



CENTRE NATIONAL D'ÉTUDES SPATIALES



The IASI Instrument

ECMWF / NWP-SAF Workshop
on the assimilation of IASI in NWP
Reading, May 6-8, 2009

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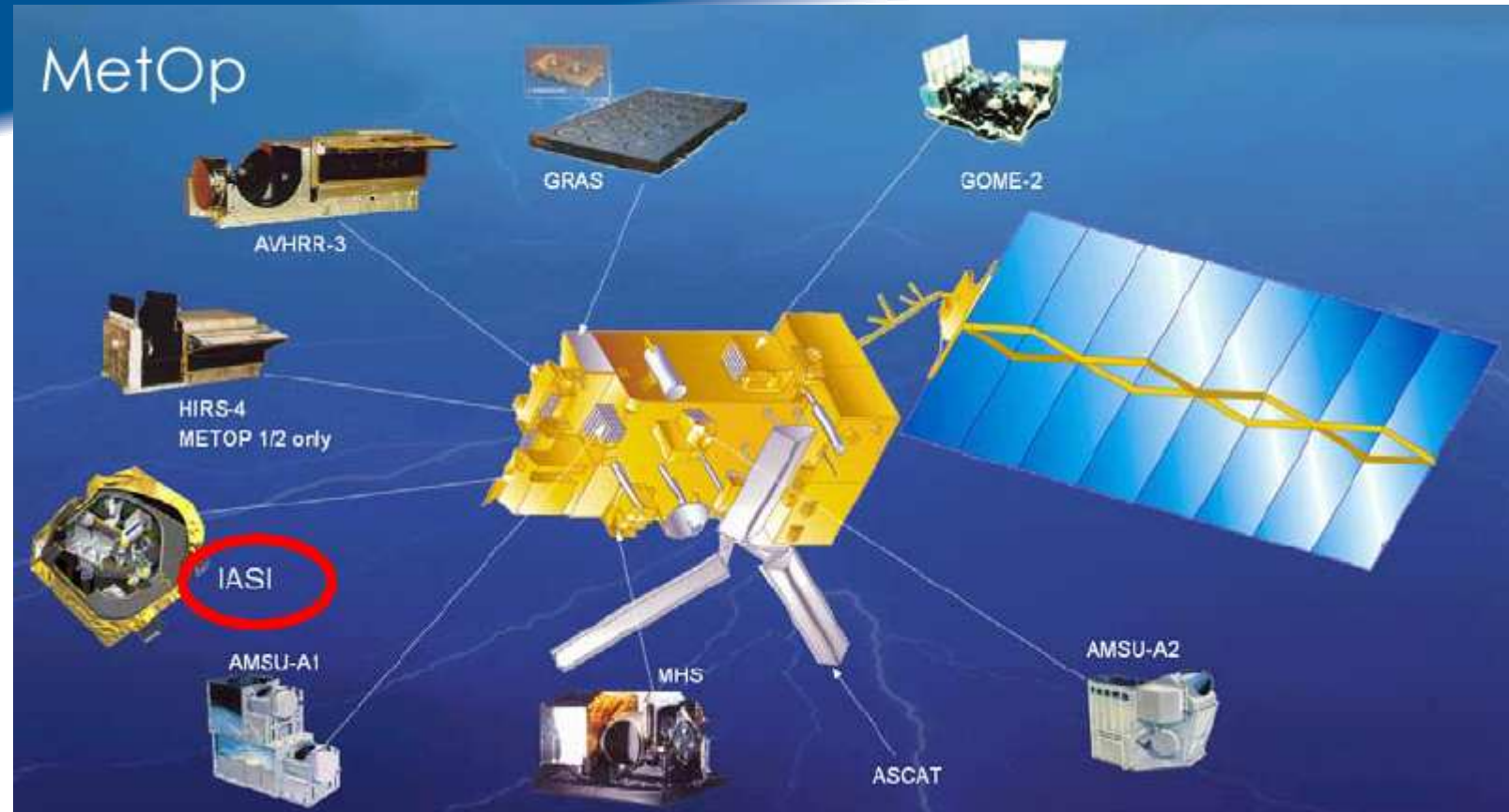
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(3) THALES ALENIA SPACE, Cannes, France



- 1. Introduction**
- 2. Functional Behaviour & Availability**
- 3. Radiometric noise & Ice contamination**
- 4. Radiometric calibration**
- 5. Spectral performances**
- 6. Geometry**
- 7. Processing and L0 / L1 data Quality**



- 3 instruments have been built → mission duration > 15 years
- MetOp-A launch : October 2006 → IASI declared operational : July 2007
 - ◆ IASI spectra assimilated by some NWP Center as early as June 2007

2.5 years of in-orbit experience



IASI : nadir Fourier Transform Interferometer

- ◆ For atmospheric sounding
- ◆ Cover without gap the thermal infrared region from 645 to 2760 cm⁻¹
- ◆ Maximum Optical Path Difference (OPD) : +/- 2 cm

◆ Spectral bands: 3.62 μm to 15.5μm

- B1 : 8.26 - 15.5 microns
- B2 : 5.0 - 8.26 microns
- B3 : 3.62 - 5.0 microns

◆ 4 off-axis pixels

◆ Field of view

- -48°20' / +48°20'

◆ Spatial resolution :

- Pixel diameter of 12Km

◆ Spectral resolution

- 0.5 cm⁻¹ (apodized spectra)

◆ Radiometric resolution :

- 0.2 to 0.4 K (apodized spectra)

◆ Data flow:

- 1.5 Mb/sec (average)

◆ Dimensions of sounder :

- 1.1 x 1.1 x 1.2 m³

◆ Mass sounder < 200 Kg

◆ Power consumption < 240 Watt (worst case EoL)

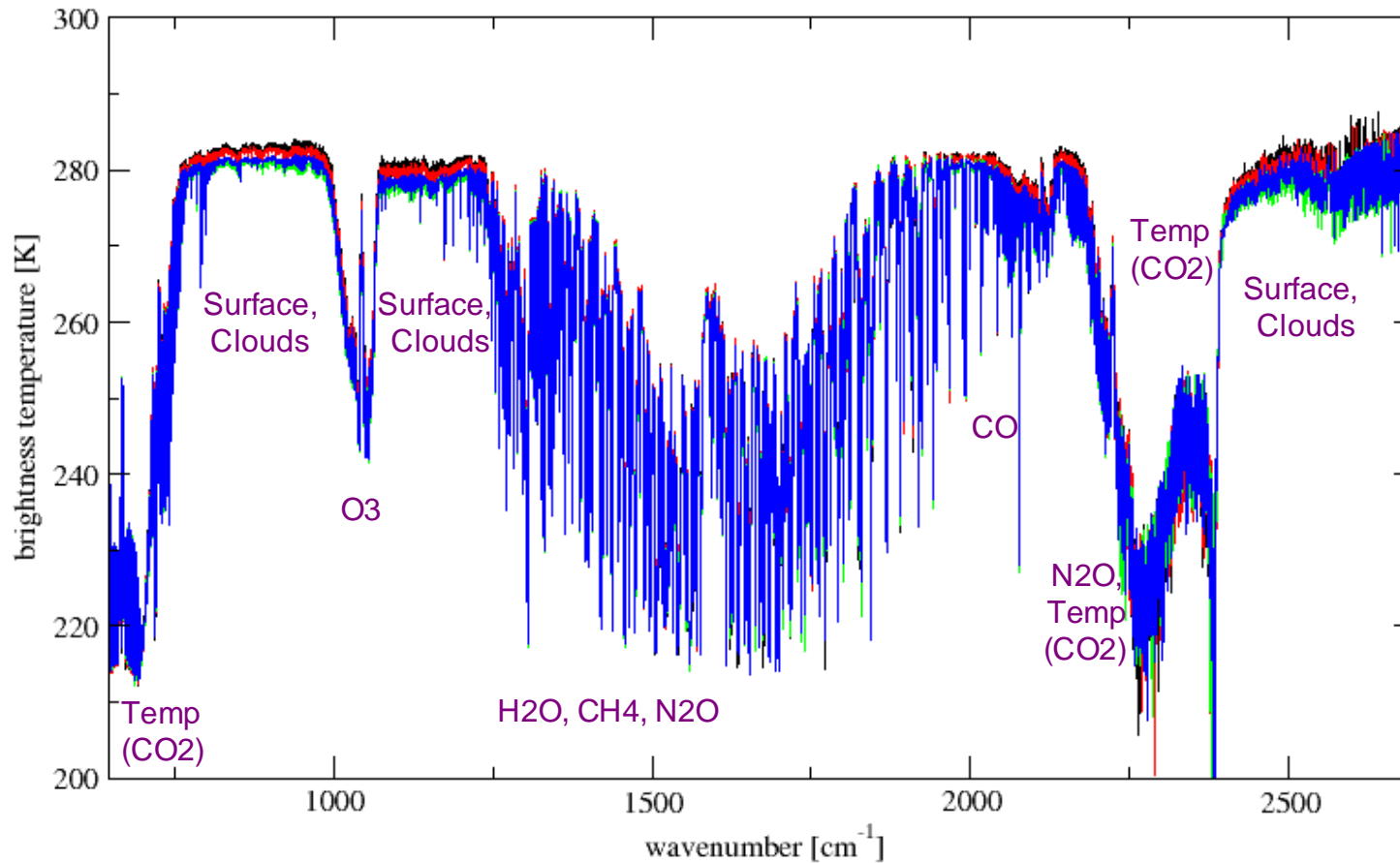
◆ Reliability > 0.8

◆ Availability > 95% over 5 years

+ Integrated Imager Subsystem 64 x 64 (0.8 km @ nadir), 10.3 μm – 12.5 μm

First IASI Level 1C Spectra

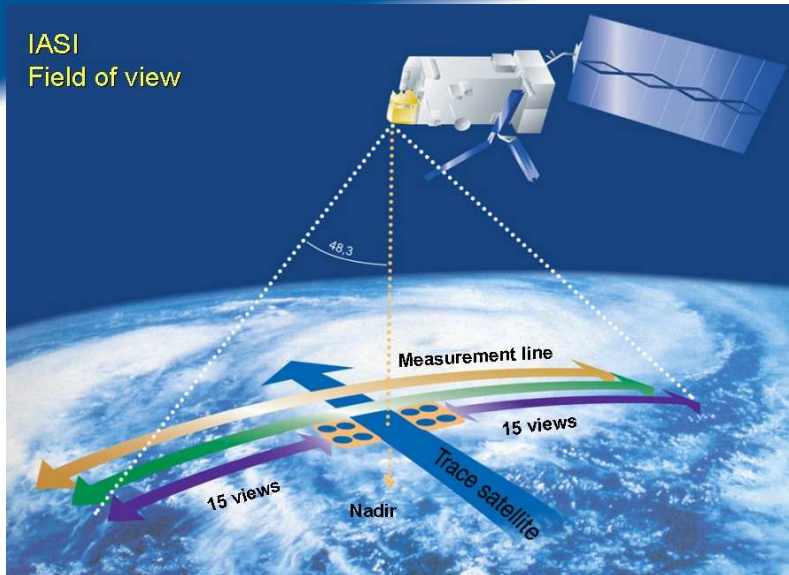
29/11/2006, 13:42:11 UTC



Generated by the IASI L1 PPF and Cal/Val Facility

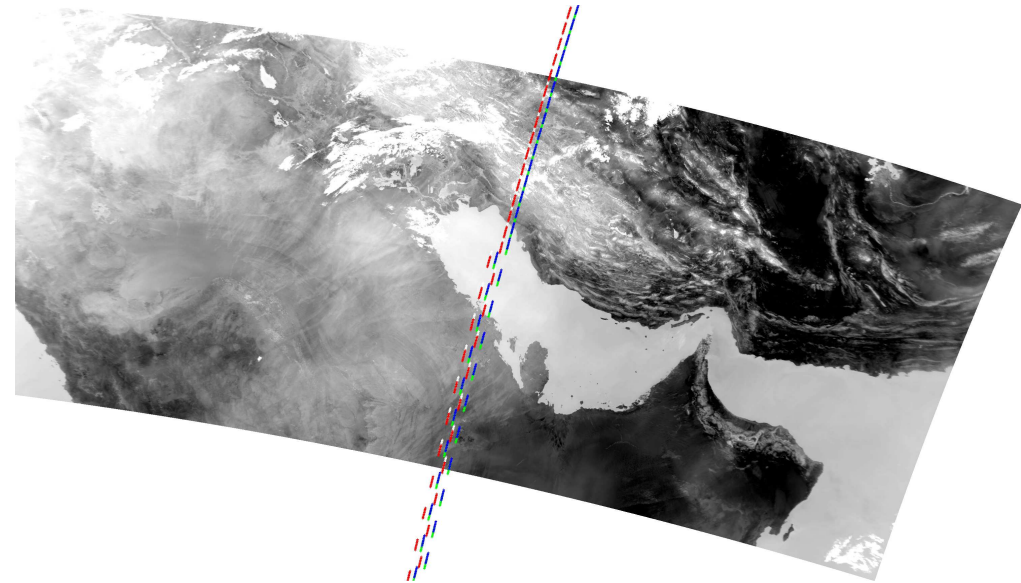


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Normal Operation Mode

- Scanning the swath
- 30 views / 8 sec



External Calibration Mode

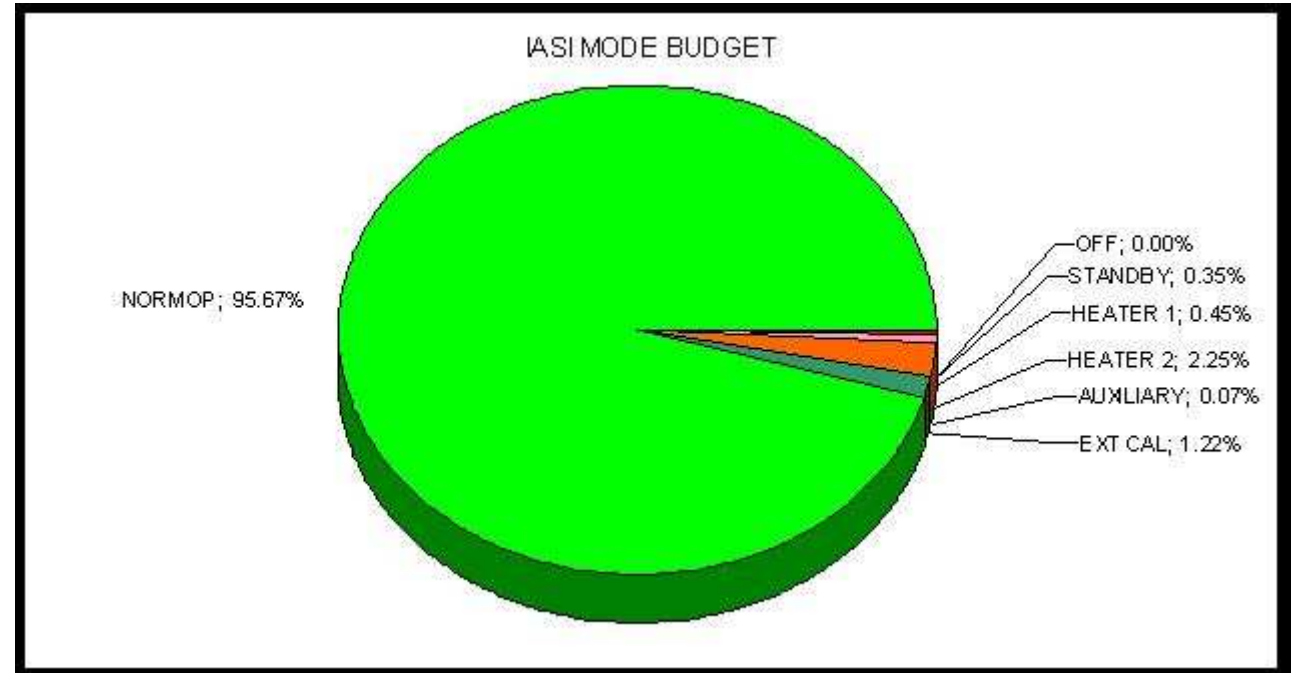
(here quasi-nadir looking)

