R^{-1}Y)^{-1} Y^T R^{-1}$$

$$p(x|y) = cp(y|x) \cdot p(x)$$

**Gaussian Mixture Case**

$$= ce^{-\frac{1}{2}(y-Hx)^T R^{-1}(y-Hx)} \sum_{\ell=1}^L p_G(x - x^{(\ell)})$$

$$= c \sum_{\ell=1}^L e^{-\frac{1}{2}(y-Hx)^T R^{-1}(y-Hx)} p_G(x - x^{(\ell)}),$$

$$p_G(x - x^{(\ell)}) = \tilde{c} e^{-\frac{1}{2}(x-x^{(\ell)})^T G^{-1}(x-x^{(\ell)})}$$

**Explicit Calculations  
possible for each term**

**We need a  
selection based on  
relative weights!**

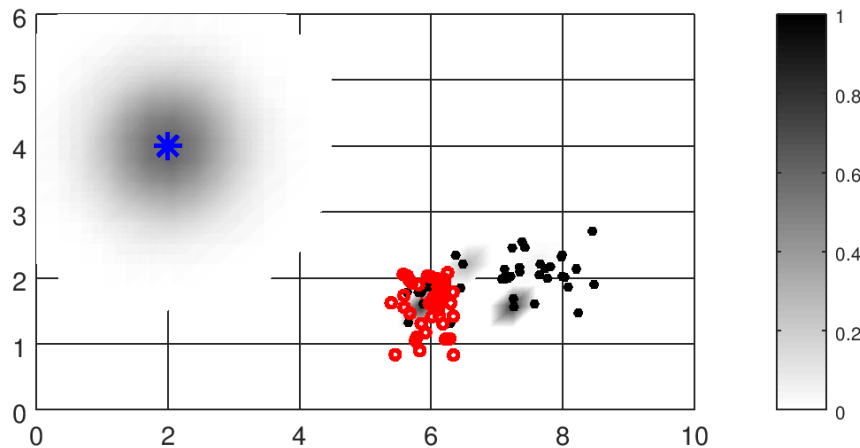
# LMCPF Basics: Relative Weights

$$w_\ell := e^{-\frac{1}{2}(y-Hx^{(\ell)})^T R^{-1}(y-Hx^{(\ell)})}, \quad \ell = 1, \dots, L$$

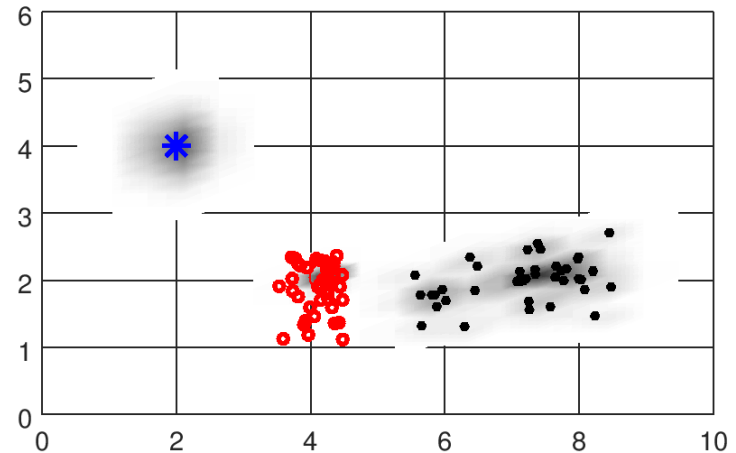
$$w_{tot} := \sum_{\ell=1}^L w_\ell.$$

**We need a selection based on relative weights!**

LAPF



LMCPF



## Projection onto Ensemble Space

Abbreviating  $A := \mathbf{Y}^T \mathbf{R}^{-1} \mathbf{Y}$  and  $C := A^{-1} \mathbf{Y}^T \mathbf{R}^{-1} (\mathbf{y}^o - \bar{\mathbf{y}}^b)$

### Projection Operator

$$P(\mathbf{y}^o - \bar{\mathbf{y}}^b) = \mathbf{Y}(\mathbf{Y}^T \mathbf{R}^{-1} \mathbf{Y})^{-1} \mathbf{Y}^T \mathbf{R}^{-1} (\mathbf{y}^o - \bar{\mathbf{y}}^b),$$

### Projected discrepancy

$$\begin{aligned} P(\mathbf{y}^o - H\mathbf{x}^{(\ell)}) &= \mathbf{Y}A^{-1}\mathbf{Y}^T\mathbf{R}^{-1}((\mathbf{y}^o - \bar{\mathbf{y}}^b) - \mathbf{Y}e_\ell) \\ &= \mathbf{Y}(C - e_\ell), \quad \ell = 1, \dots, L. \end{aligned}$$

### Exponent

$$P(\mathbf{y}^o - H\mathbf{x}^{(\ell)})^T \mathbf{R}^{-1} P(\mathbf{y}^o - H\mathbf{x}^{(\ell)}) = [C - e_\ell]^T A [C - e_\ell], \quad \ell = 1, \dots, L,$$

### Weight

$$w_{k,\ell} = ce^{-\frac{1}{2}[C - e_\ell]^T A [C - e_\ell]}, \quad \ell = 1, \dots, L.$$



## Classical versus projected weights

$$\begin{aligned}w_{k,\ell}^{classical} &= e^{-\frac{1}{2}[(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]} \\ &= e^{-\frac{1}{2}[(P+(I-P))(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[(P+(I-P))(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]} \\ &= e^{-\frac{1}{2}[P(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[P(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]} \cdot \underbrace{e^{-\frac{1}{2}[(I-P)(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]^T \mathbf{R}^{-1}[(I-P)(\mathbf{y}^o - H\mathbf{x}^{(\ell)})]}}_{=\tilde{c}},\end{aligned}$$

Factor is a constant term, since we have

$$\begin{aligned}(I-P)(\mathbf{y}^o - H\mathbf{x}^{(\ell)}) &= (I-P)(\mathbf{y}^o - \bar{\mathbf{y}}^b + \mathbf{Y}e_\ell) \\ &= (I-P)(\mathbf{y}^o - \bar{\mathbf{y}}^b) - \underbrace{(I-P)\mathbf{Y}e_\ell}_{=0}.\end{aligned}$$