- ◆ Fixed viewing direction for 8 sec
- ◆ 27 views / 8 sec

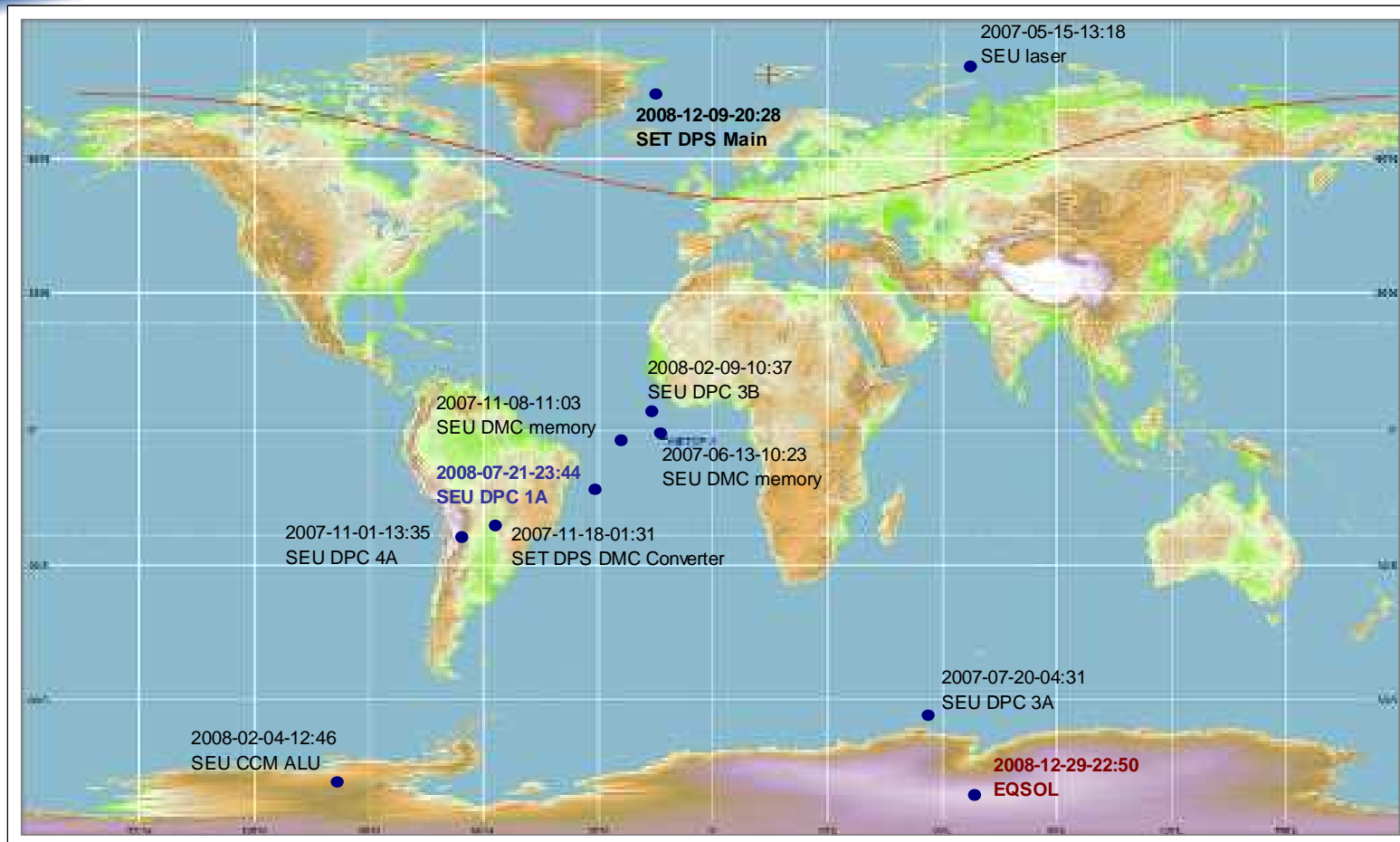
Pre-calibrated spectra computed on-board → science data TM
+ 1 raw interferogram available on ground every 8 seconds (over 408)
 ► selection fully programmable



- **Normal Op Mode**
 - ◆ 95.7 %
- **External Cal Mode**
 - ◆ 1.2 %
- **Instrument outage**
 - ◆ 3.1 %
 - Strong pressure from users to minimize outage duration



- **On-board software update designed to mitigate SEU affecting Data Processing Subsystem (most of the events)**
 - ◆ Automatic restart of suspended Data Processing
 - ◆ Will be uploaded before summer
 - ◆ Will not cover all the anomalies
 - In case focal plane temperature is lost, recovery still takes at least 2 days and 14 hours (passive cooling)



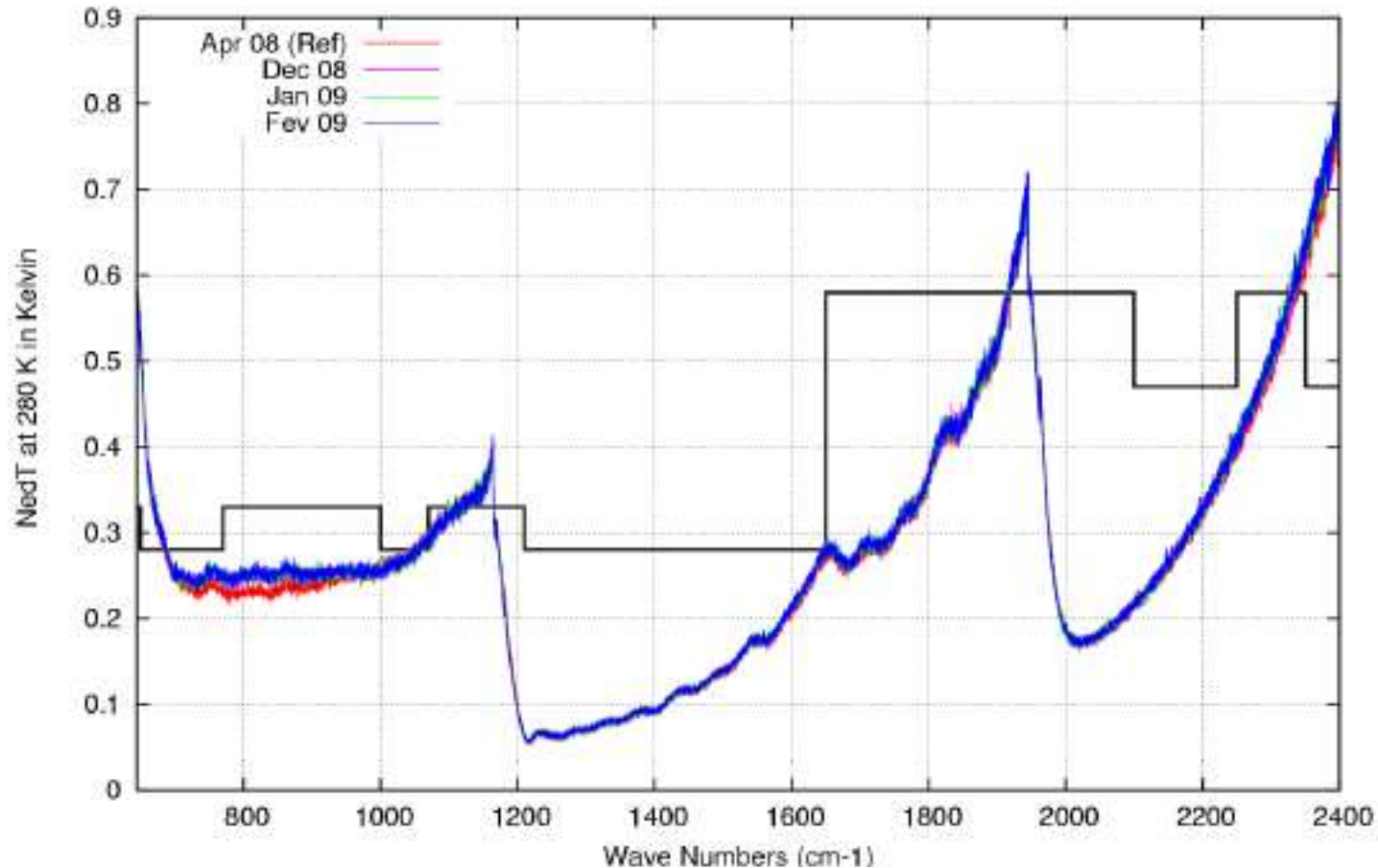
- Instrument outages caused by protons or heavy ions
 - ◆ Mainly over South Atlantic Anomaly & at High latitudes (North & South)



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Monthly Estimation (Ext.Cal.) by using Hot Black Body target

NeDT evolution between April 2008 and February 2009



Stable since last decontamination, except ice effect between 750 et 900 cm^{-1}

Estimation by using radiometric calibration coefficient (slope)

Physical phenomenon : water released by materials at 300K (MLI, electronics)

→ condensation on field lens at 100K (entrance of Cold Box Subsystem)