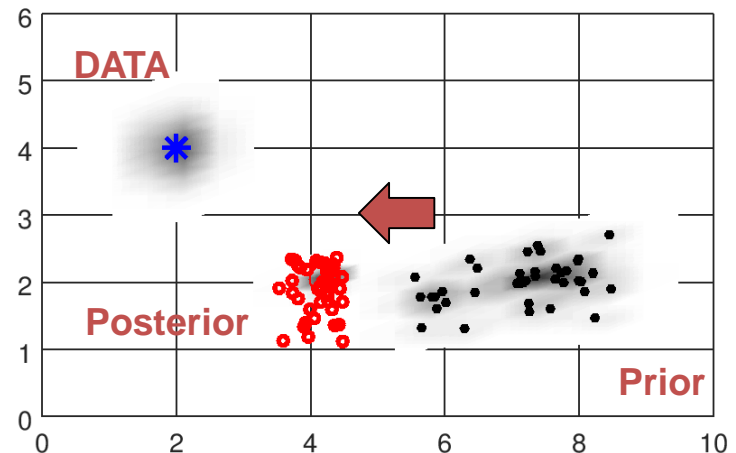
Projected particle filter weights and classical particle filter weights are equivalent theoretically, but numerically remove a very small common factor

## LMCPF = Local Markov Chain Particle Filter

- **Weights  $W$**  are calculated by drawing from the posterior

$$W = W + A_{shift} * W + B_{post} * R_{nd} * \sigma$$

with  $R_{nd}$  normally distributed random numbers,  
 $A_{shift}$  and  $B_{post}$  calculated with Gaussian radial basis function (rbf) Approximation for prior density and observation error



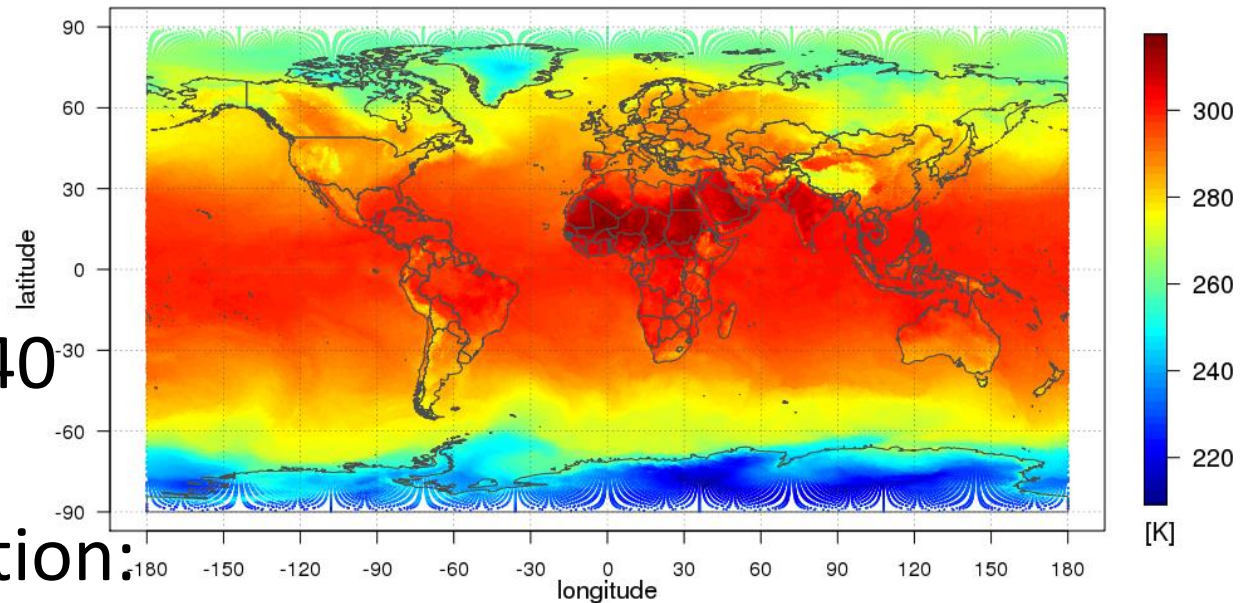
- ✓ It is an **explicit calculation of the Bayes posterior based on radial basis function approximation of the prior, with subsequent draws from that distribution in the MCMC sense.**

# Large-Scale Experimental Set-up

Deutscher Wetterdienst  
Wetter und Klima aus einer Hand



- Full ensemble: 40 members
- Reduced resolution:
  - 26km deterministic
  - 52km ensembles
- Period:  
01.05.2016 –  
31.05.2016

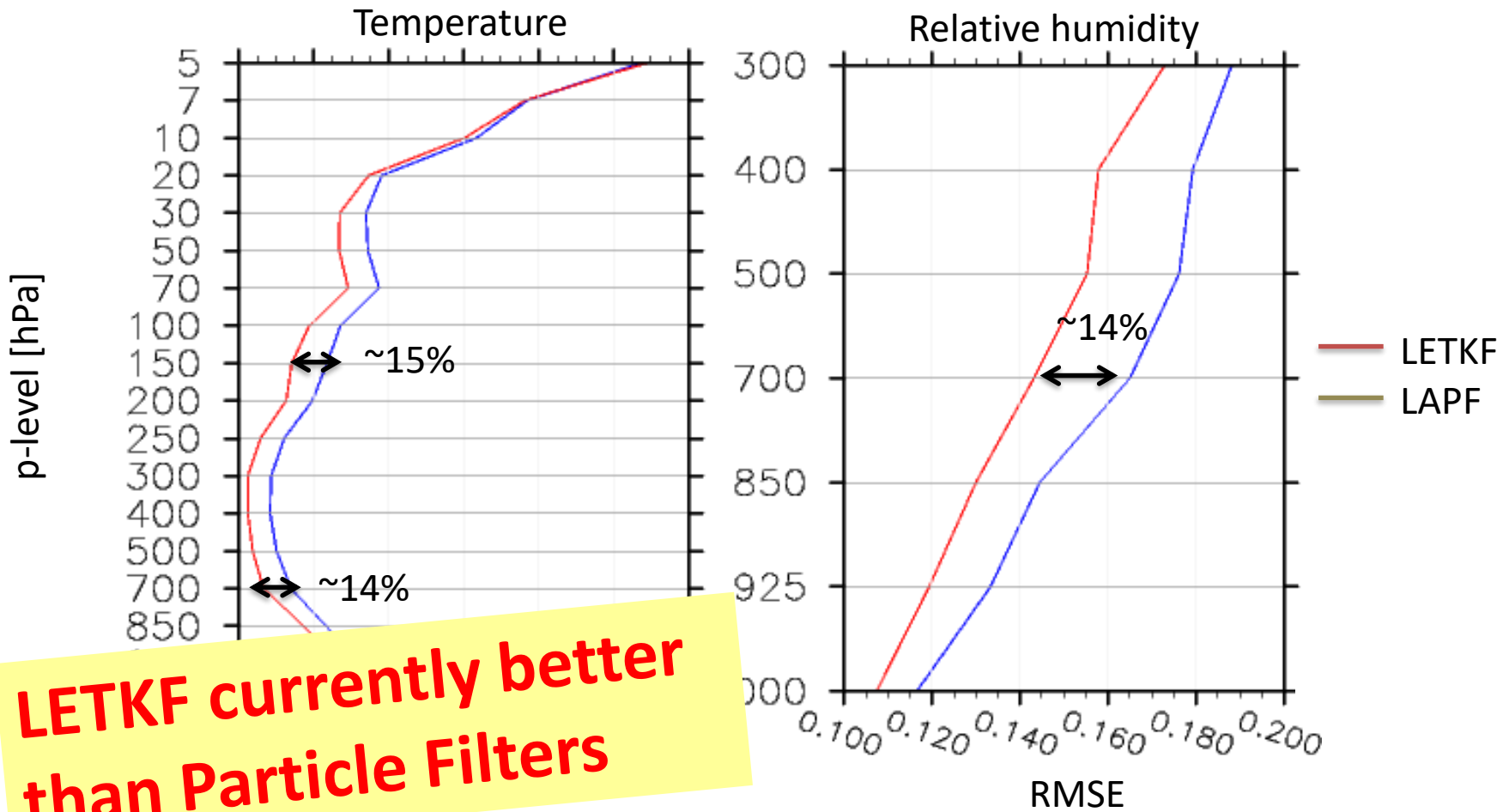


Experiments programmed and carried out  
by **Anne Walter, DWD& Uni Reading**, and  
Roland Potthast, DWD& Uni Reading

In Cooperation with Peter-Jan van  
Leeuwen , Uni Reading

# LAPF Scores vs LETKF

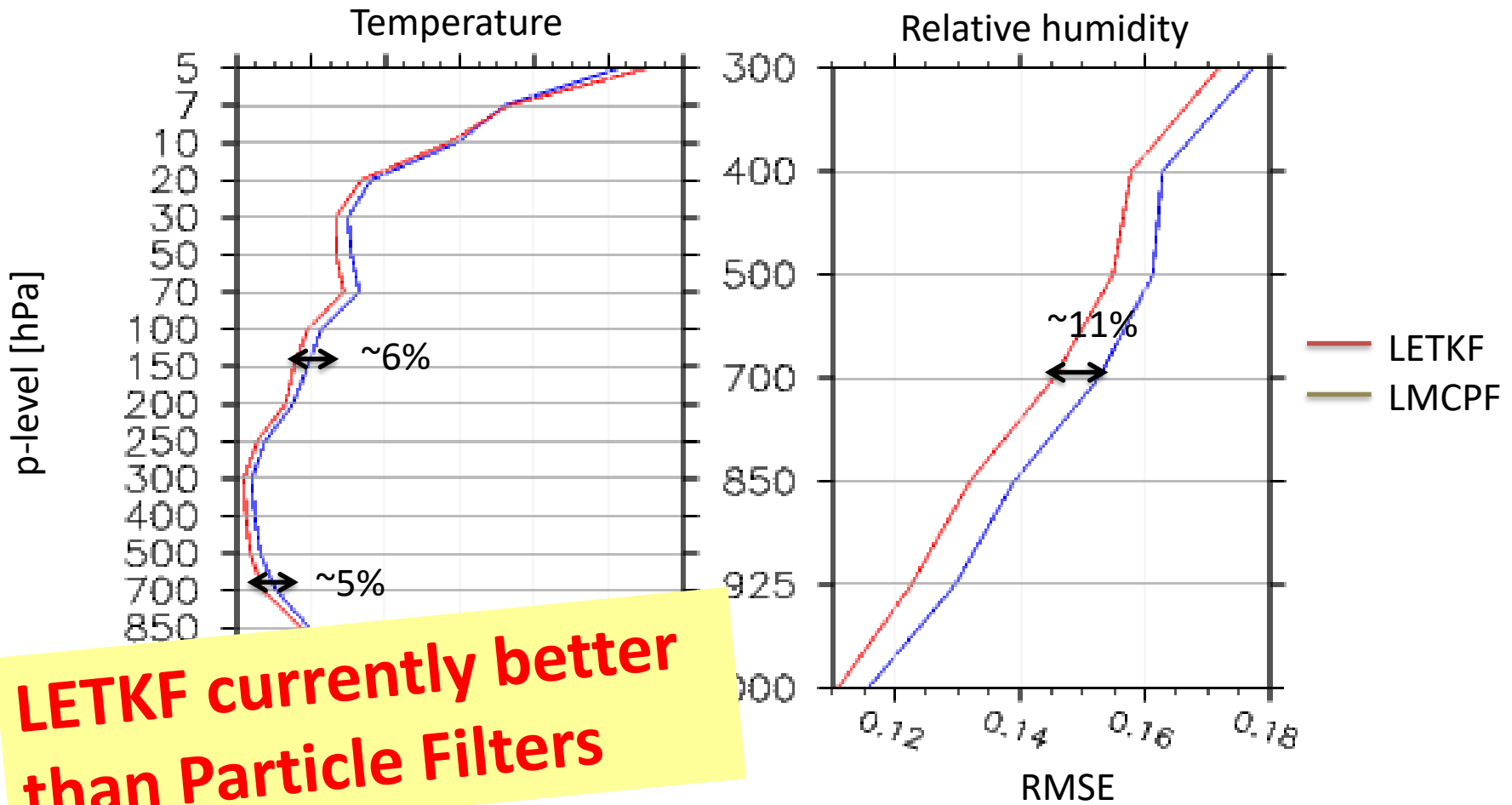
Global **RMSE** for **obs-fg** statistics (Radiosondes vs. Model)  
Period: 08.05.2016 – 31.05.2016