→ formation of ice

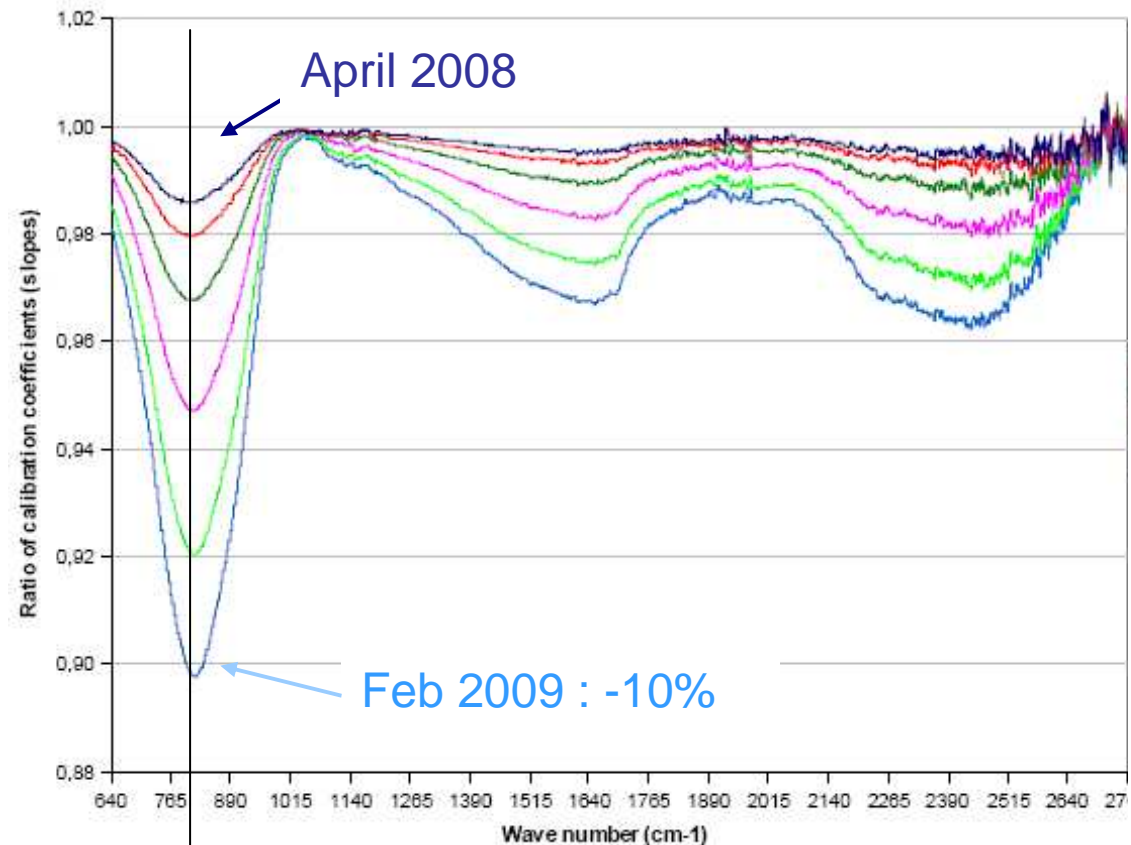
→ instrument transmission decreases

→ less signal

→ SNR decreases

→ NeDT increases

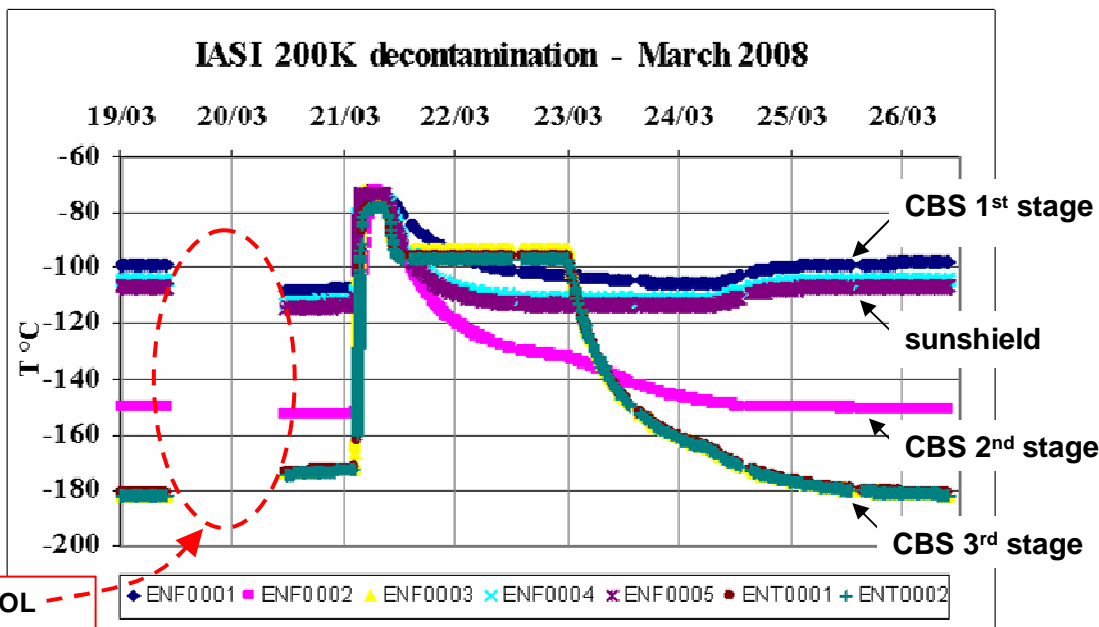
Instrument transmission evolution



Maximum loss due to water ice at 850 cm⁻¹

■ Decontamination

The decontamination lines heat the different parts of the Cold Box Subsystem (the three passive radiator stages and the sunshield) up to a temperature of 200 K (-73°C) for a duration of 4 hours. Then during the cooling down of the first and second stages, the third stage is maintained at -93°C in order to avoid re-deposition of ice on the cold optics. About 1.5 day later, when the second stage reaches -131°C, the third stage decontamination line can be switched off and the cooling of the first stage begins. It takes about 4 days to cool down the CBS third stage from -73°C to -181.8°C, the final temperature being exactly the same as before the outgassing phase.

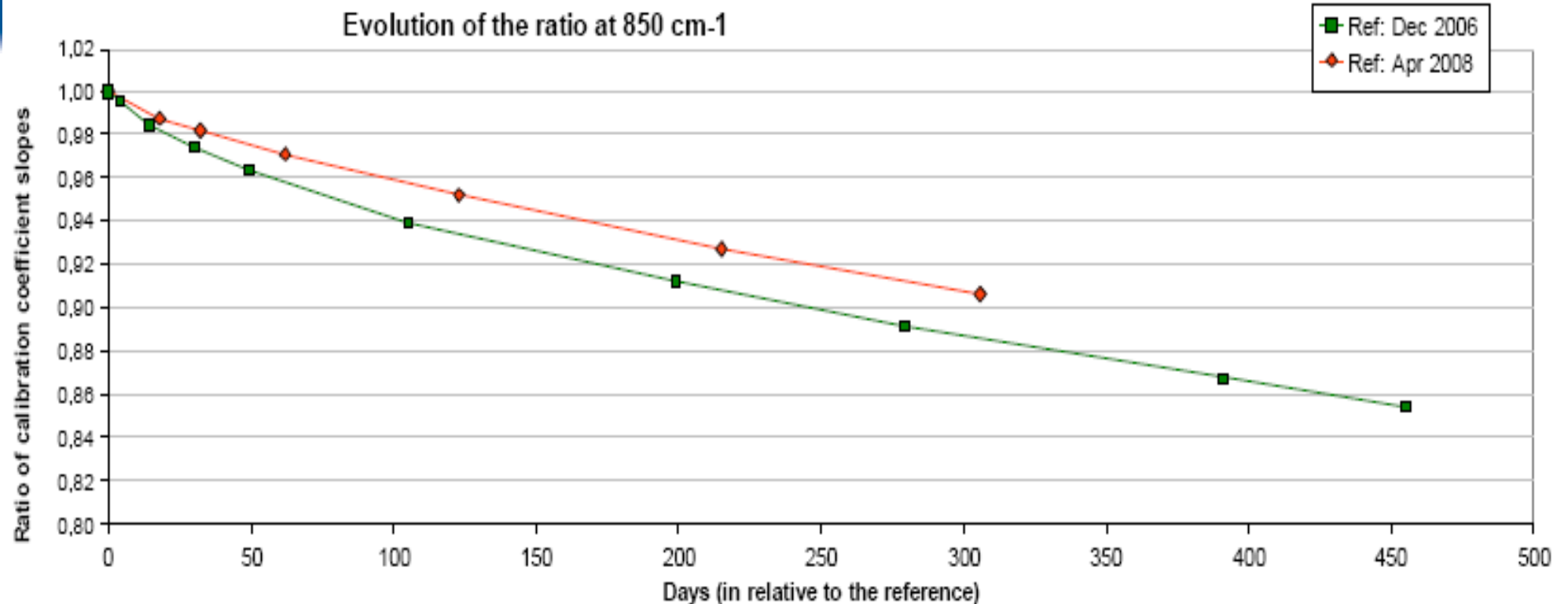


■ Performed during MetOp outage recovery

■ mission outage < 5 days.

■ After 1.5 year in-orbit

■ Recovery of the initial noise measured end of 2006



- **Criteria** : maximum noise increase of 20% (= transmission loss of 20%)
- **Last IASI decontamination** : 21-24th March 2008 (1.5 year after launch)
- **Next one** : Mid 2010

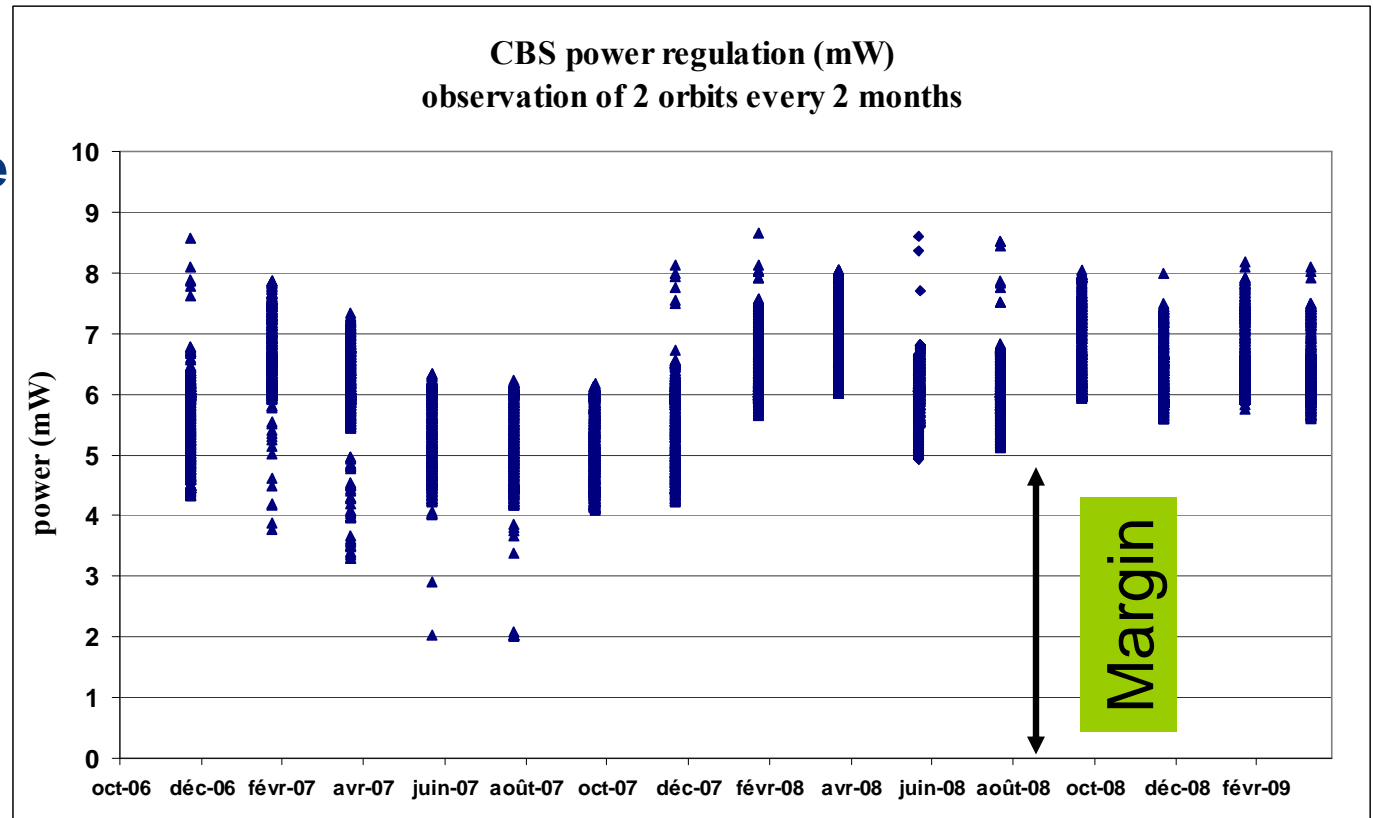
- Focal plane temperature regulated at 91.7 K

- Power regulation

- ◆ Stable (No trend towards 0)
- ◆ Seasonal effect

- Conclusions

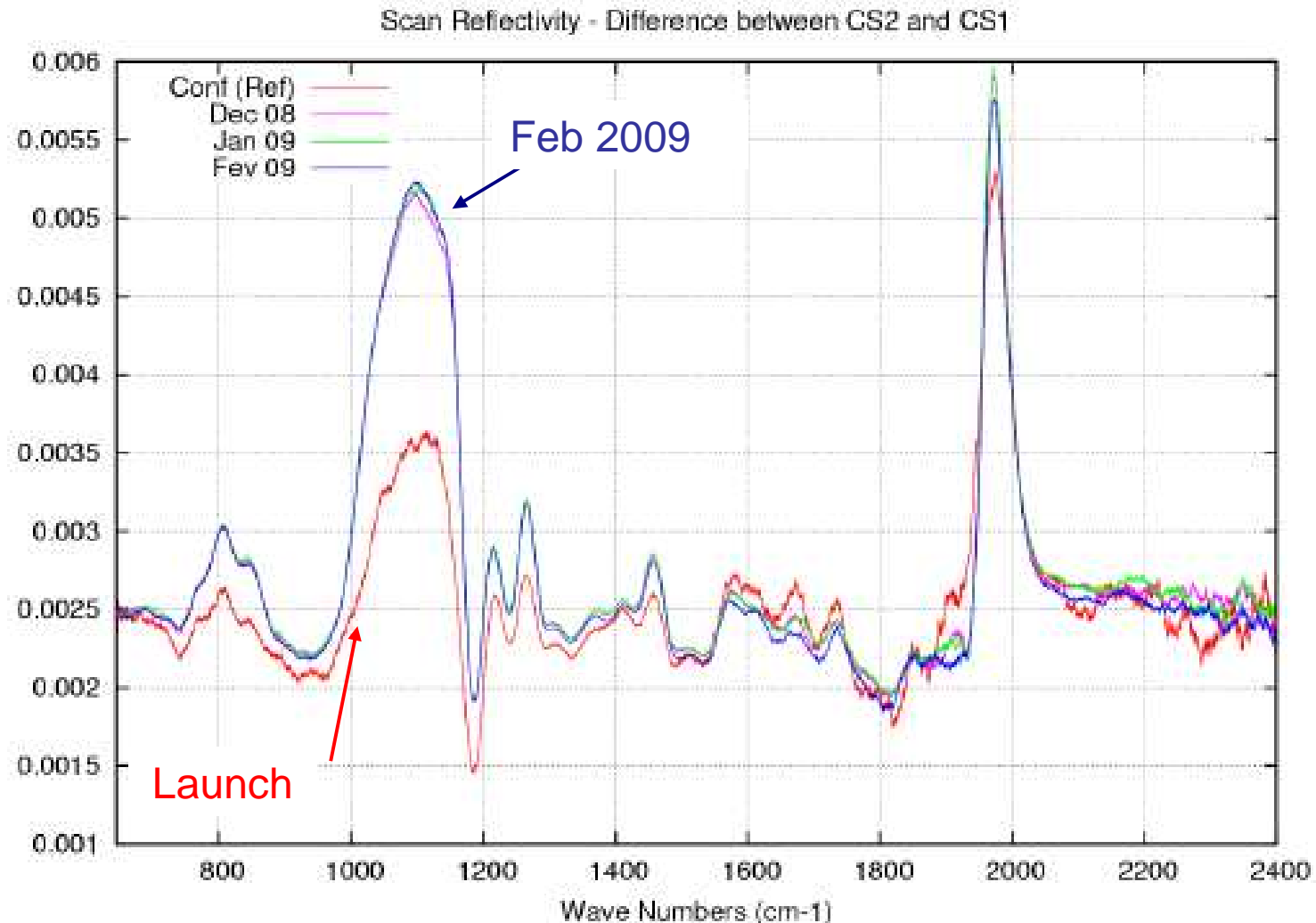
- ◆ Contamination of the sun shield is low
- ◆ Margin sufficient : No need to increase focal plane temperature target
- Stability of the radiometric noise expected in the next years