**LETKF currently better than Particle Filters**

# LMCPF Scores vs LETKF

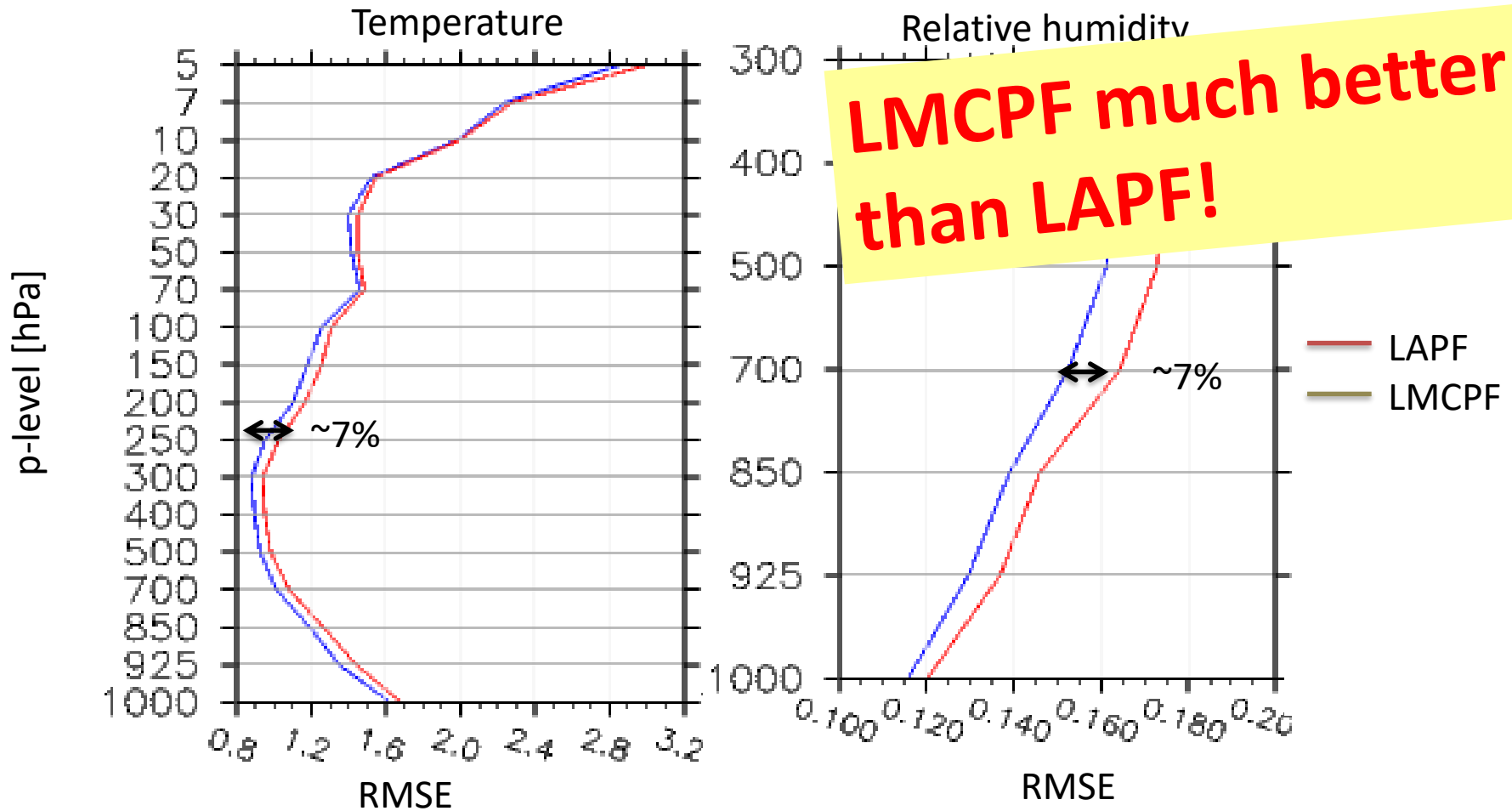
Global **RMSE** for **obs-fg** statistics (Radiosondes vs. Model)  
Period: 08.05.2016 – 22.05.2016



# LMCPF Scores vs LAPF

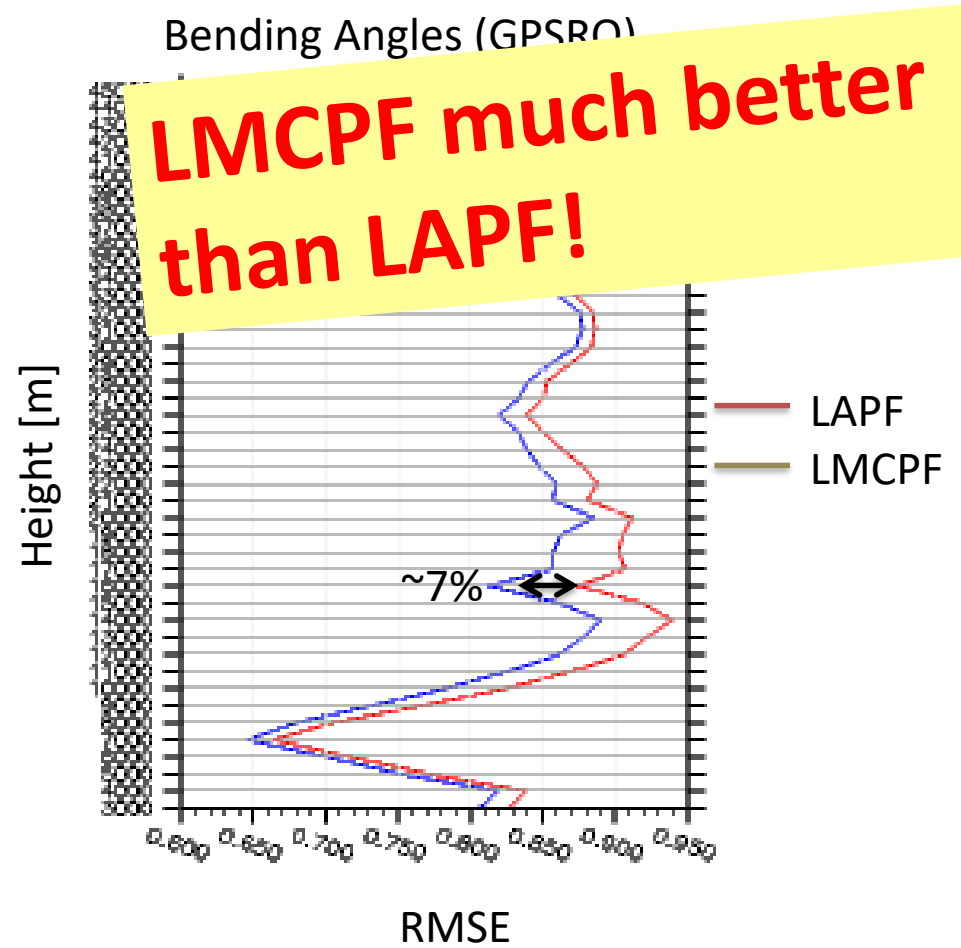
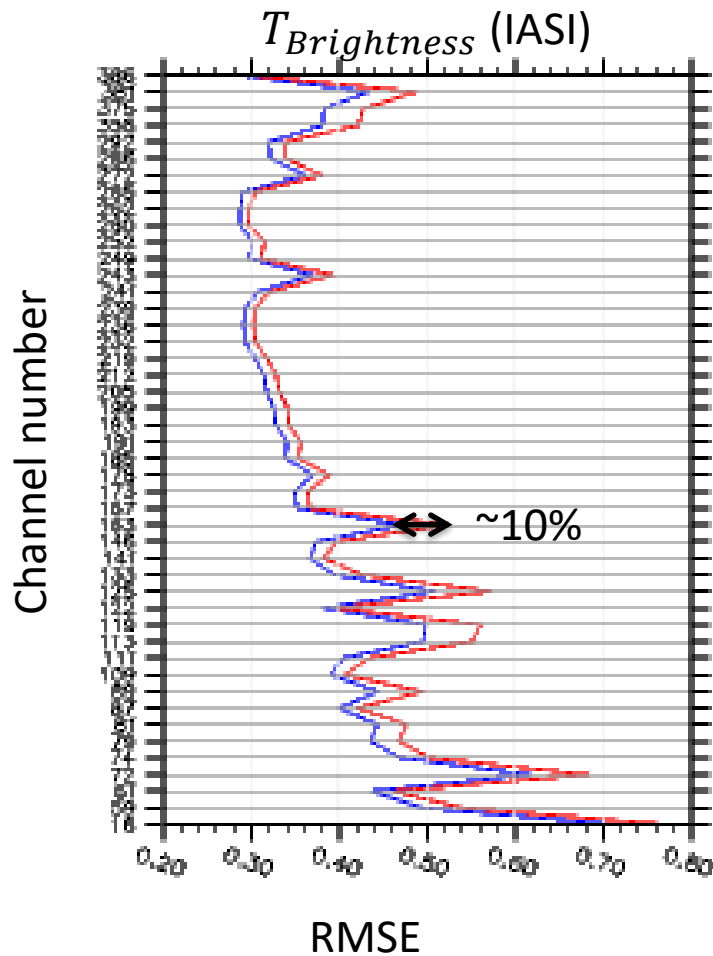


Global **RMSE** for **obs-fg** statistics (Radiosondes vs. Model)  
Period: 08.05.2016 – 22.05.2016



# LMCPF Scores vs LAPF

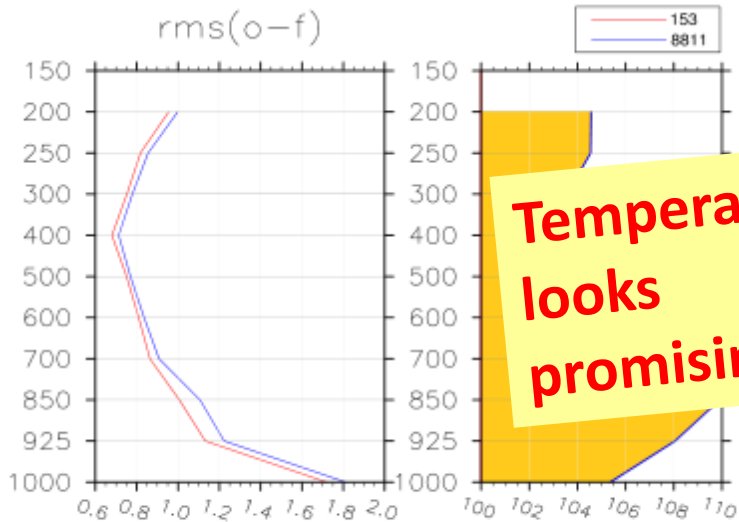
Global **RMSE** for **obs-fg** statistics  
Period: 08.05.2016 – 22.05.2016



# New LAMCPF Scores vs LETKF

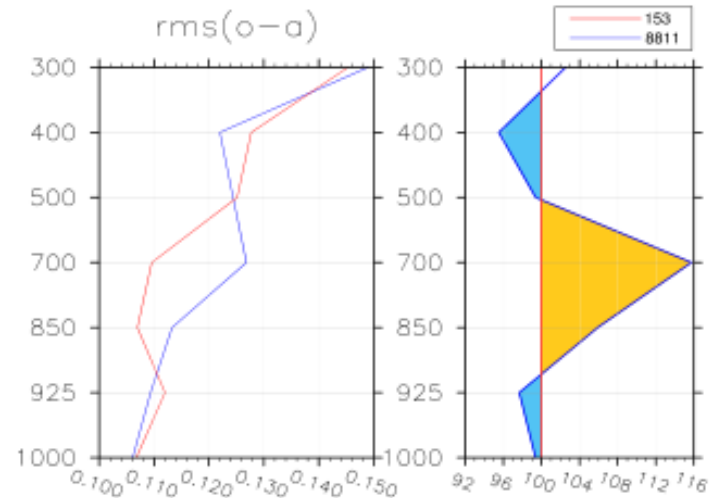
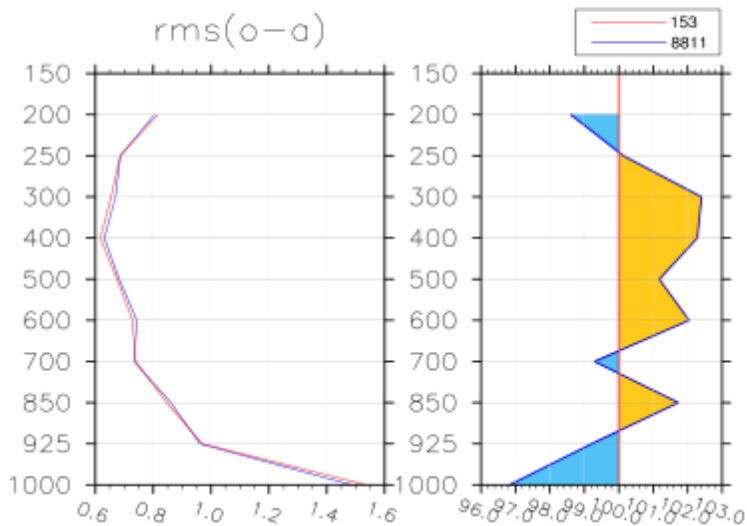
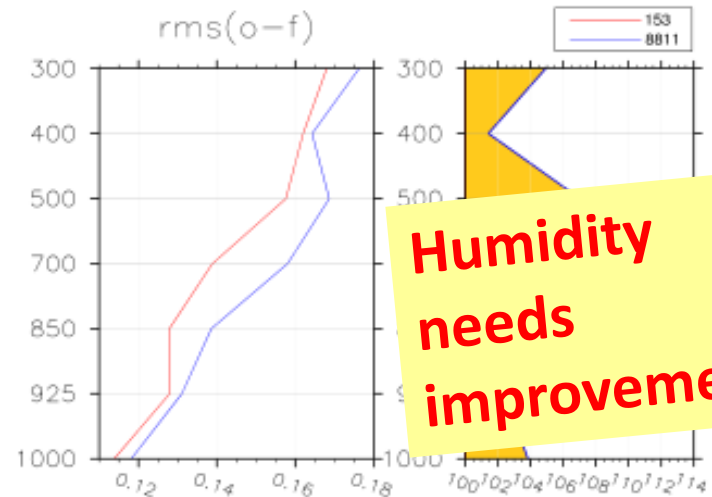
8811-153 AIREP T global other statistics dt=3h