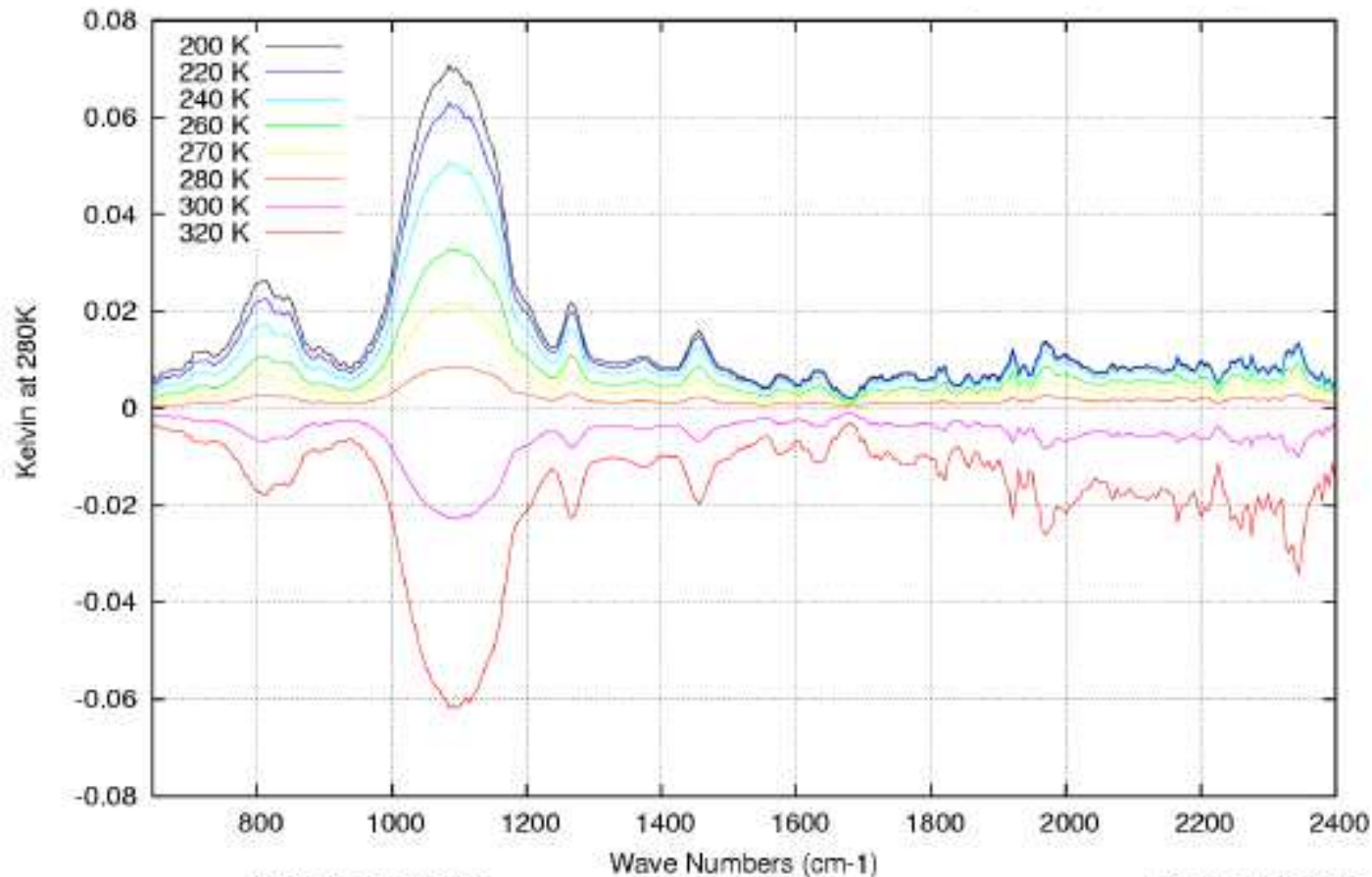


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Monthly Estimation (Ext.Cal.) by using spectra from CS1 (10°) and CS2 (60°) targets



Maximum impact of scan reflectivity variation on radiometric calibration within a scan line for different scene temperature



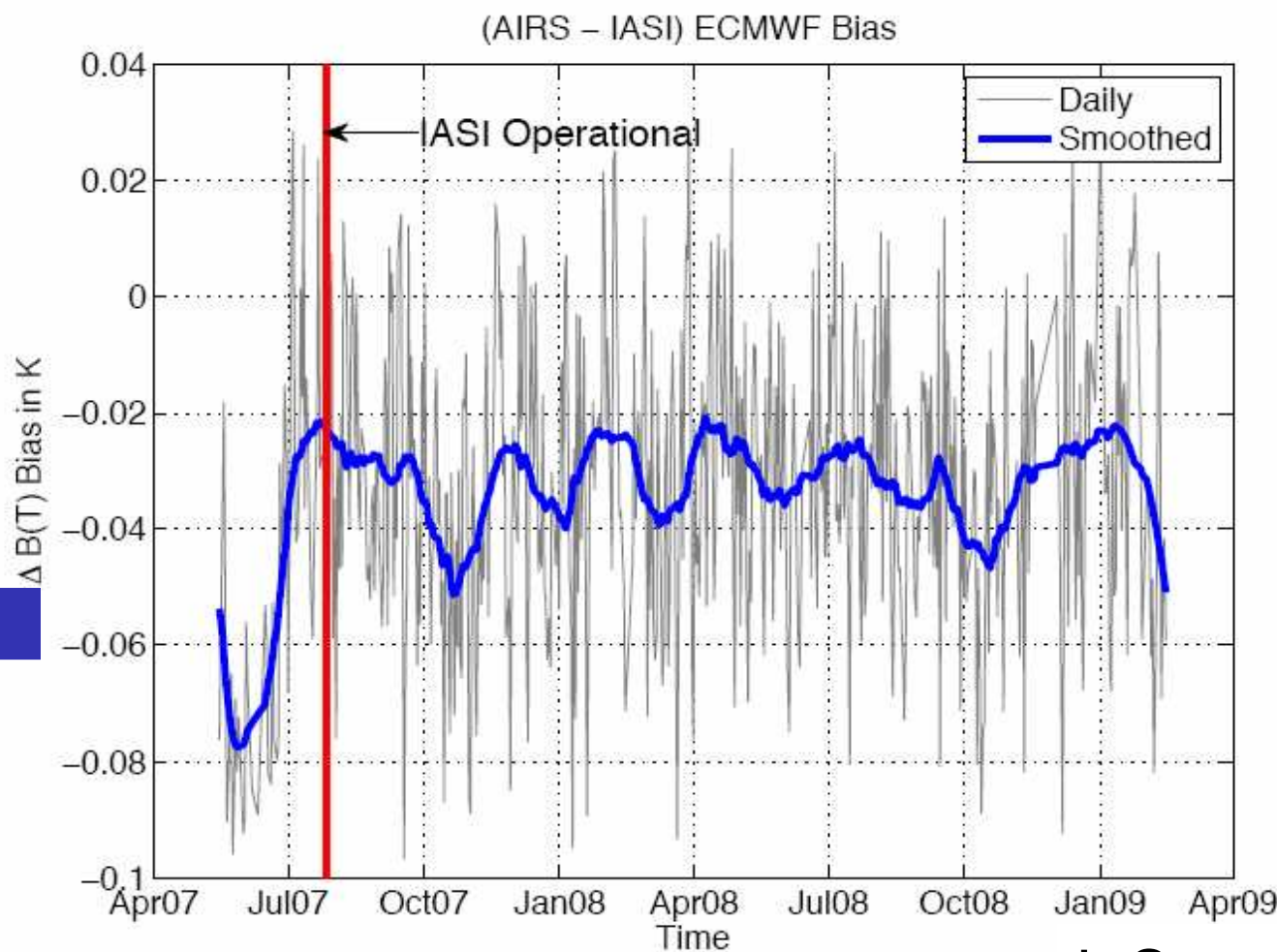
- specification = 0.1K
- Update of scan reflectivity in April 2009 (ground segment)
- Used in Level 1 Processing to correct for this effect (radiometric post-calibration)

AIRS versus IASI Stability

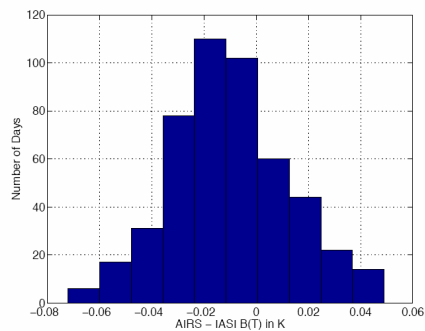
-0.0019K/year \pm 0.008K/year (corrected for lag-1 correlation of 0.45)

IASI/AIRS

L. Strow
UMBC



Histogram of Daily Observations



L. Strow (2009)



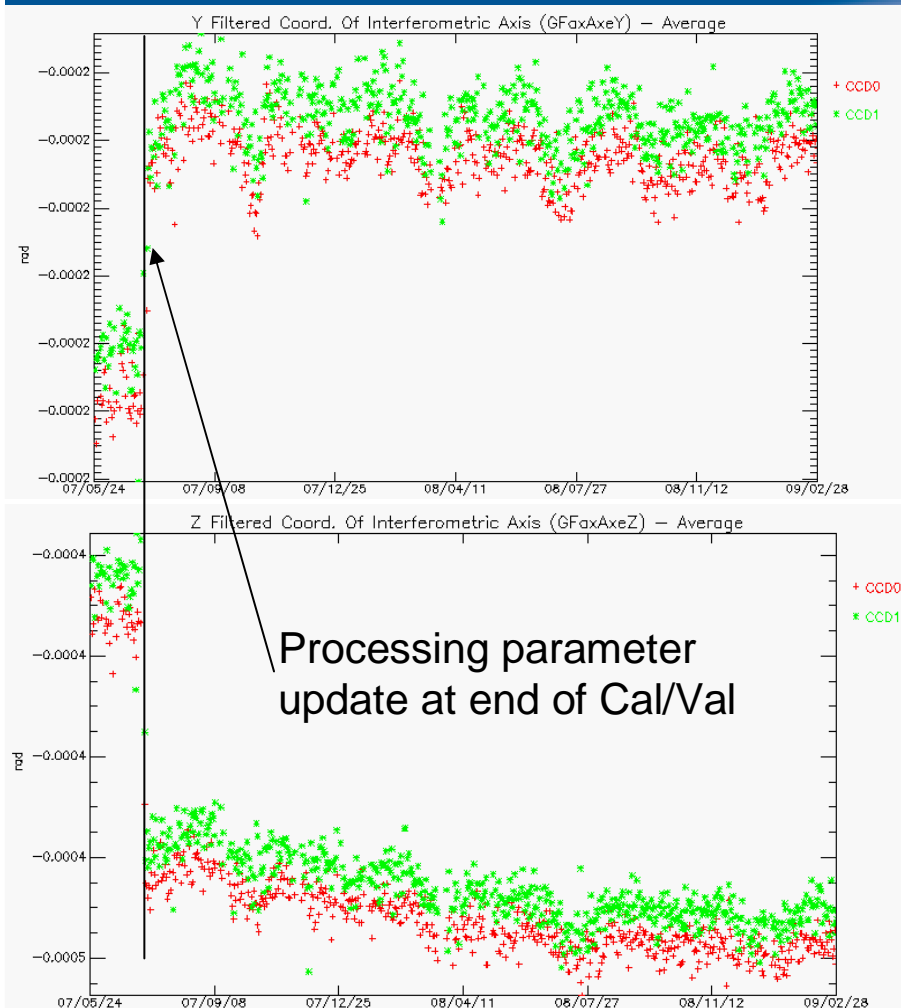
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Specification for IASI spectral calibration :

- A priori knowledge (instrument design) $\delta\sigma/\sigma < 2 \cdot 10^{-4}$
 - It means $\delta\sigma = 0.5 \text{ cm}^{-1} @ 2500 \text{ cm}^{-1} = \text{IASI spectral resolution}$ ($\sim 1/3$ of the spacing between two CO_2 absorption lines in $[2340 - 2380 \text{ cm}^{-1}]$ band used operationally for IASI spectral calibration)
- A posteriori knowledge (after on-ground spectral calibration) $\delta\sigma/\sigma < 2 \cdot 10^{-6}$
 - It means $\delta\sigma = 0.005 \text{ cm}^{-1} @ 2500 \text{ cm}^{-1} = 1\%$ of IASI spectral resolution)

For a good accuracy of IASI spectral calibration, we need a very good knowledge of Instrument Spectral Response Function (ISRF) => model



Y position of IA at the end of Cal/Val : $Y_0 = -159 \mu\text{rad}$

Long term drift: $Y - Y_0 = -8 \mu\text{rad}$

Seasonal cycle amplitude: $15 \mu\text{rad}$

Z position of IA at the end of Cal/Val : $Z_0 = -443 \mu\text{rad}$

Long term drift: $Z - Z_0 = -25 \mu\text{rad}$

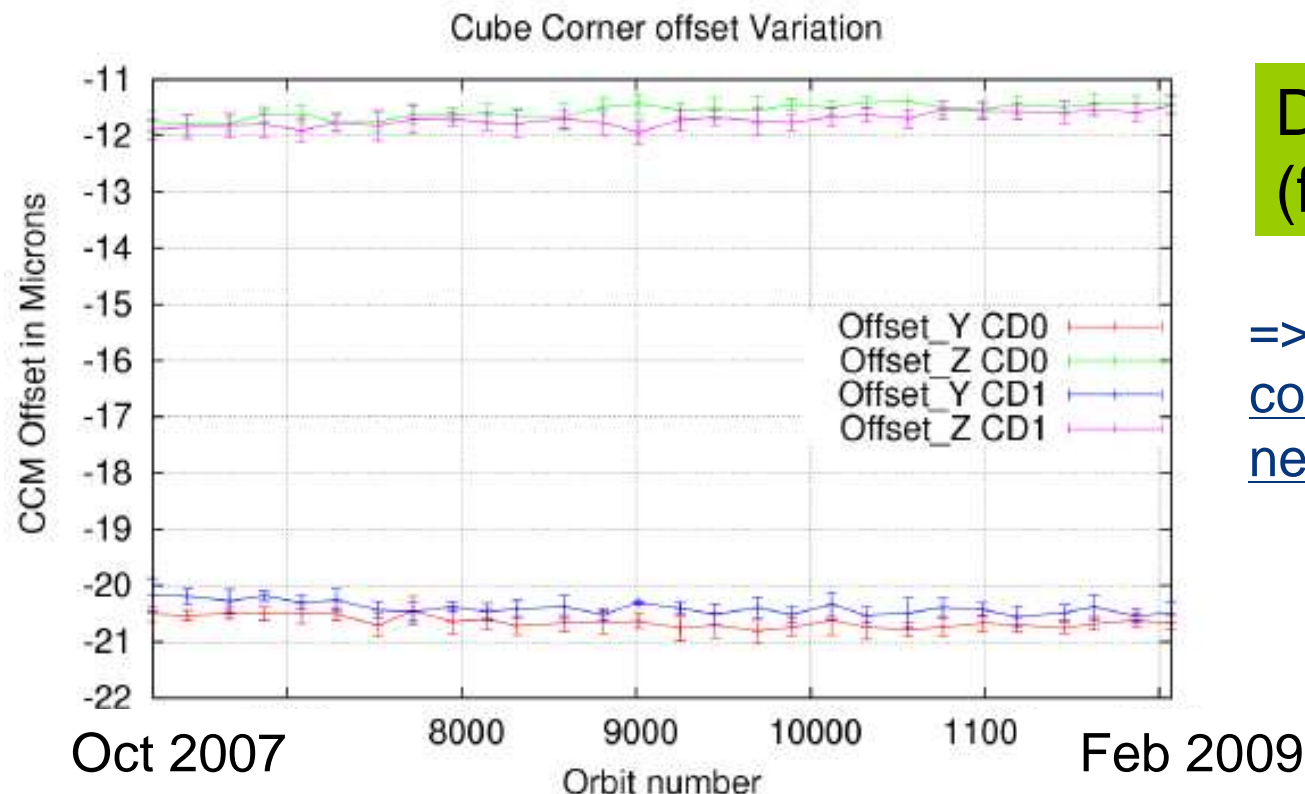
Seasonal cycle amplitude: $20 \mu\text{rad}$