relative differences 2016050700-2016050700



8811-153 TEMP RH global other statistics dt=3h

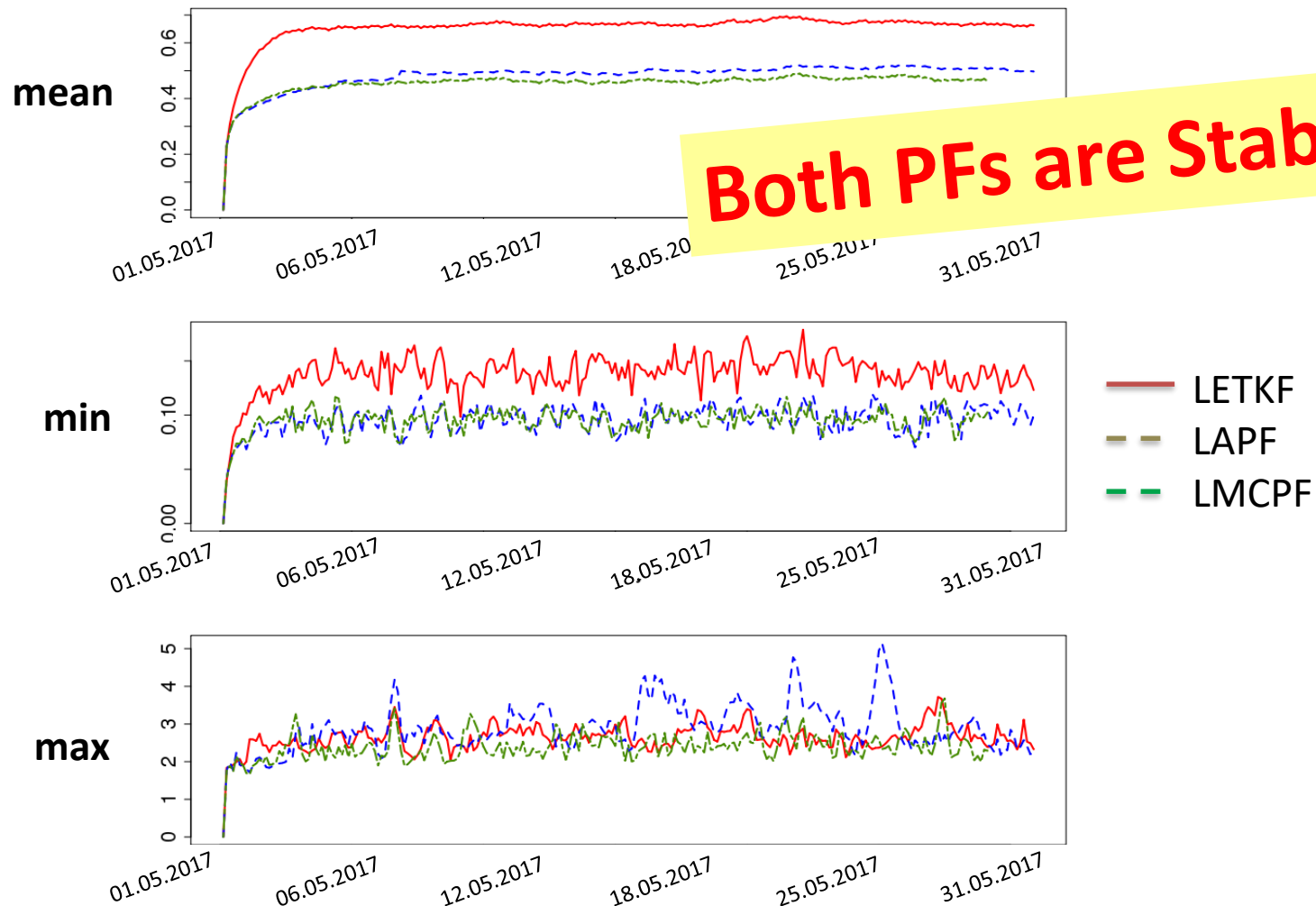
relative differences 2016050700-2016050700





# LAPF Spread vs LMCPF & LETKF

Global spread of T [K] ~ 500 hPa

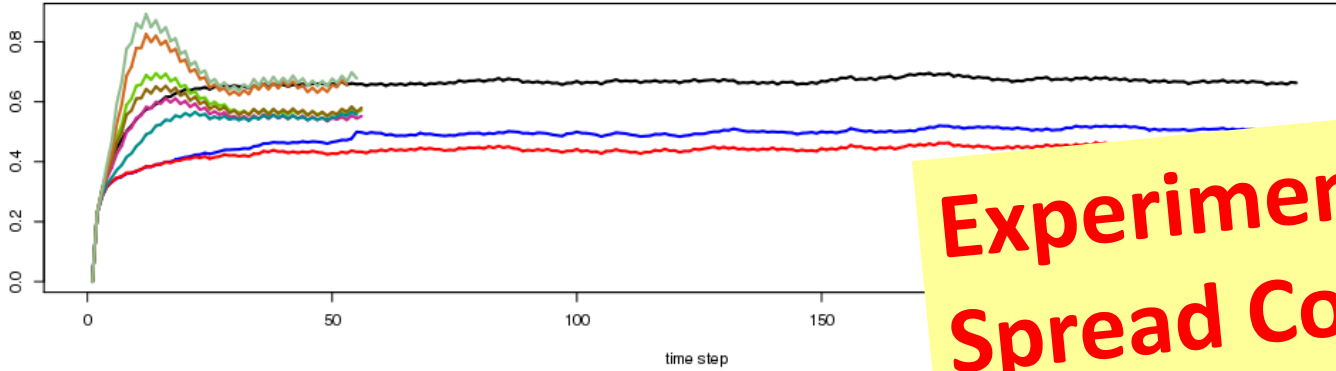


# Statistics for spread at level 64 for variable T



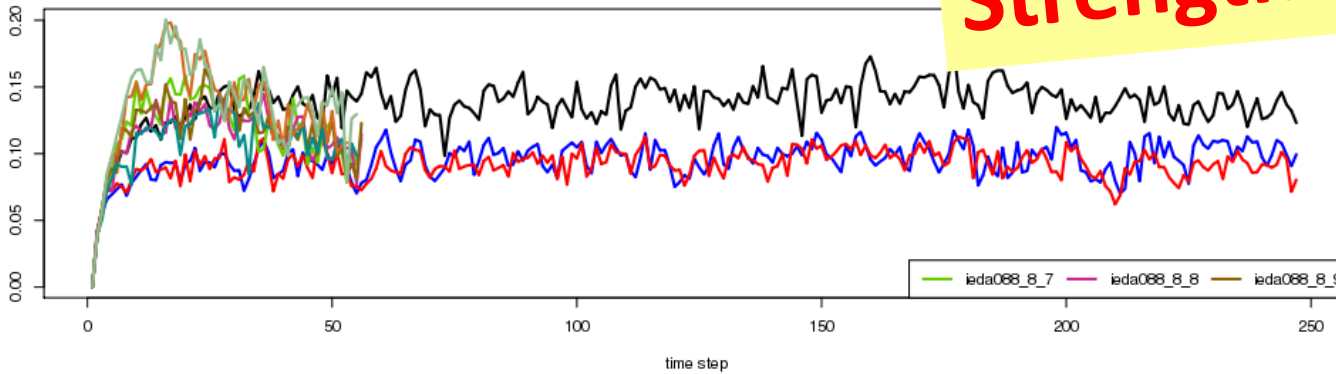
Deutscher Wetterdienst  
Wetter und Klima aus einer Hand