- Total drift with respect to reference position in the spectral database: $(+40 \mu\text{rad}, -60 \mu\text{rad})$
- As soon as $|Total\ Drift| < 300 \mu\text{rad} \Rightarrow$ No spectral database configuration update needed



Velocity : continuous on-board monitoring
+ regular in-depth checks (no evolution)

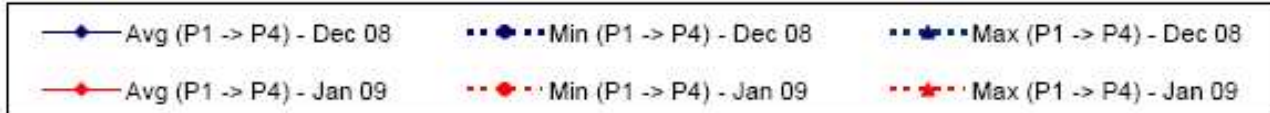
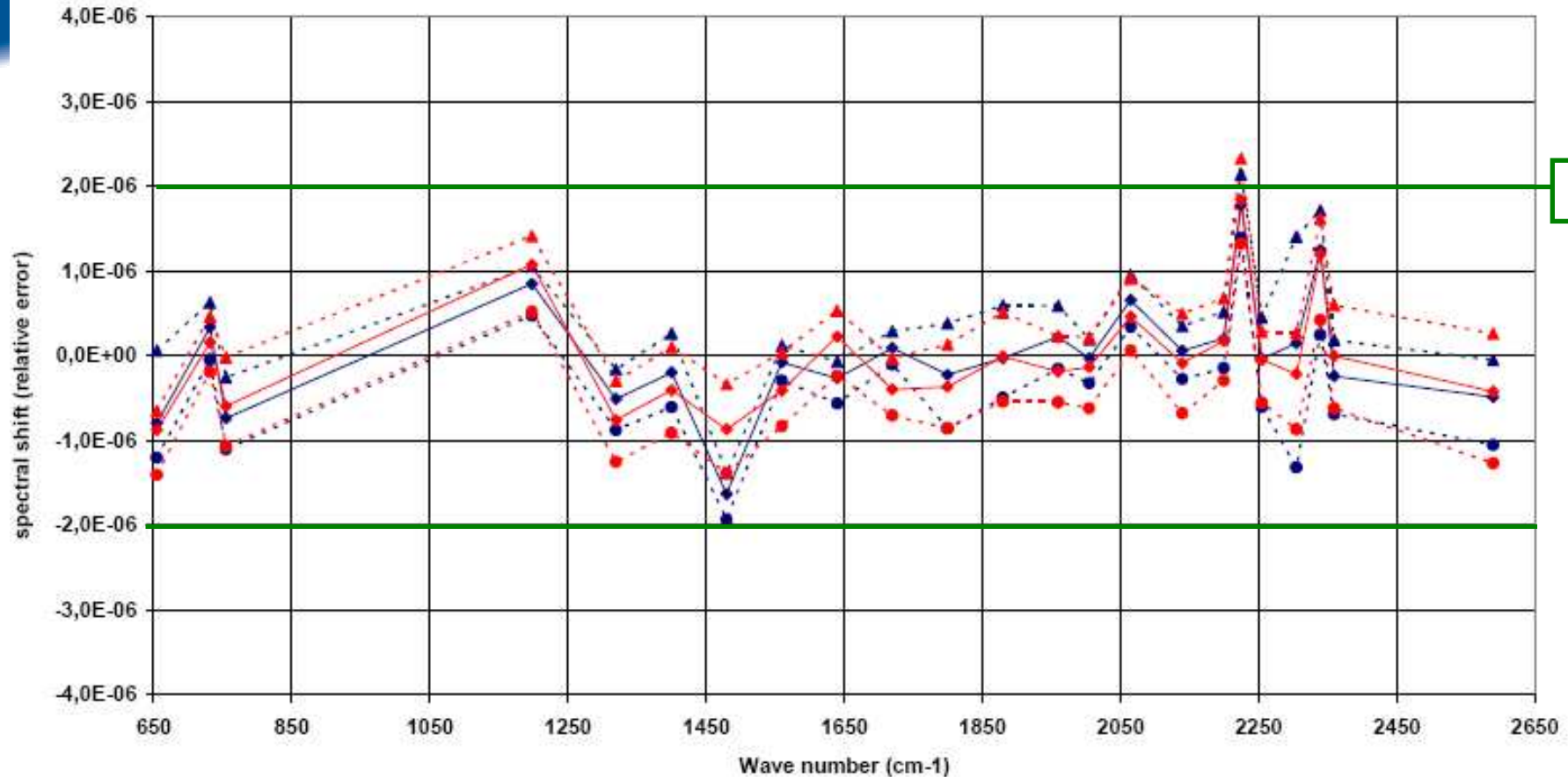
Position : cube corner offset (shear)



Drift < 1 μm
(for 2.5 years)

=> No spectral database
configuration update
needed (up to 4 μm)

Spectral calibration drift (Ref : end of Cal/Val - July 2007)

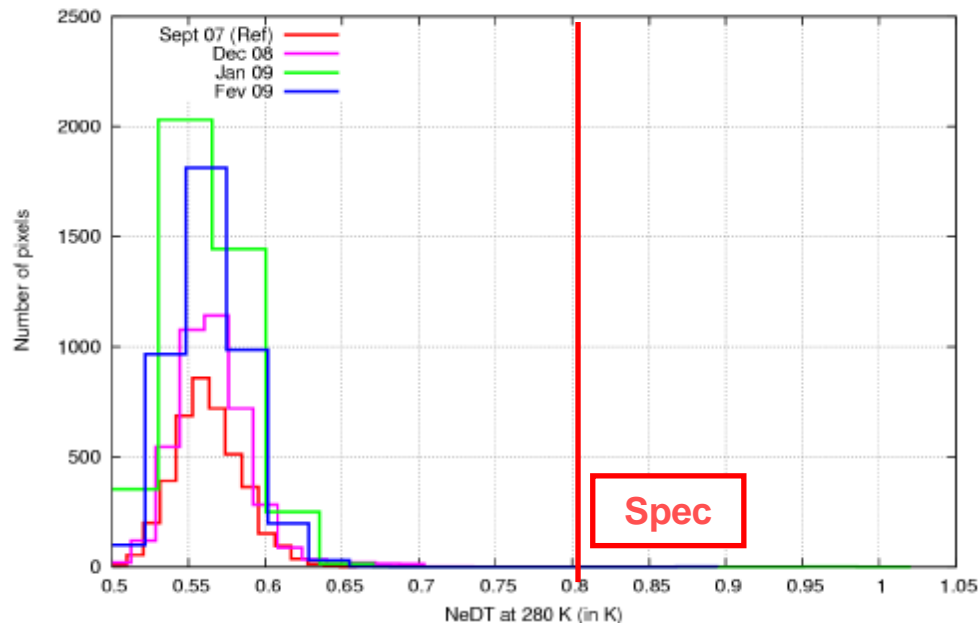




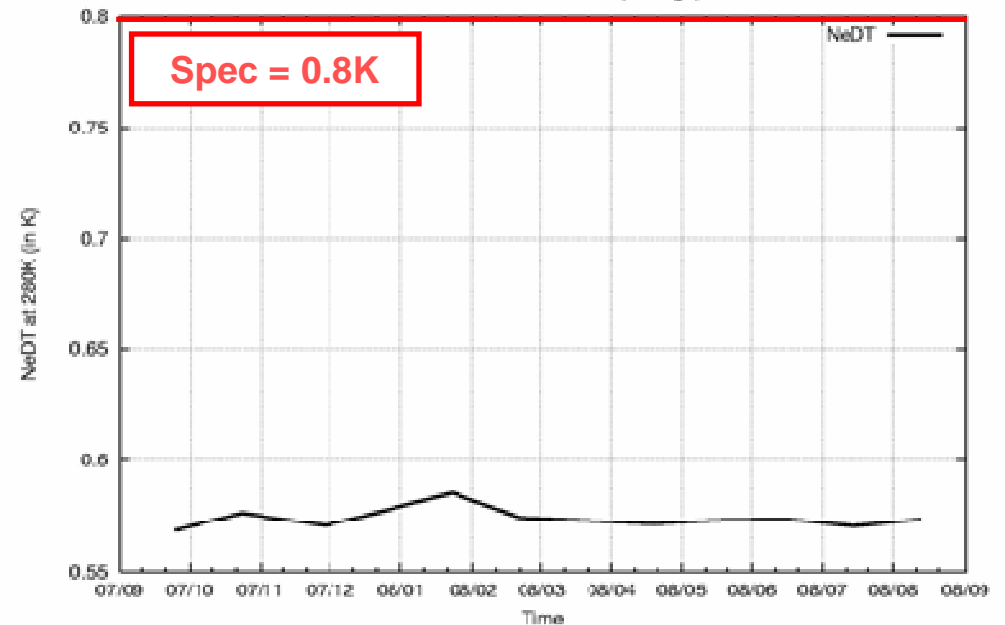
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Geometry of IASI sounder controlled with respect to the companion imager (IIS)

Temporal evolution of IIS noise (histogram with all the pixels)

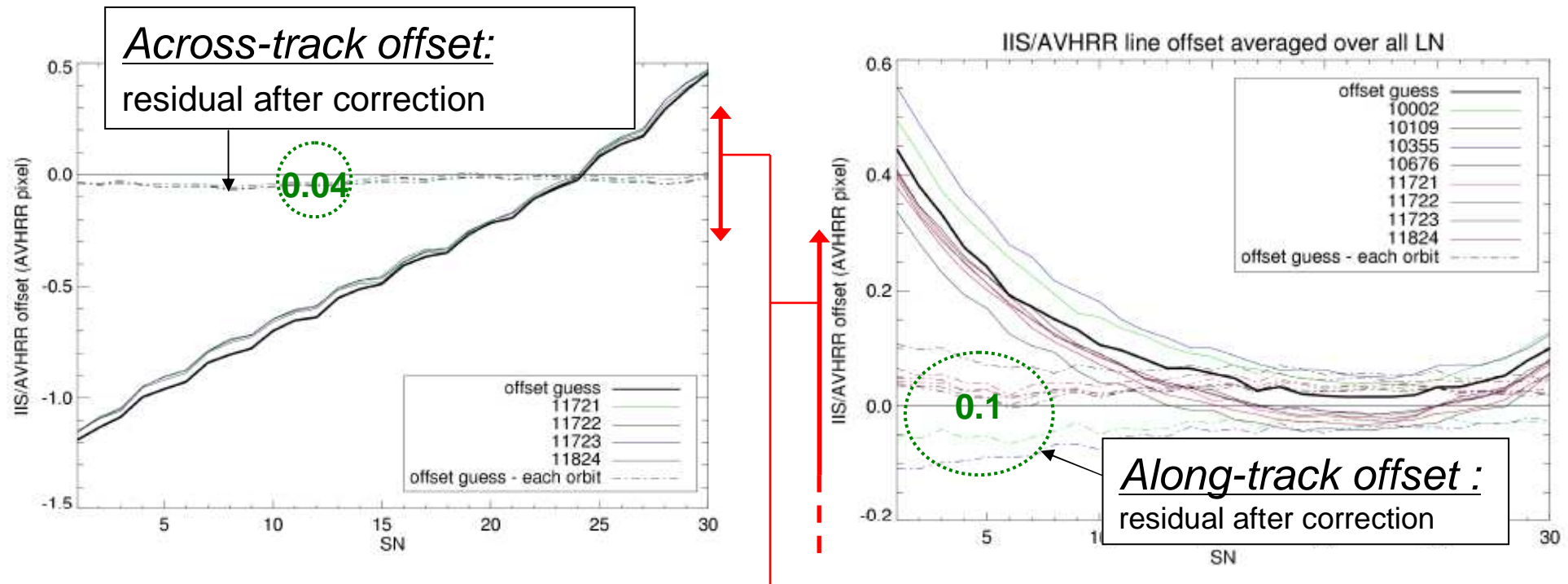


Temporal evolution of IIS average noise



- Stable (0.57K) and widely within the specification (0.8K)

- IIS offset in AVHRR raster : along-track (0.1 AVHRR pixel), across-track (0.04 AVHRR pixel)
- IASI pixel centre localisation accuracy in AVHRR raster ~ 100m



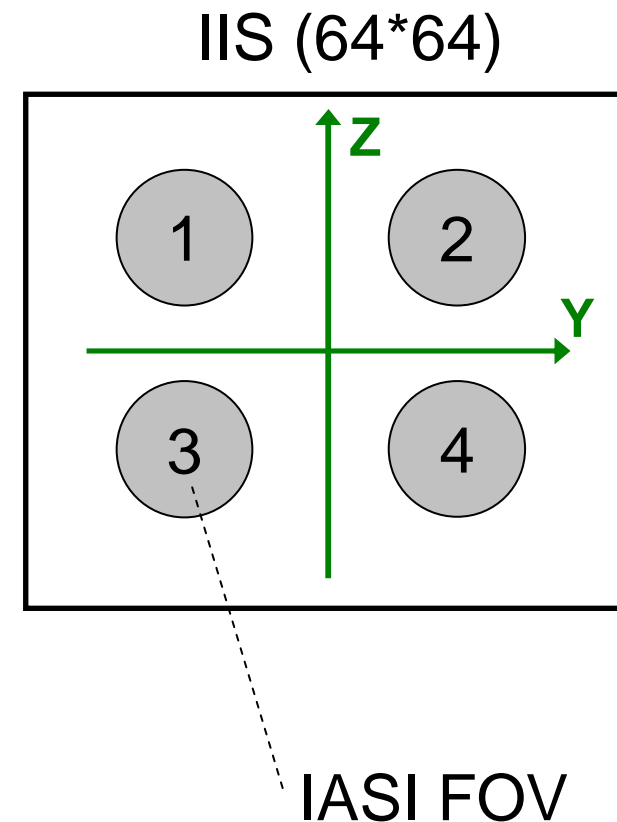
Very good stability since the end of the Cal/Val

→ Health check for scanning mirror mechanism



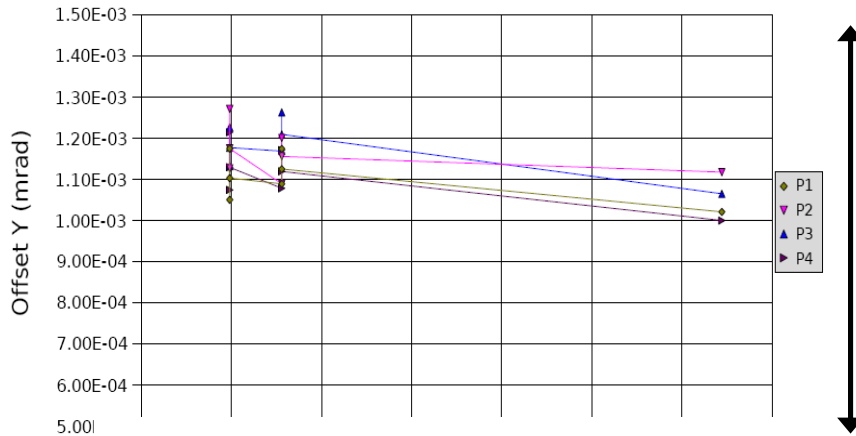
Method:

- Selection of a continuous sequence of scenes with important contrast (coast line, fractional clouds)
- Spectral integration of IASI spectra in IIS spectral band => $U_{ref}(i)$
- Spatial integration of IIS pixels in IASI FOVs for j different positions of IASI FOVs => $U(i,j)$
- Correlation between U_{ref} and U series for all j
- Look for the maximum of correlation => IASI FOVs positions in IIS



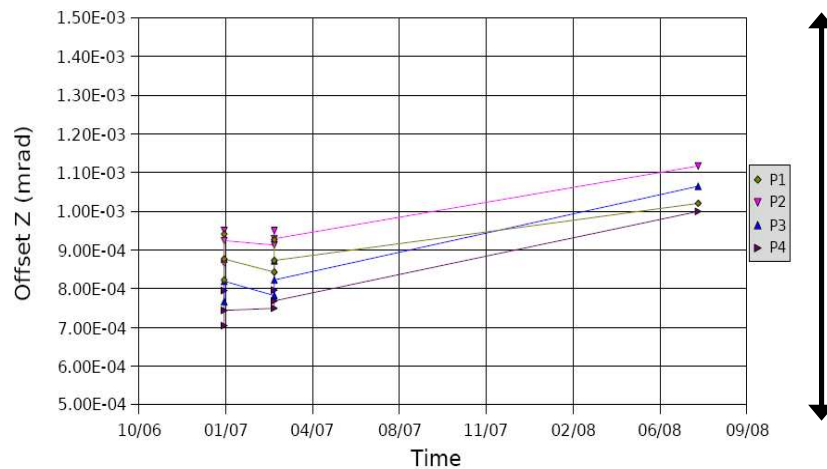


Temporal evolution



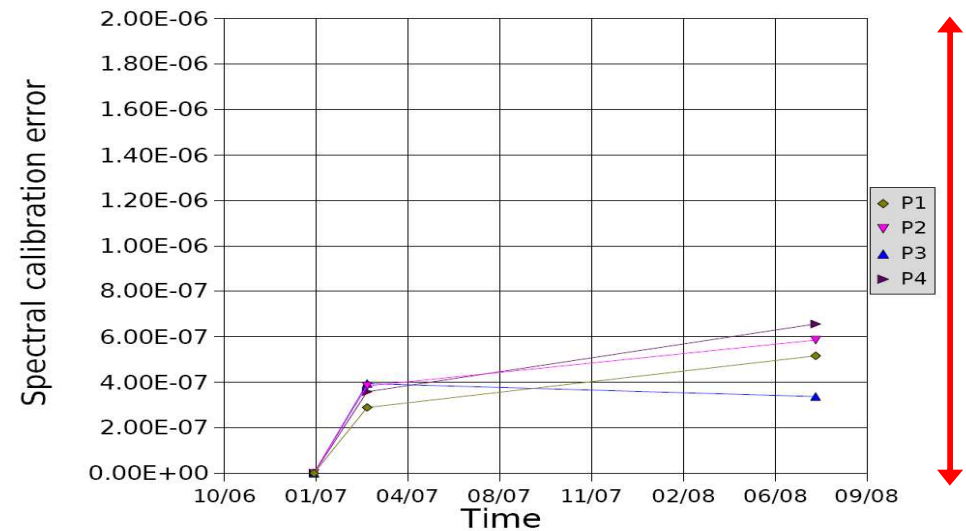
- Specification = +/- 0.8 mrad
- Target = +/- 0.5 mrad

=>



Effect of a IASI FOV offset on spectral calibration quality

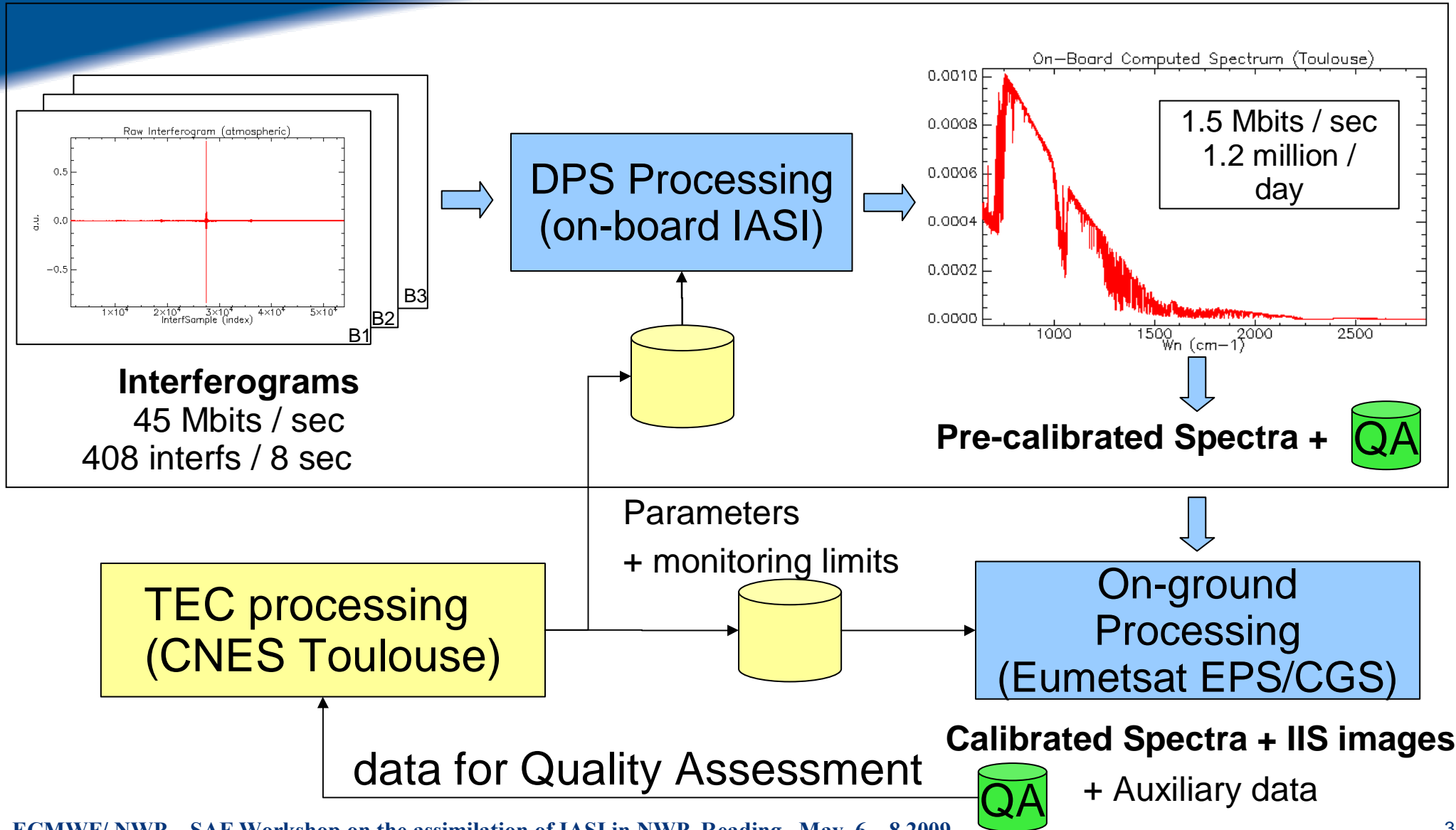
Temporal evolution



Specification: $\Delta\sigma/\sigma < 2 \cdot 10^{-6}$



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■ 8 seconds cycle

- ◆ 30 views (times 4 pixels) for the Normal Op. Mode (27 in Ext.Cal Mode)
- ◆ 2 x 2 calibration views : hot (Black Body), cold (space), 2 scanning directions

■ Main functions

◆ Preprocessing of the interferograms (raw measurements of the interferometer)

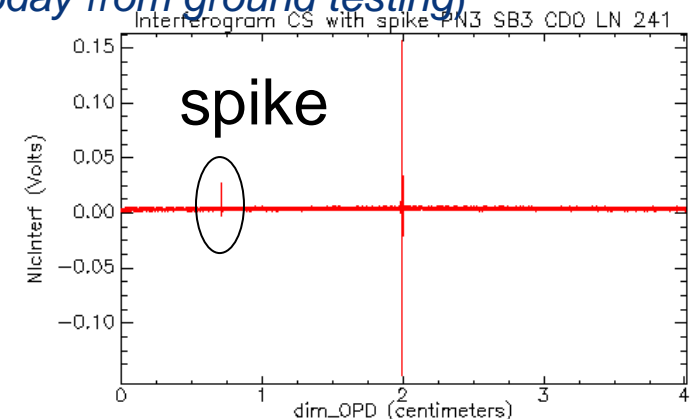
- Integrity checks (spikes detections, etc.) : limits provided by the ground
- Non-Linearity correction : tables provided by the ground (*today from ground testing*)

◆ Computation of calibrated spectra (radiometry)

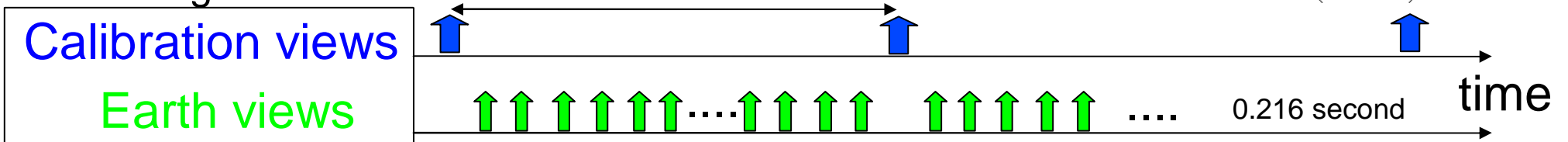
- Internal tables used by calibration updated every 8 sec
 - Reduced spectra Initial values provided by the ground
 - Integrity checks : limits provided by the ground