### Mean of spread

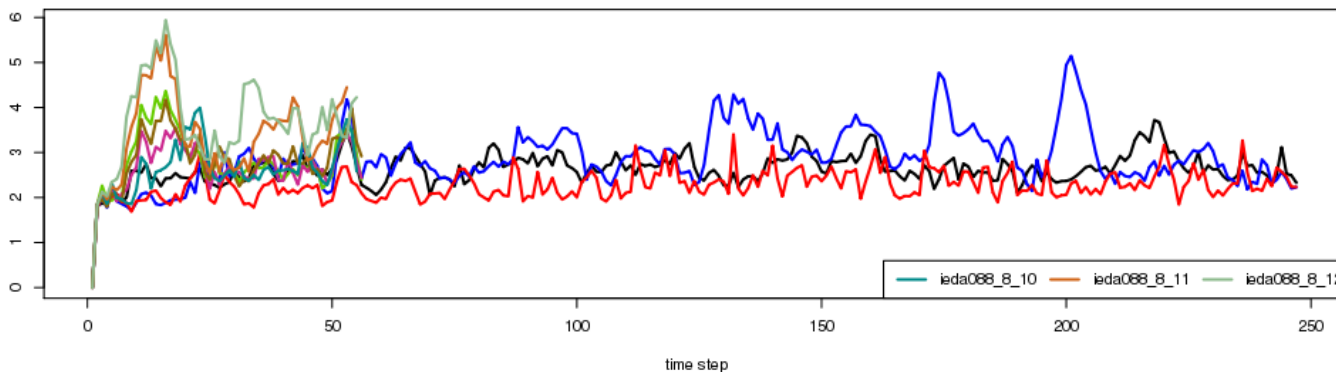


**Experiments with  
Spread Control and  
Strength of Shifts**

### Minimum of spread



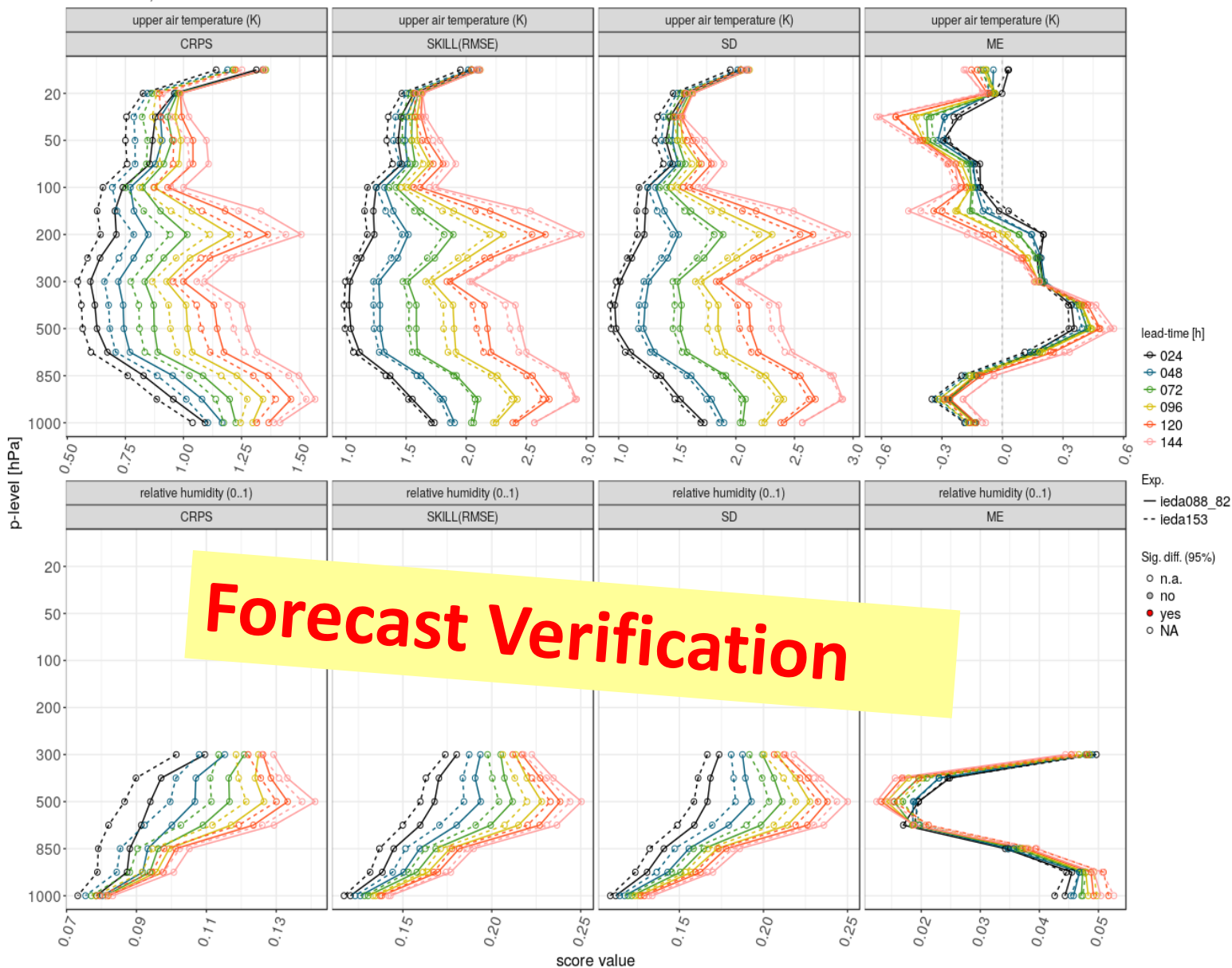
### Maximum of spread



# LMCPF Scores vs LETKF



2016/05/02 - 2016/05/24  
INI: ALL UTC, DOM: ALL

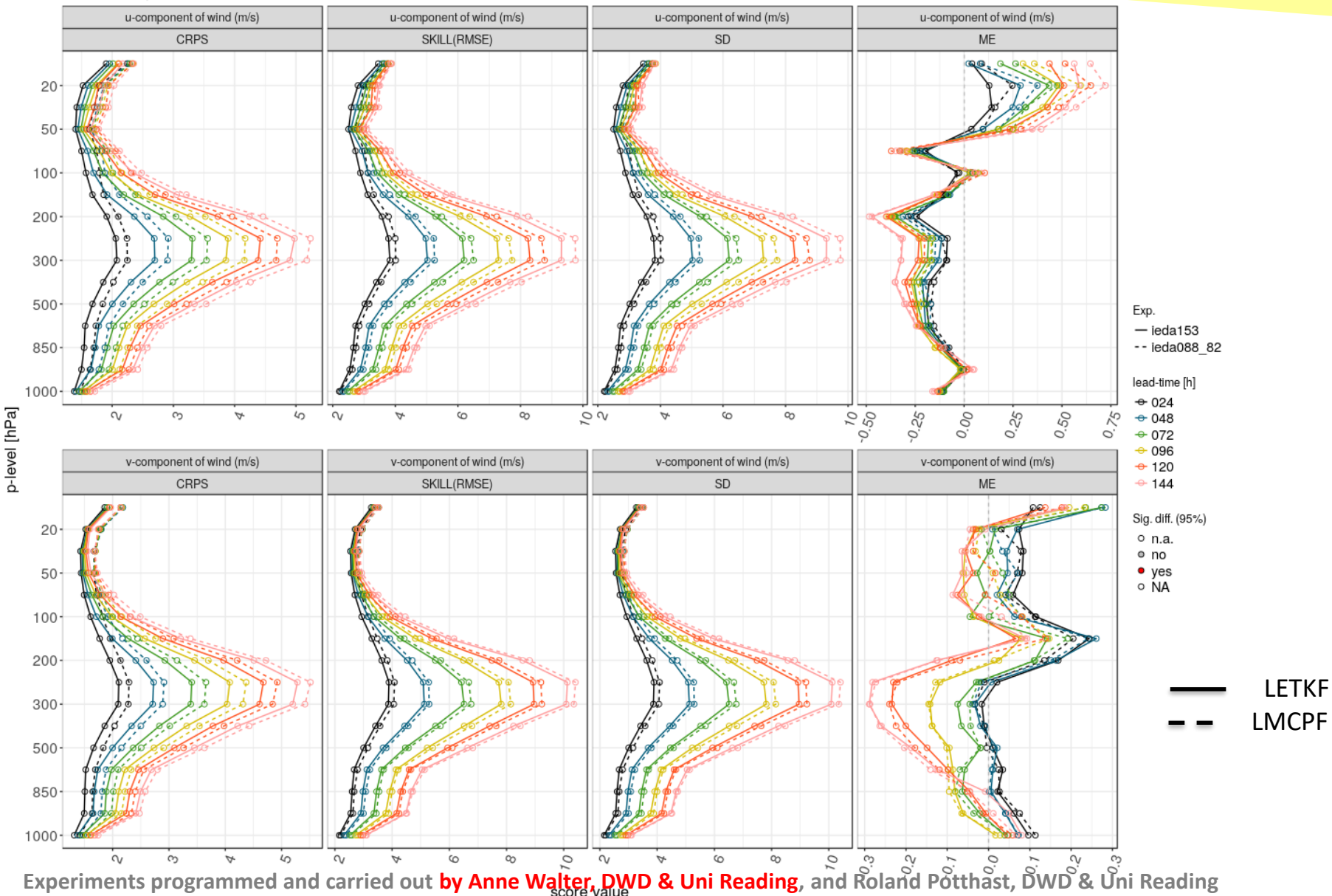


# LMCPF Scores vs LETKF



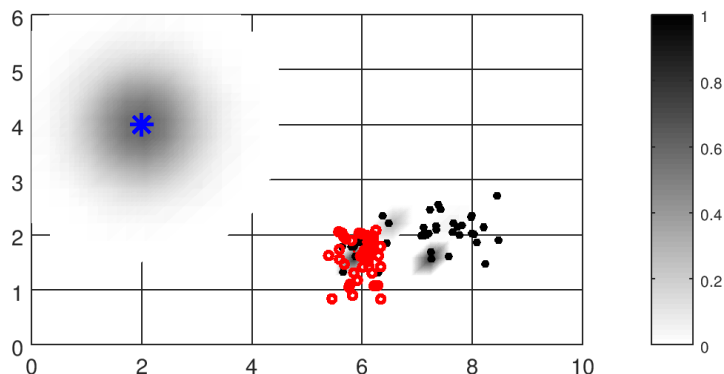
## Forecast Verification

2016/05/02 - 2016/05/24  
INI: ALL UTC, DOM: ALL

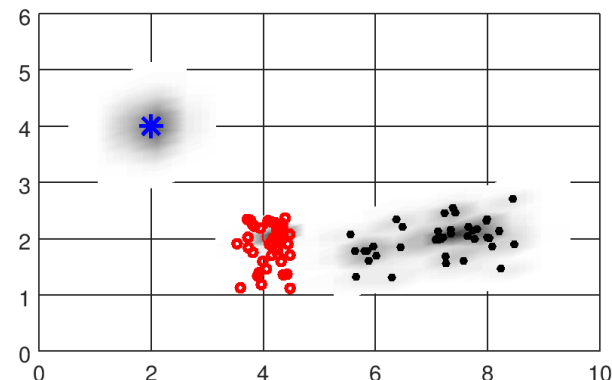


# Summary LAPF and LMCPF

## LAPF



## LMCPF



- LAPF and LMCPF are implemented in an **operational NWP system**: **Globally + mesoscale, convective scale**
- Both Particle Filters are able to provide **reasonable atmospheric analysis** in a large-scale (high-dimensional) environment and are running stably over a period of one month
- The LMCPF outperforms the LAPF but not yet the LETKF, but both Particle Filters are not far behind the operational LETKF

**Both Particle Filters are showing promising results; further tuning and development is in progress.**

# Many Thanks!



## Inverse Modeling

An introduction to the theory and methods of inverse problems and data assimilation

Gen Nakamura  
Roland Potthast