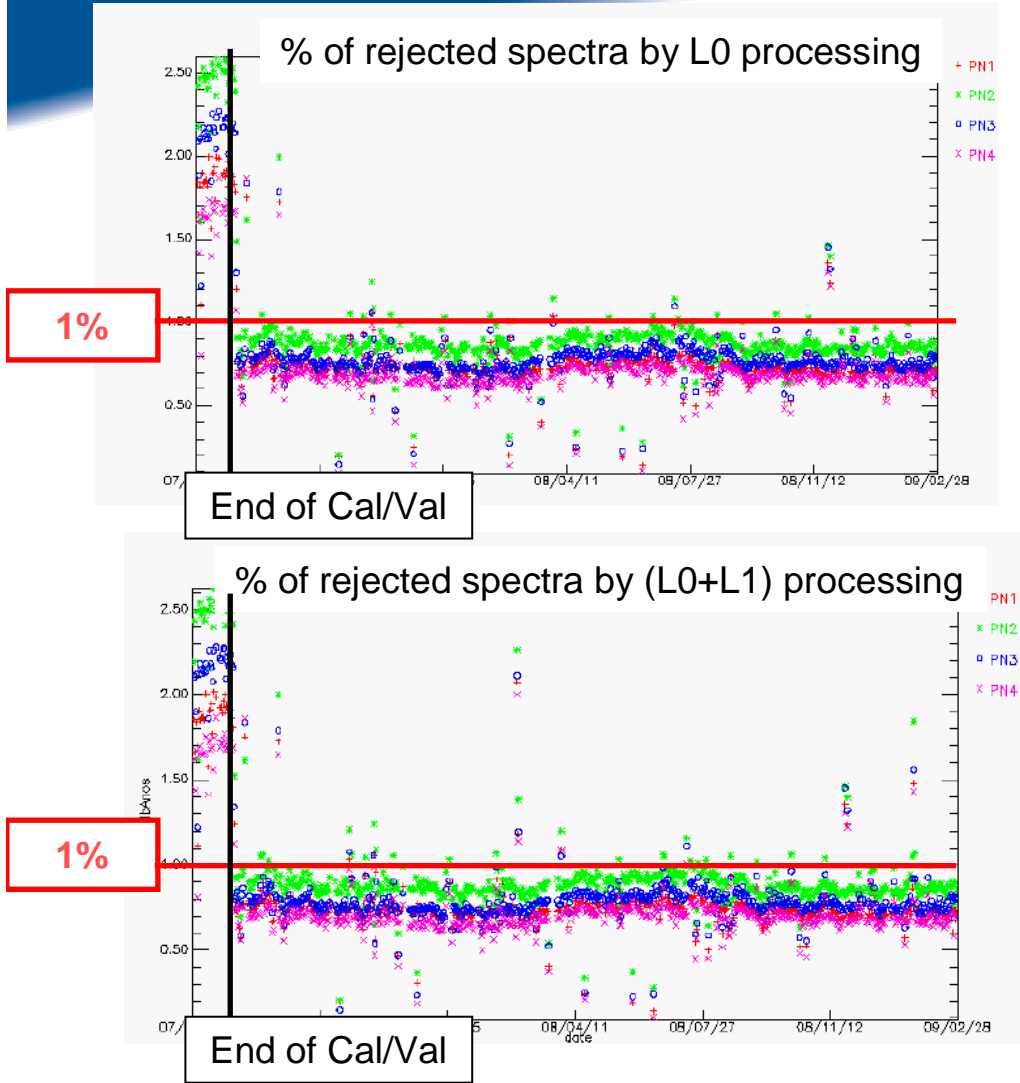
◆ Spectra encoding to reduce data rate

- Programmable Coding Tables provided by the ground



Processing





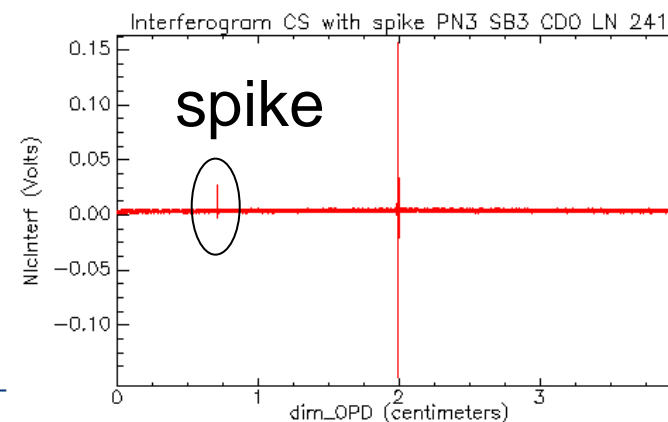
	PN 1	PN 2	PN 3	PN 4
Total % of rejected spectra	0.83	1.01	0.88	0.77
% of rejected spectra by L0 processing (on-board)	0.81	0.99	0.86	0.75

- In NOp, 99% of good quality spectra
- Ground segment is very reliable

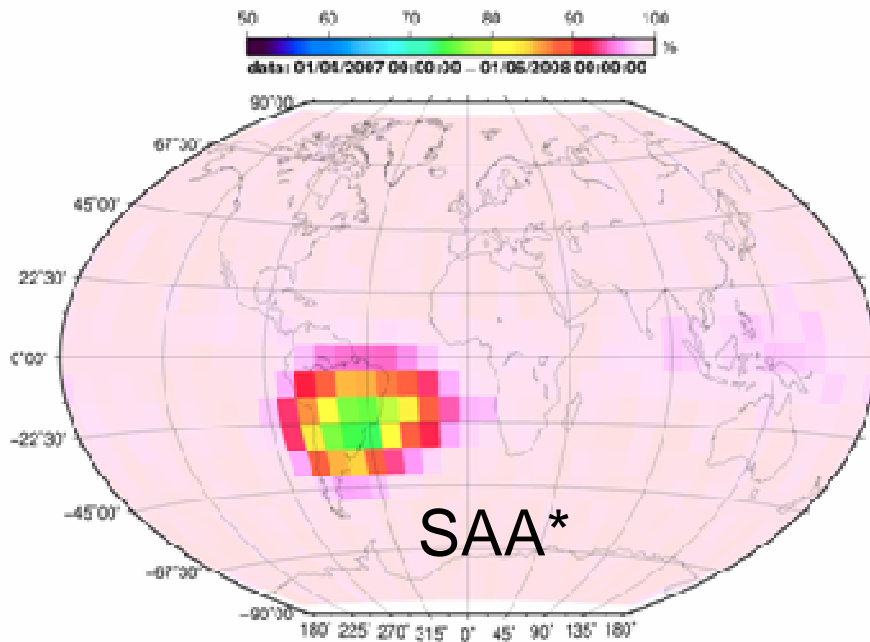
Stable since end of Cal/Val

	Pixel 1	Pixel 2	Pixel 3	Pixel 4
ON-BOARD				
% Spikes (mainly in B3)	0.55*	0.55*	0.55*	0.55*
% NZPD calculation failure	0.15	0.29	0.24	0.15
% radiometric calibration failure	0.02	0.02	0.02	0.02
GROUND				
% Over/Underflows	0.02	0.02	0.02	0.02
TOTAL	0.74	0.88	0.83	0.74
All other parameters	0.09	0.13	0.05	0.03

* Part of « DAY-2 » evolutions



Geographical distribution

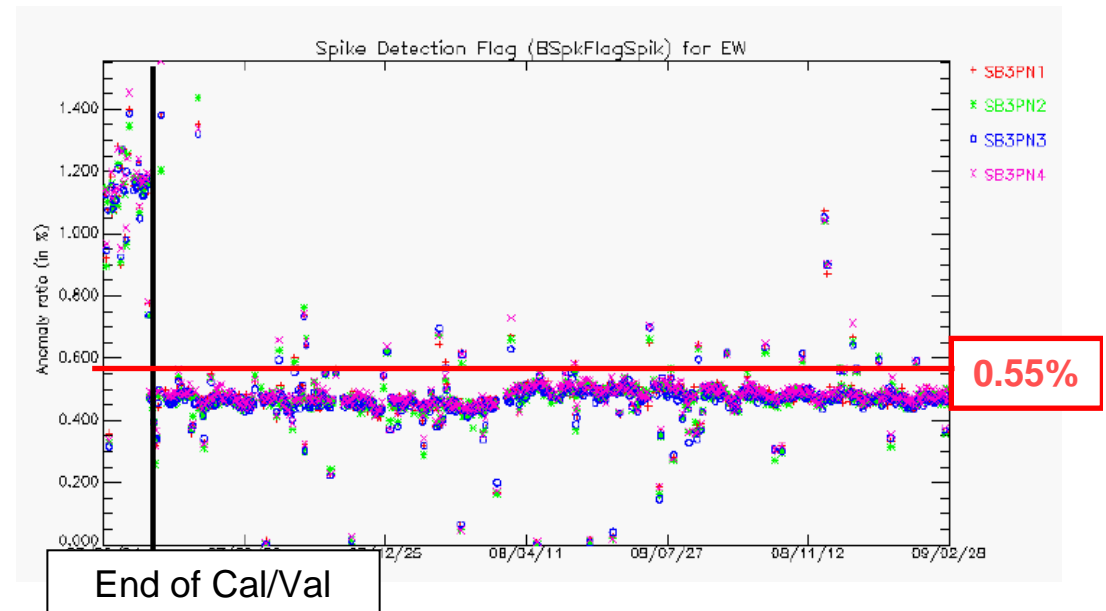


ECMWF 2009 Jun 8 11:27 AM

Courtesy L.Fiedler (EUMETSAT)

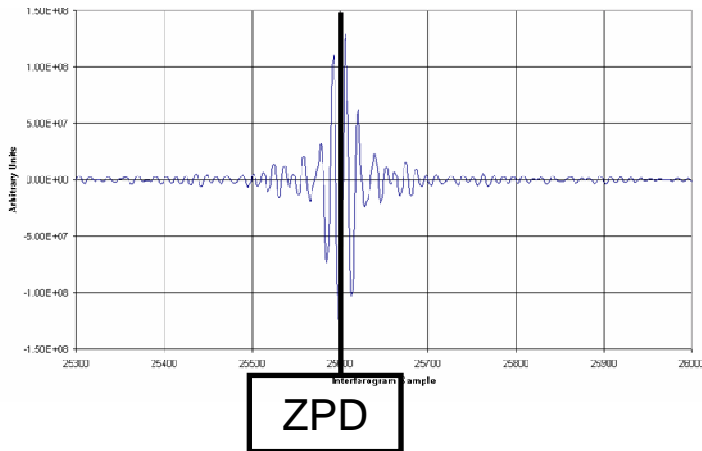
*South Atlantic Anomaly

Temporal evolution

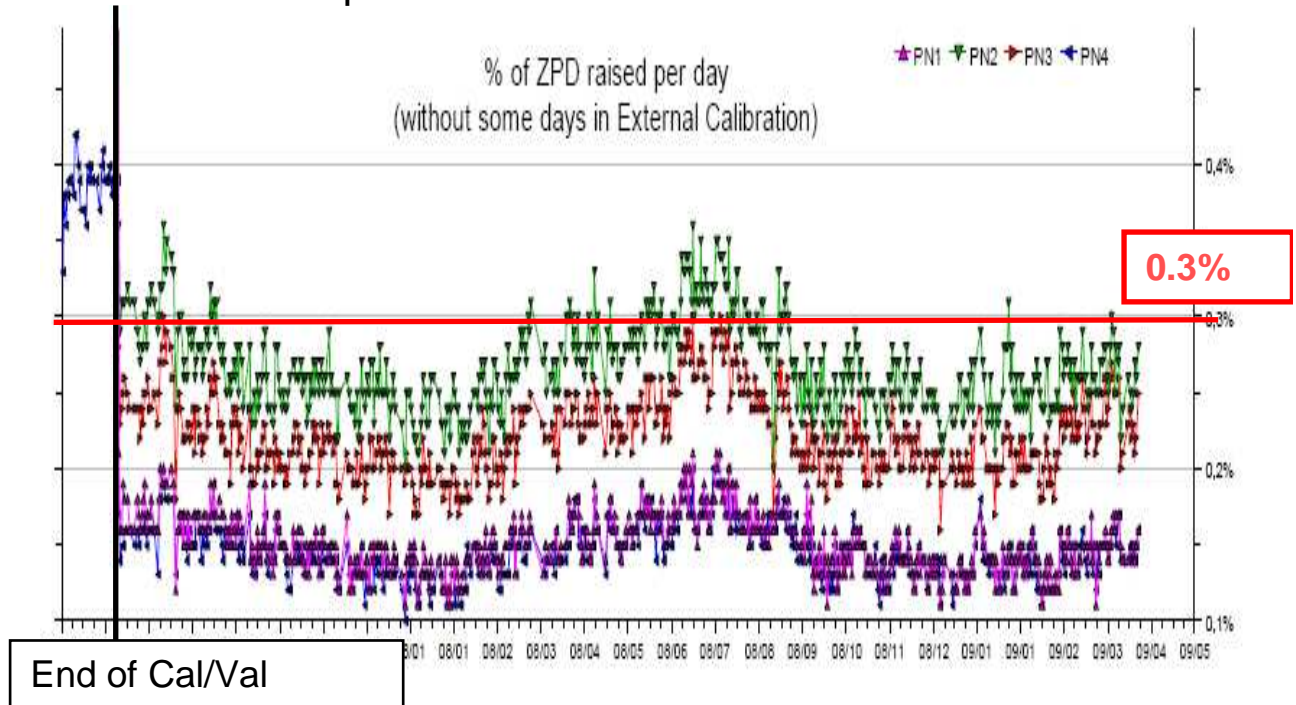


Stable since end of Cal/Val

IASI Interferogram



Temporal evolution of NZPD detection failure

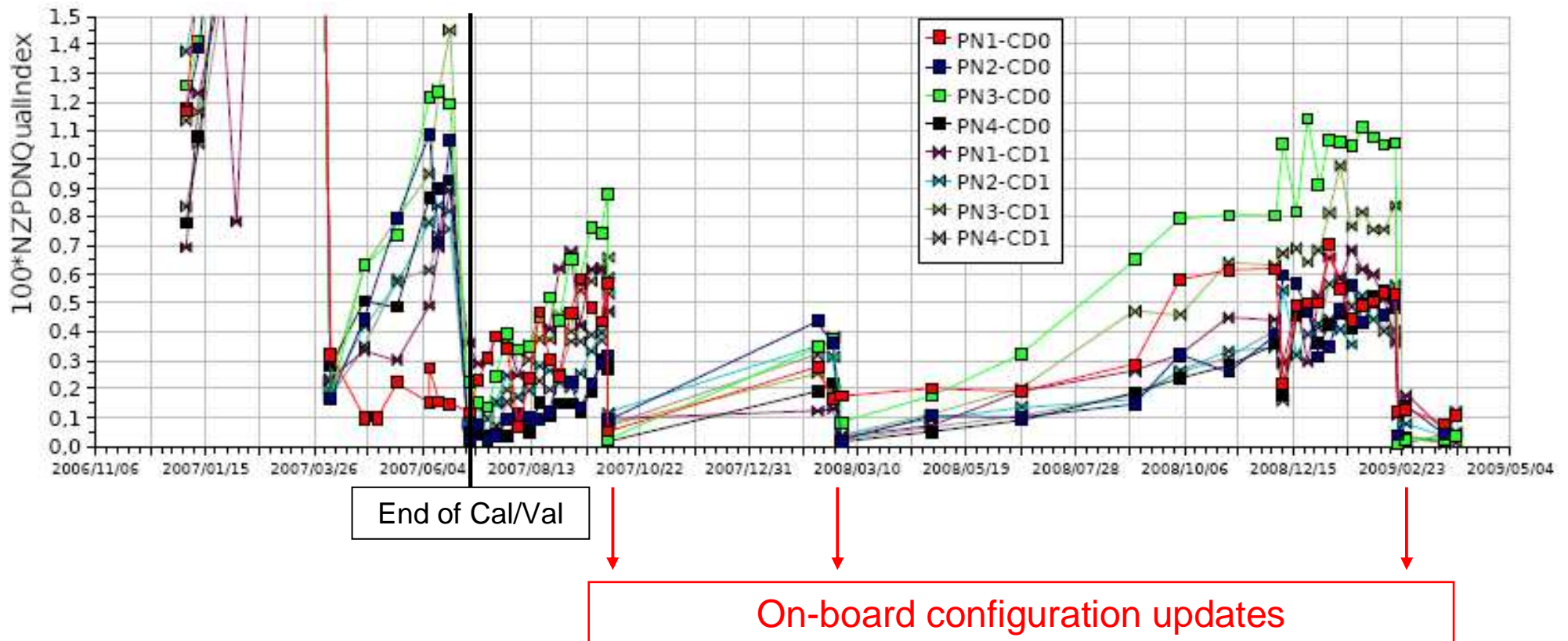


NZPD = sample number at ZPD (calculated by algorithm)

Its knowledge and stability over a calibration period (80s) are necessary for a good radiometric calibration of spectra

Stable since end of Cal/Val
(slight seasonal variation of 0.1% => cold scene)

One of the most important monitoring of on-board processing : it ensures that radiometric calibration of IASI spectra remains good after a re-initialization of on-board configuration (typically after a mode transition NOP \leftrightarrow EC)





- **Instrument design provides good stability**
 - ◆ **In-flight behavior very close to the one measured on-ground**
 - ◆ **Optical bench accurate thermal control (at ambient temperature)**
 - Dimensional stability (hence spectral calibration stability)
 - Radiometric calibration stability
 - But effect of “warm” optical bench on noise in band B3
 - ◆ **Modifications after PFM ground testing against ice contamination**
 - In-flight confirmation of good results obtained on ground
 - Contamination rate continuously decreasing. BUT not very fast (in particular MLI keep desorbing for very long time in orbit)

- **Instrument design provides good testability**
 - ◆ **External Calibration Mode**
 - ◆ **Verification Data Selection (raw interferograms)**



■ Integrated Imager very valuable

- ◆ Easy registration with sounder and AVHRR
- ◆ Very useful for test scenes selections
- ◆ Provide images for calibration views (CS1,CS2, ... moon)
- ◆ Provide images during the ground testing

■ On-board processing working flawlessly

- ◆ All on-board monitoring algorithms proved useful to cope with real data
 - Spikes detection, Reduced Spectra and Radiometric Calibration integrity checks



- ✓ After 30 months in orbit
 - IASI is performing very well
 - **no redundancy used**
 - **all mission requirements are met : both instrument and processing**
 - **the instrument is extremely stable : radiometry, spectral, geometry**
 - **mechanisms (Cube Corner, Scan) show no evolution in orbit**
 - **radiator (passive cooling) show no evolution in orbit**
 - There is still a lot of science to be done with IASI data
 - **Meteorology and Climatology**
 - **Atmospheric chemistry**

- ✓ During the routine phase, IASI Technical Expertise Center (IASI TEC in CNES premises in Toulouse) takes care of :
 - In-depth Performance monitoring, Processing parameters updating

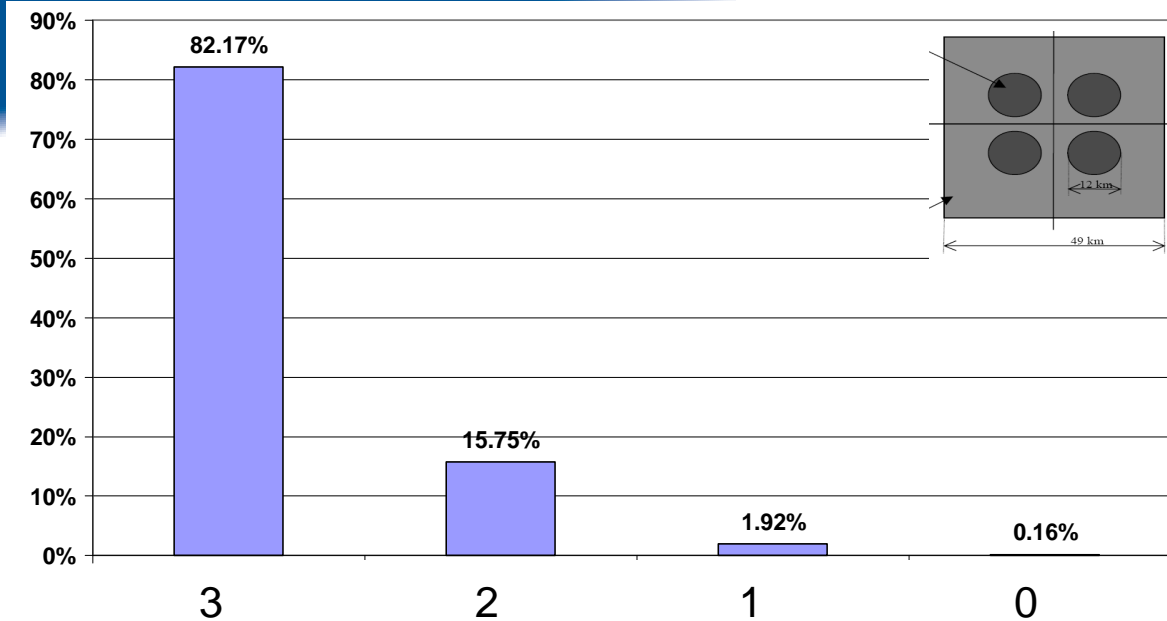
- ✓ In parallel with the operational monitoring performed by the EUMETSAT EPS/CGS teams :
 - Near Real Time PDU analyses, Radiance monitoring

Thank you for your attention!

■ **Visit our web sites :**

- ◆ www.cnes.fr
- ◆ www.smsc.cnes.fr/IASI

BACK-UP



Number of IASI spectra available for each sounding in case of a spike

- 98.5 % of earth views (groups of 4 pixels) are not affected by spikes
- Among the 1.5 % of earth views affected by spikes
 - ◆ Mainly over the South Atlantic Anomaly (SAA) in band B3
 - ◆ 82.2 % have more than 3 spectra available
 - ◆ 97.9 % have more than 2 spectra available
 - ◆ **99.8 % have more than 1 spectrum**

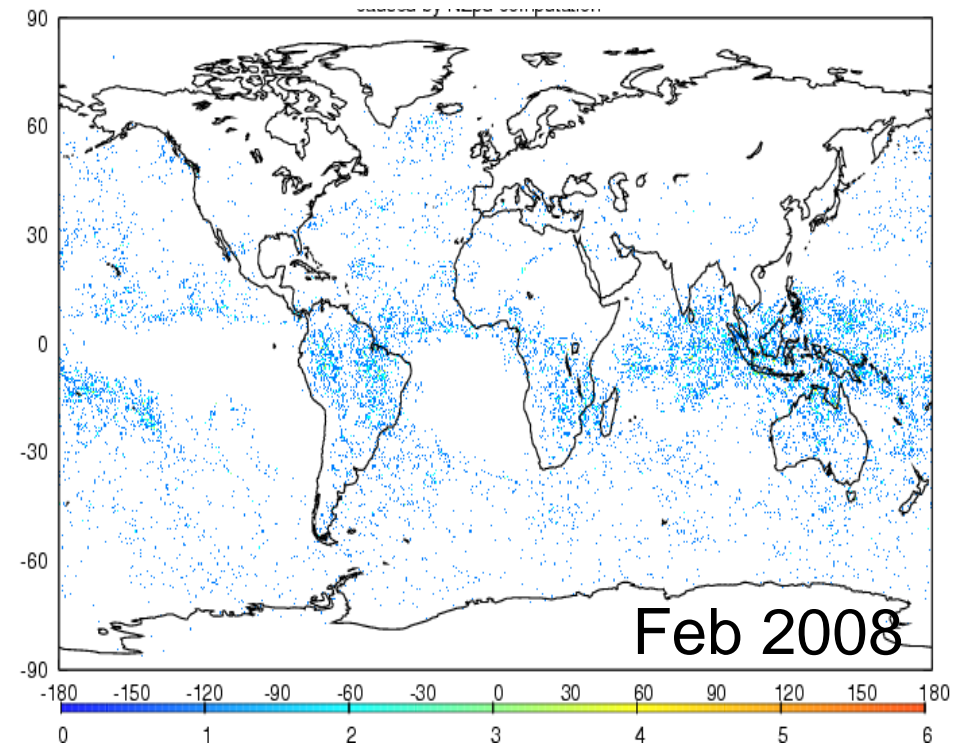
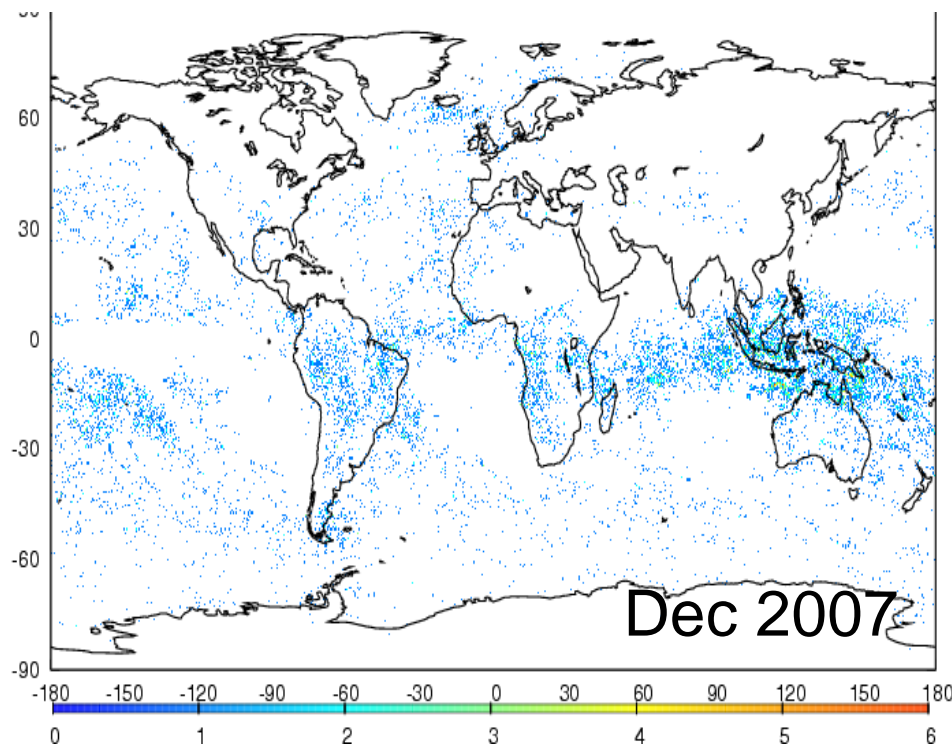
■ Suggestion for short term

- ◆ since the 4 IASI pixels are not assimilated, a dynamic selection of the selected pixel would increase drastically measurements availability.

■ For long term

- ◆ Day 2 evolution of ground processing : make spectra available for bands B1 and B2 when a spike occurs in B3.

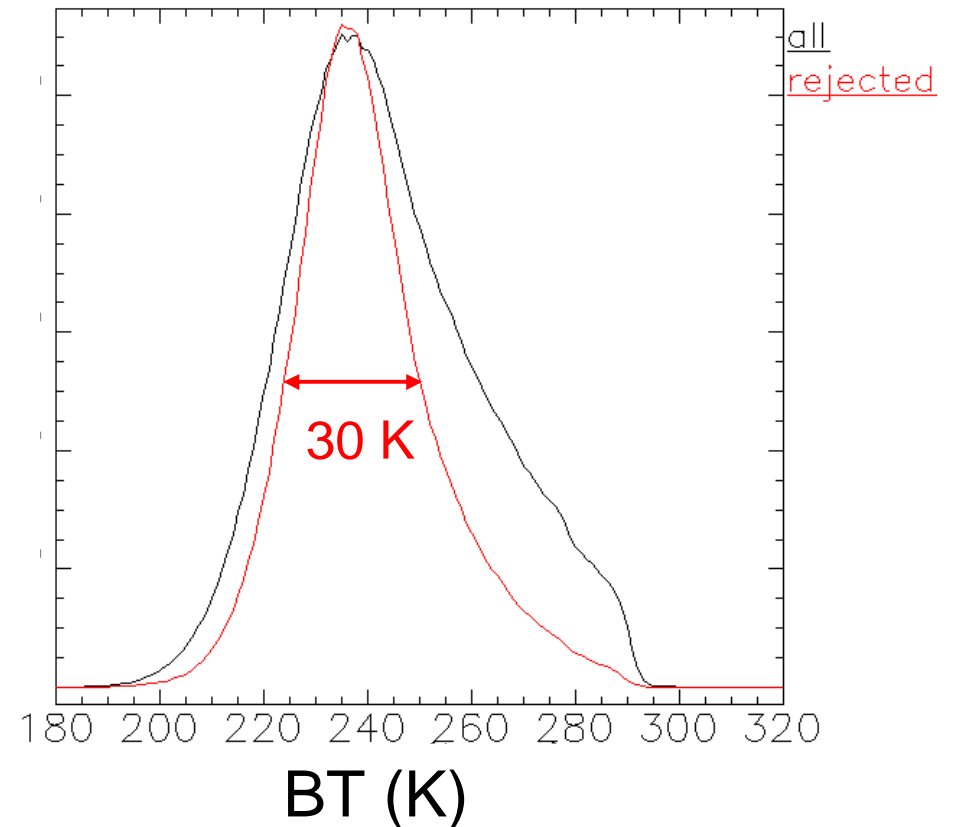
- A small fraction of spectra are not available because of not computed NZPD by on-board processing
 - ◆ Less than 0.3 % for all pixels
 - ◆ Stable since end of Cal/Val (slight seasonal cycle / amplitude ~ 0.1%)
- Geographic repartition
 - ◆ 1 or 2 occurrences max per month per box of 0.5 x 0.5 deg



Brightness Temperatures from the IIS imager

Black curve : Histogram of BT in the vicinity of rejected spectra (about 1/4 of the IIS image)

Red curve : Histogram of BT in the IAS footprint for rejected spectra



■ Conclusions

- ◆ Affected pixels : 0.3 %
- ◆ Histogram of rejected pixels : FWHM = 30 K
- ◆ Close shape of the 2 histograms → no significant impact on climatology